



Space Science & Technology  
Rutherford Appleton Laboratory

Imperial College  
London

# Modelling of the Highly Miniaturised Radiation Monitor

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On behalf of the HMRM collaboration

(STFC Rutherford Appleton Laboratory & Imperial College London)

Space Radiation and Plasma Environment Monitoring Workshop

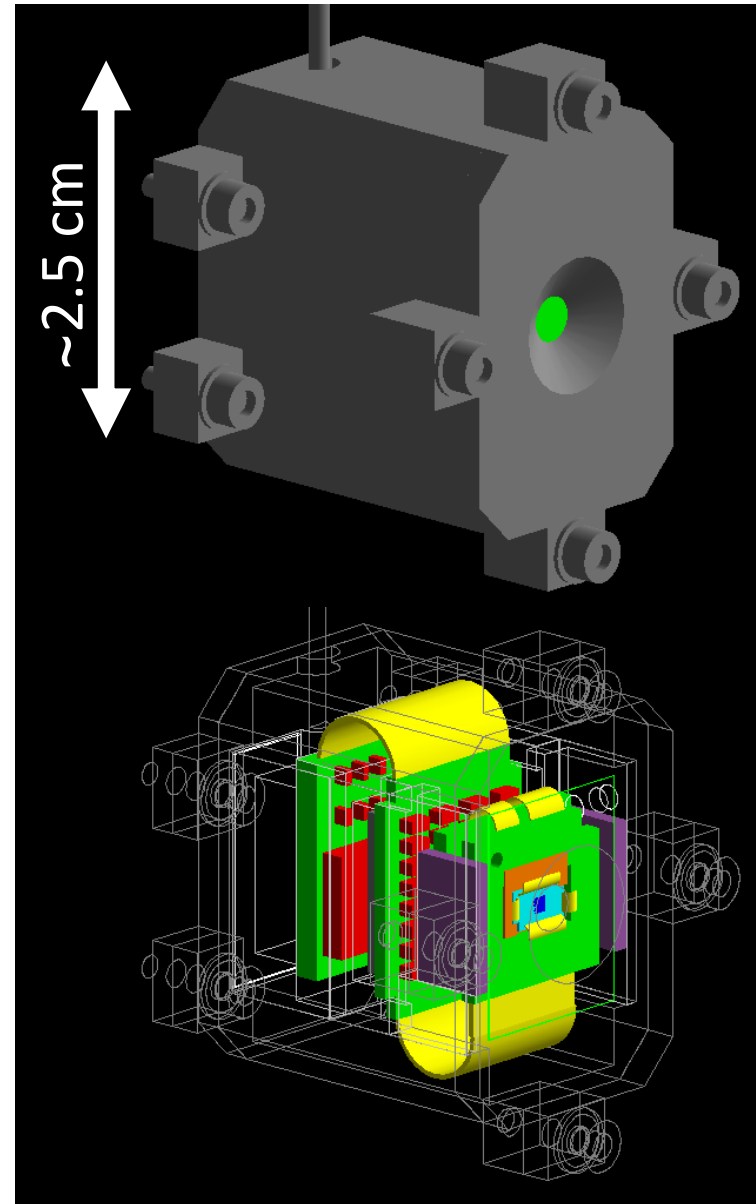
ESA/ESTEC, Noordwijk, The Netherlands

10<sup>th</sup> May 2012

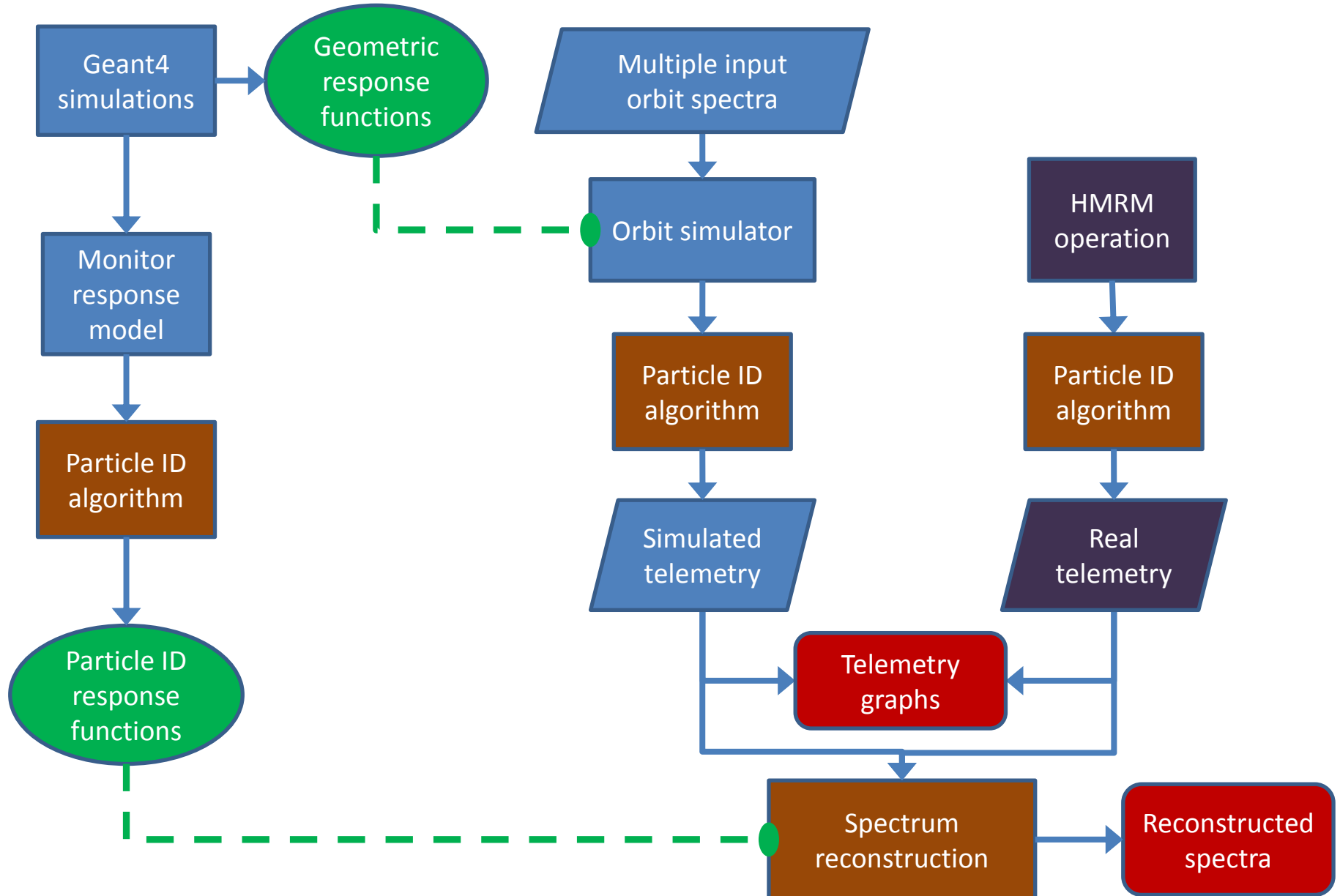
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# HMRM operation summary

- Miniature monitor for energetic charged particles in a range of earth orbits
- Sensors read out at 10 kHz: up to 100  $\mu\text{s}$  exposure time, divisible down to 3.125  $\mu\text{s}$
- FPGA executes particle ID algorithm on each read out
- Particle data output:
  - Count rate, dose rates
  - Identified particle rates } In monitor, in real-time
- Spectrum reconstruction (in development) ground segment, offline



# Modelling overview



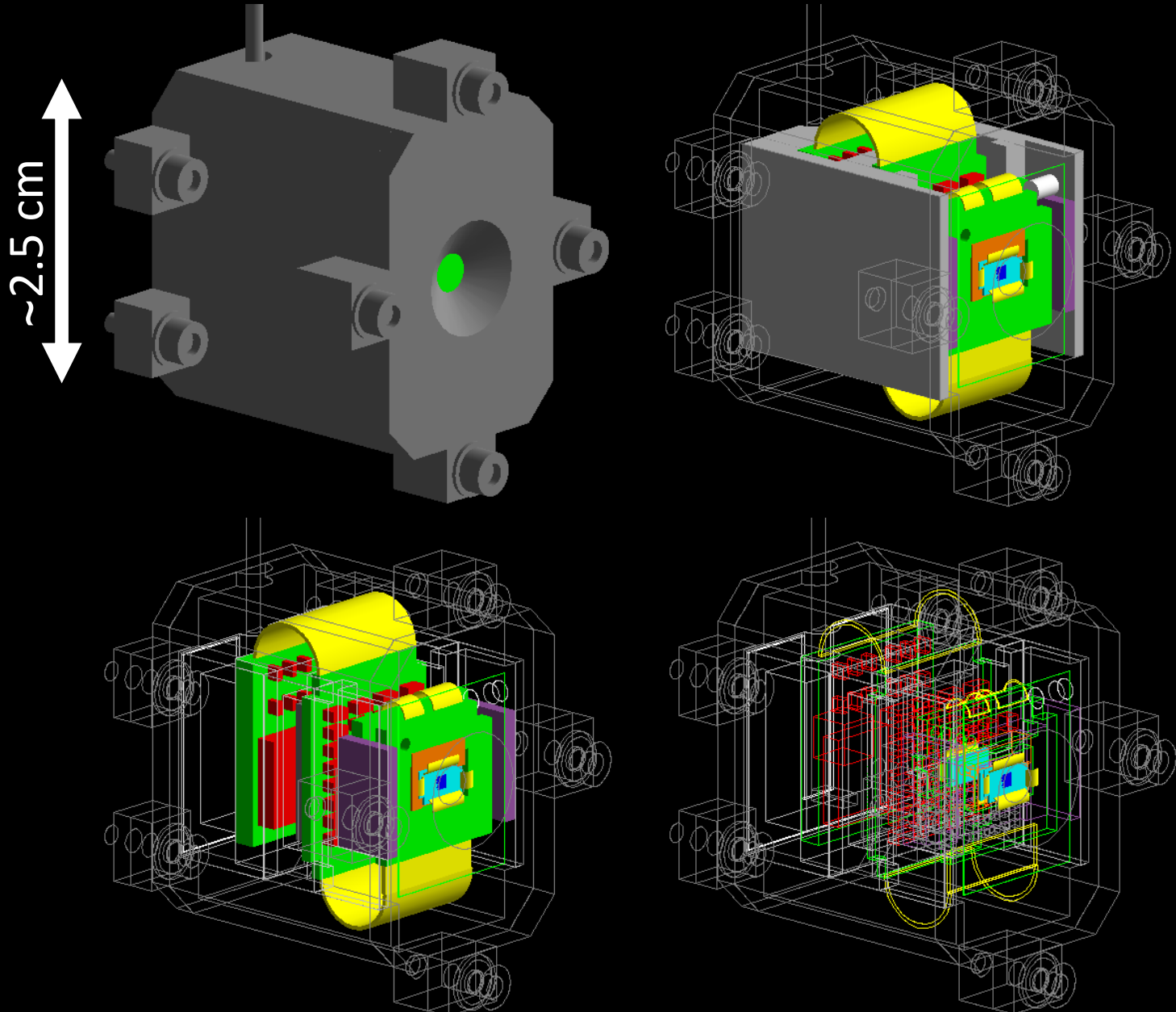
# Particle identification

1. Pixel signals summed to find total charge for each sensor
2. Event allocated to one of 32 channels via a lookup table. Some channels are of high purity
3. Table imposes coincidence, and anticoincidence constraints
4. No-hit event counter informs integration time “shuttering”
5. Table chosen through:
  - Theory (similar to  $\Delta E, E$  detectors<sup>[1] [2] [3]</sup>)
  - Experiment
  - **Simulation**

Channel	Sensor 1		Sensor 2		Sensor 3		Sensor 4		ID
	L1	U1	L2	U2	L3	U3	L4	U4	
1	0.0	0.2	0.0	0.2	0.0	0.2	0.0	0.2	No Hit
2	361.0	624.2	23.4	208.8	0.0	0.2	0.0	0.2	Proton 2.5 - 4.0 MeV
3	23.4	69.8	23.4	69.8	23.4	40.4	0.0	0.2	Proton 20 - 60 MeV
4	4.2	6.6	4.2	6.6	4.1	11.9	1.5	13.5	Proton 180 - 500 MeV
5	4.5	13.5	1.5	2.6	0.0	0.2	0.0	0.2	Electron > 0.1 MeV
6	1.5	2.6	1.5	2.6	2.6	4.5	0.0	0.2	Electron > 0.5 MeV
7	1.5	7.8	1.5	2.6	0.9	1.5	0.0	0.2	Electron > 1.0 MeV
8	13.5	23.4	0.0	0.2	0.0	0.2	0.0	0.2	Mixed
9	120.7	208.8	120.8	208.8	0.0	0.2	0.0	0.2	Mixed

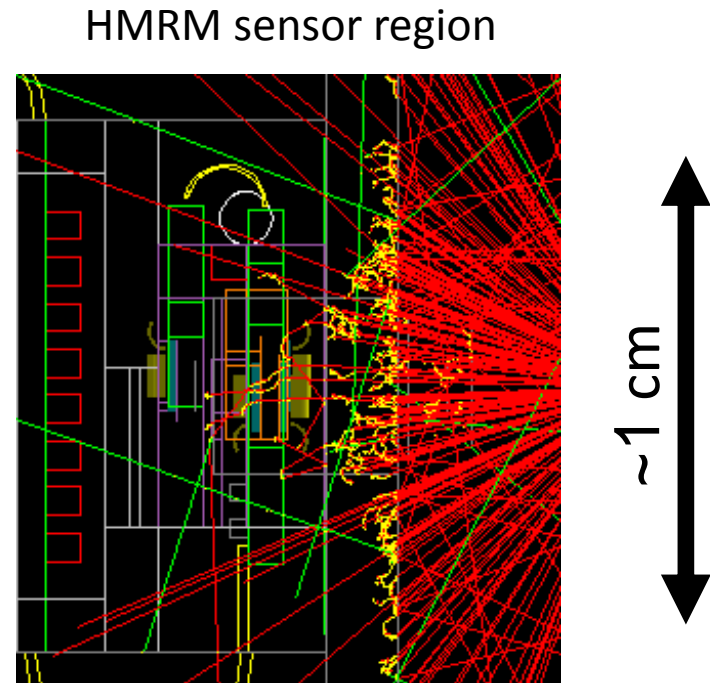
Part of an example table (scaled units)

# Geant4 simulation geometry



# Geant4 simulation

- Highly accurate Monte Carlo, extra validation with data for
  - Energy losses in thin films<sup>[4]</sup> <sup>[5]</sup>
  - Backscattering angular distributions<sup>[6]</sup>
- More than  $10^9$  primaries simulated
- Electrons: 0.04 – 6 MeV
- Protons: 1 – 500 MeV
- Obtain sensor energy deposits as a function of incident particle species, energy, angle



1 MeV electrons demonstrating front shielding and acceptance

# Response functions (RFs)

- Avoid the need for further Geant4 simulations
- Doubly differential in particle energy and solid angle, per unit incident particle fluence

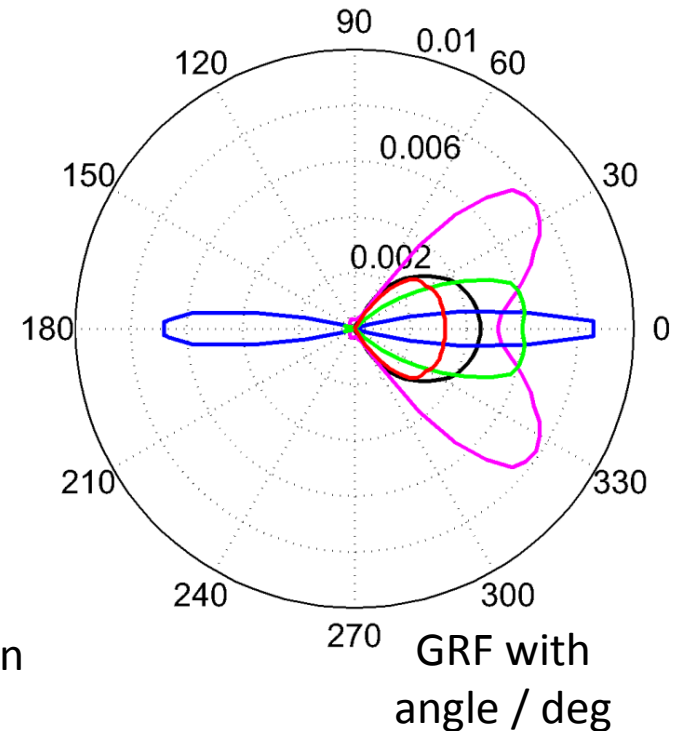
Particle	Energy range / MeV	Sensor coincidence
Proton	1 – 10	1
	10 – 60	1 + 2 + 3
	60 – 500	1 + 2 + 3 + 4
Electron	0.04 – 0.3	1
	1 - 6	1 + 2

- Geometric RF
  - Probability of a hit to each or any sensor
- Energy deposit RF
  - Energy deposit p.d.f for each sensor, per hit

Used in orbit simulator

- Particle ID RF
  - Probability of obtaining a count in each channel
  - Assuming a single hit (no pile up)

Used in spectrum reconstruction



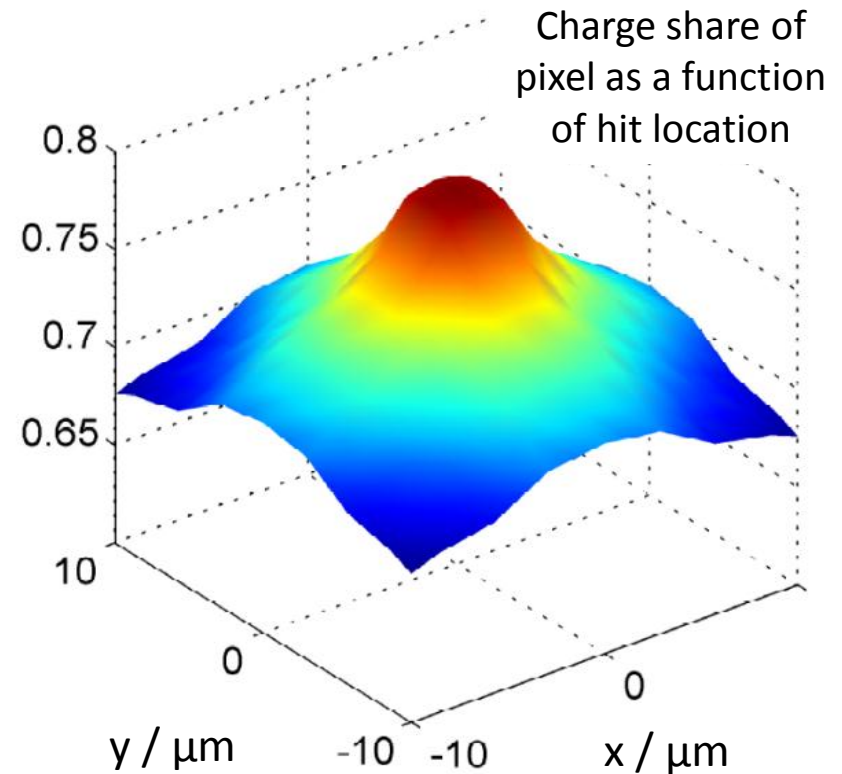
# Monitor response model

Convert Geant4 energy deposits into realistic signal by introducing:

1. **Pixellation effects:** lateral charge diffusion equation
2. **Noise:** simple Gaussian model
3. **Analogue to digital conversion:** 7 programmable comparator levels

Diffusion and noise model parameters fitted via experiment

Proposed 2D charge diffusion model result using arbitrary parameters



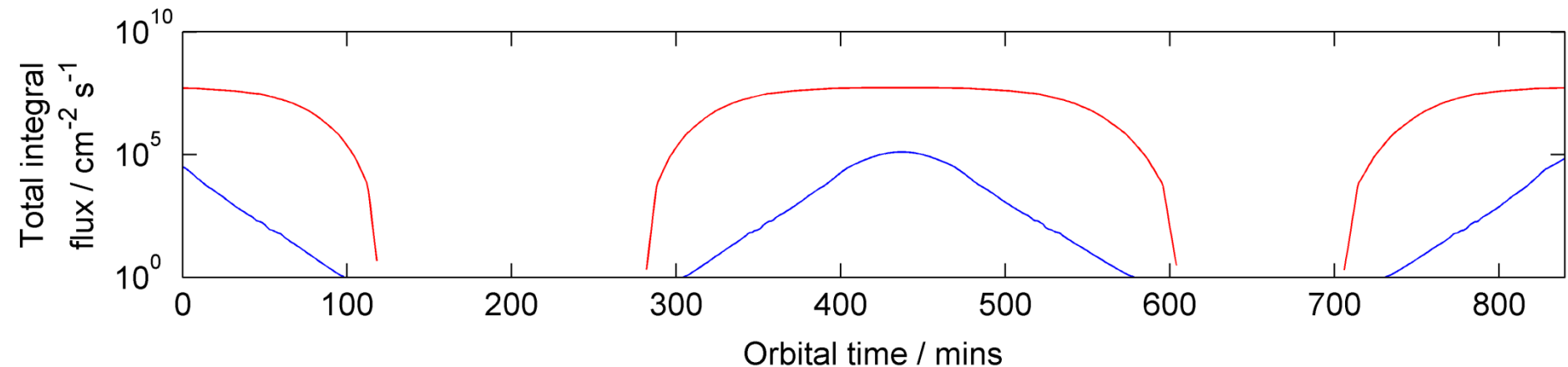


# Orbit simulator

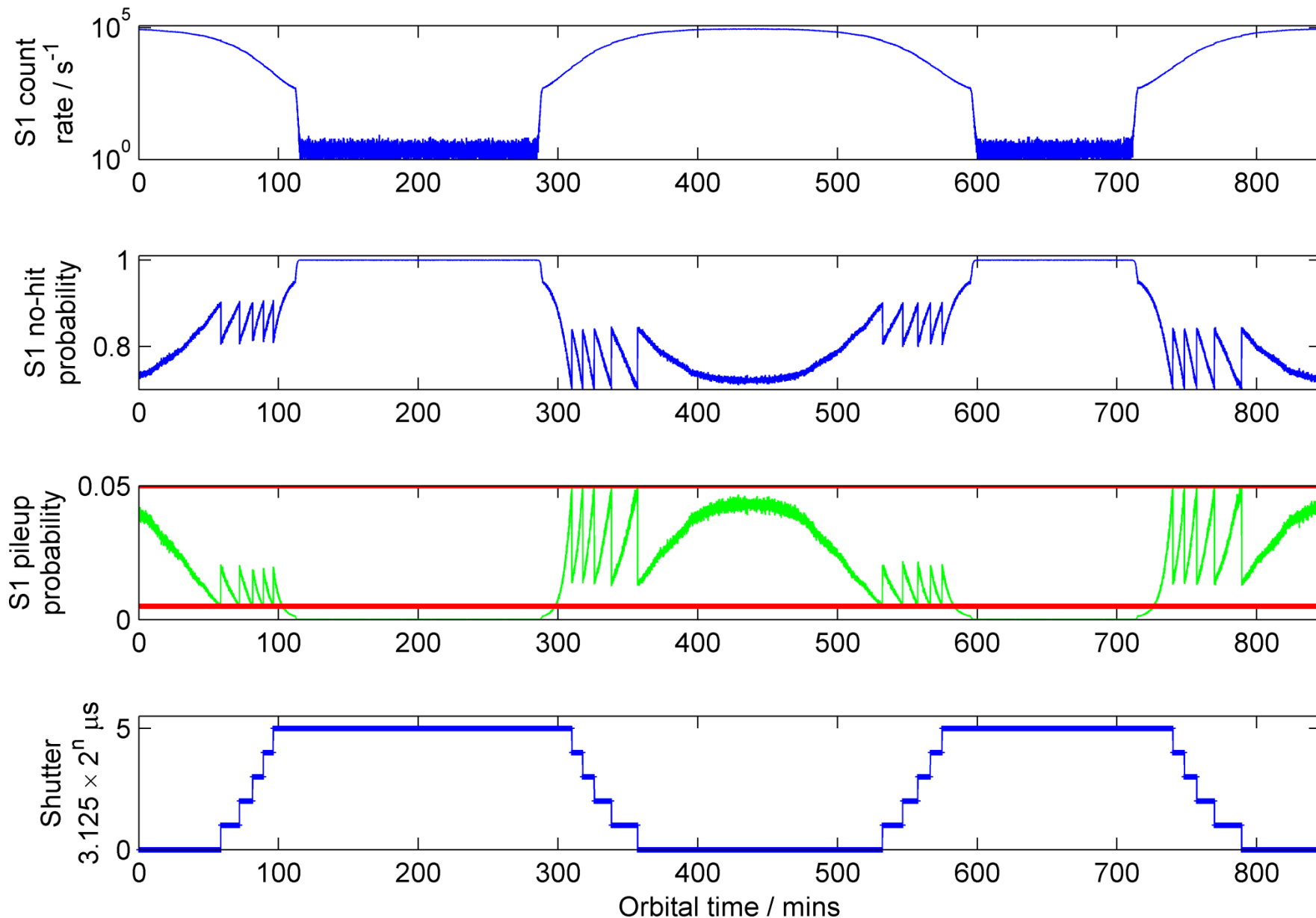
- Introduce:
  - Time-dependence (transient spectra, relative motion, pile up effects, exposure “shuttering”)
  - Multiple simultaneous species and spectral components (trapped particles, GCR, SEP etc) compatible with SPENVIS<sup>[7]</sup> output files
- At each point in time:
  1. Fold response functions with instantaneous incident spectra
  2. Monte Carlo sample the resulting distributions
  3. Apply monitor response model
  4. Execute HMRRM algorithm

# Orbit simulator: example results

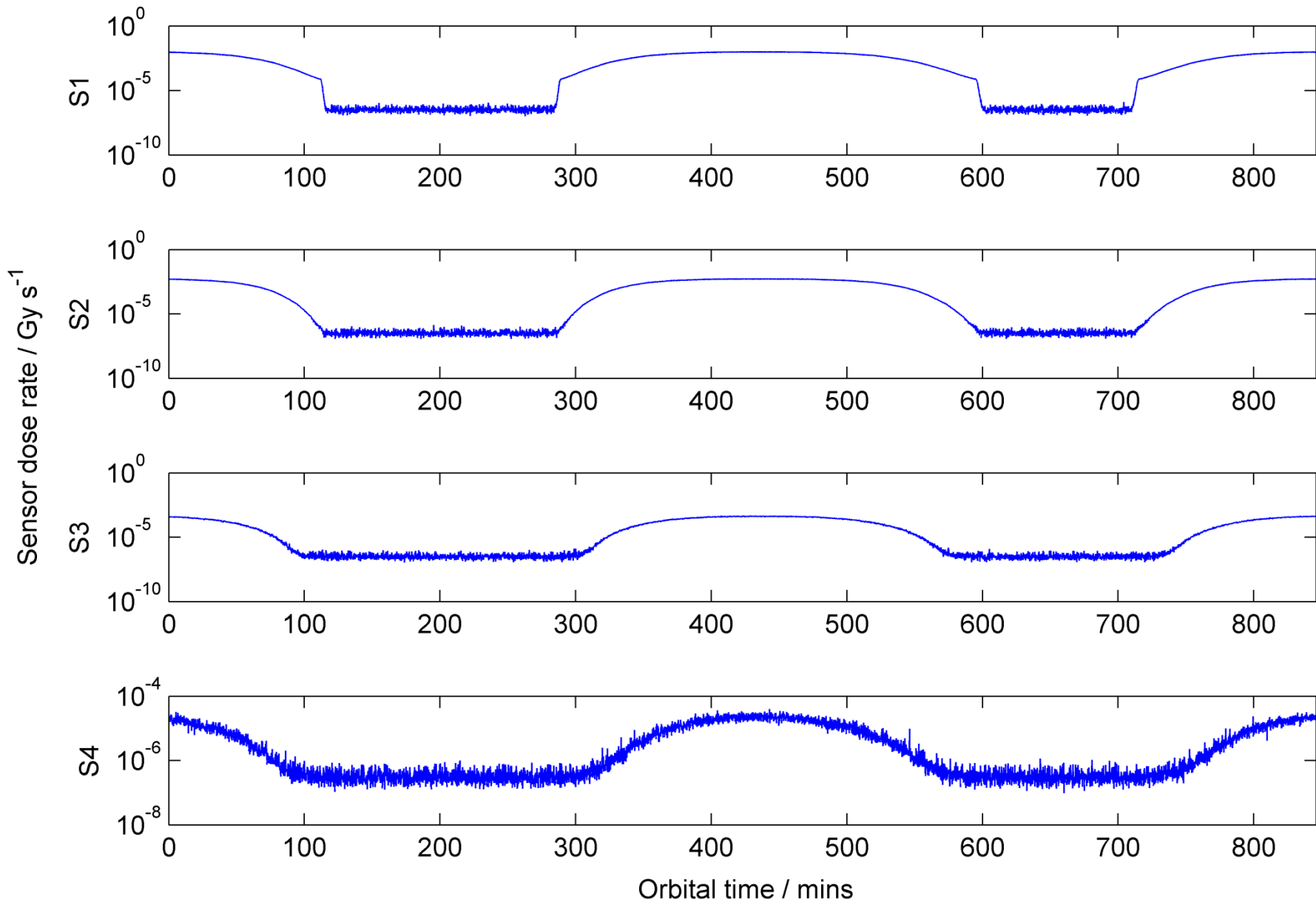
- Example orbit: 23,222 km MEO at 56° inc. (840 min period)
- Assume isotropic fluence, using omnidirectional AP-8 and AE-8 model spectra
- Total integral flux for **protons** and **electrons** shown below (simulator uses full energy spectra).



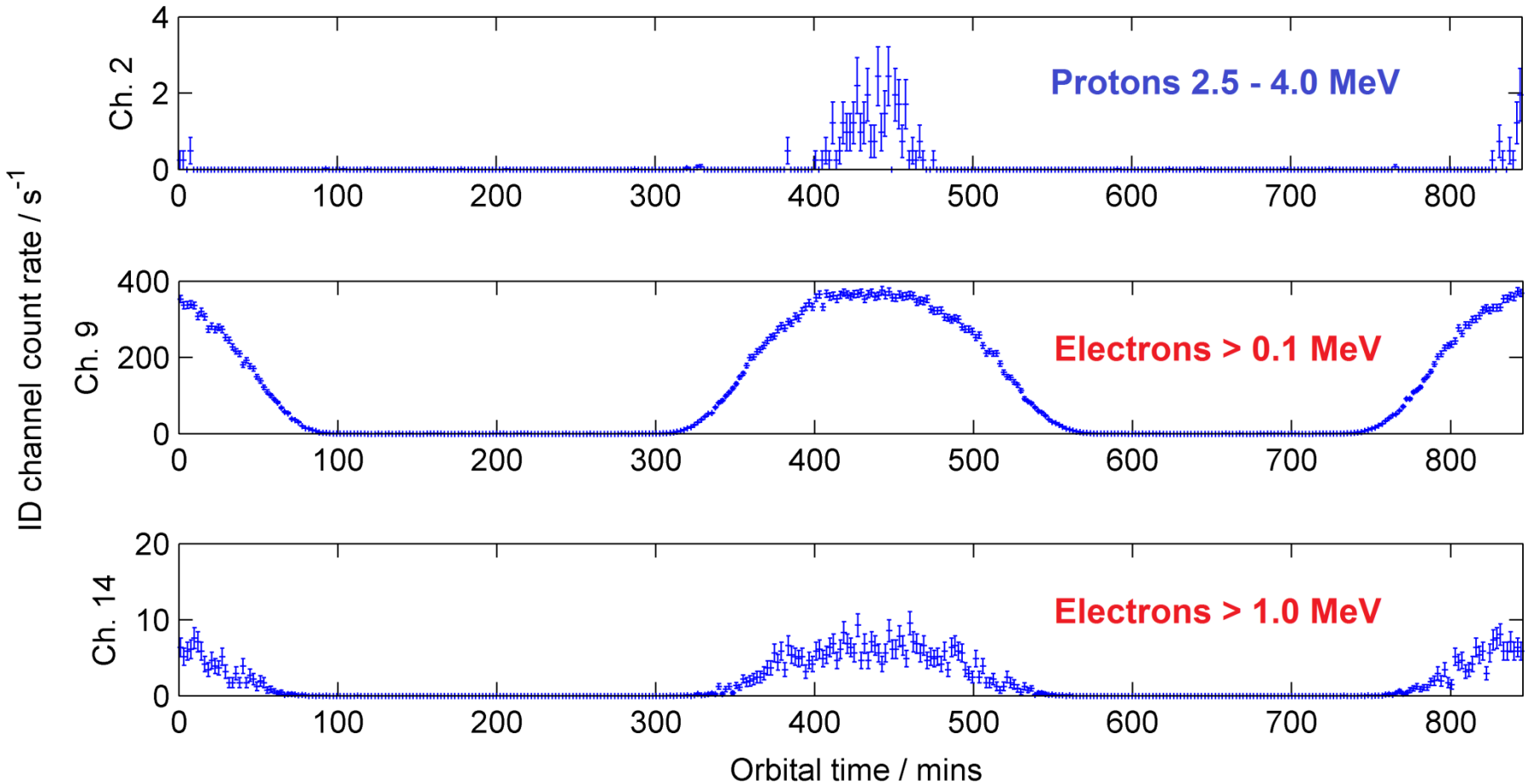
# Orbit simulator: example results



# Orbit simulator: example results

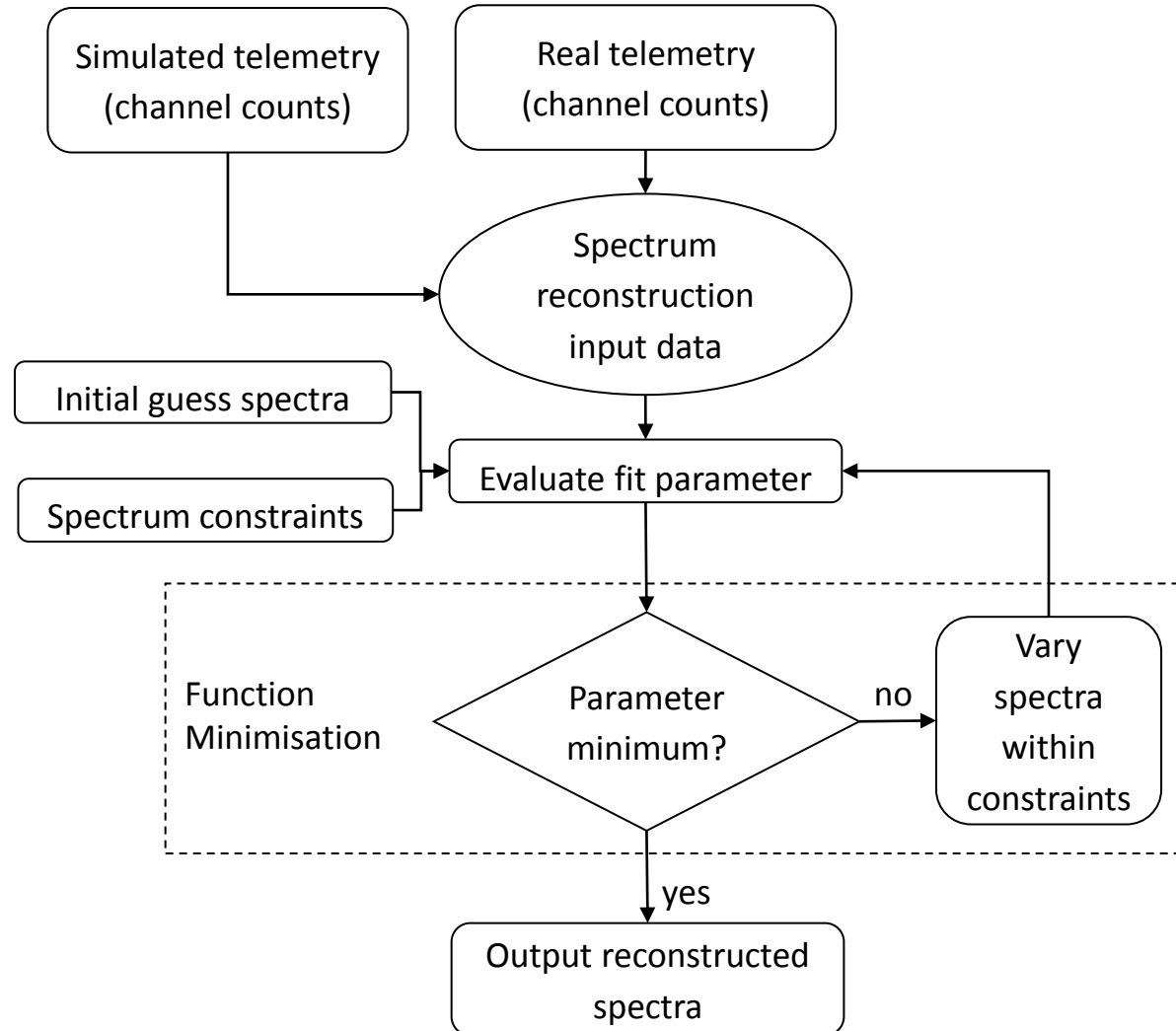


# Orbit simulator: example results



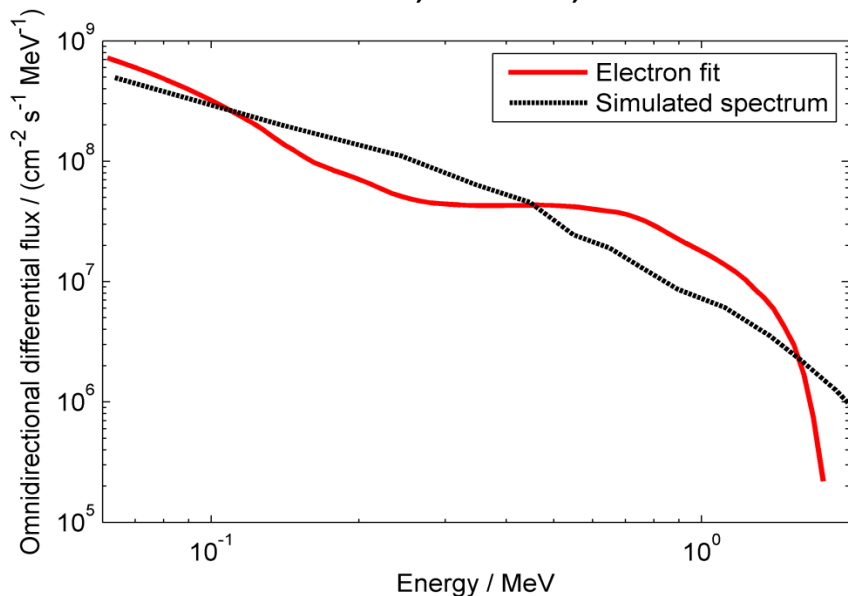
# Spectrum reconstruction development

- Simplistic matrix method rejected
  - $n = Ms$  ;  $s = M^{-1}n$
  - Intolerant of fluctuations
  - Requires a large number of channels, resulting in small counts
- Proposed method: Iterative fitting with Likelihood or Least Squares parameter
- Electron and proton spectra are simultaneously fitted (combined hypothesis)

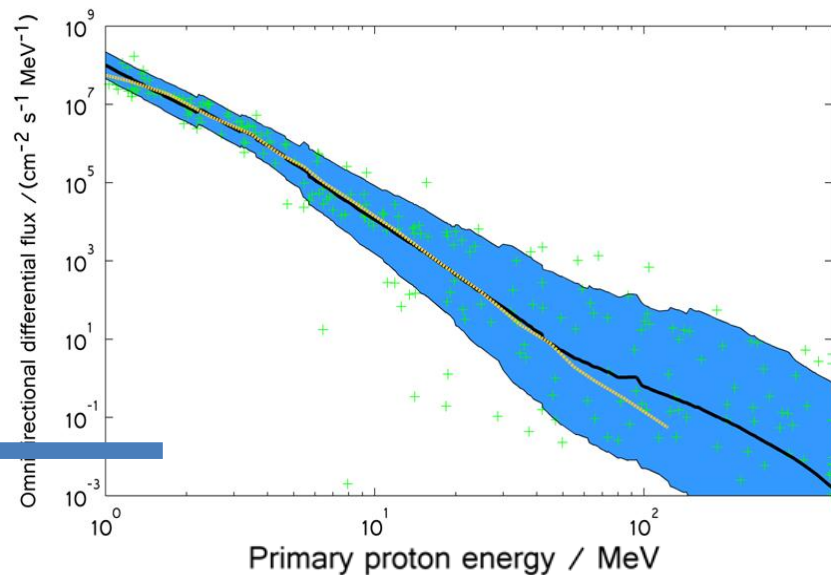
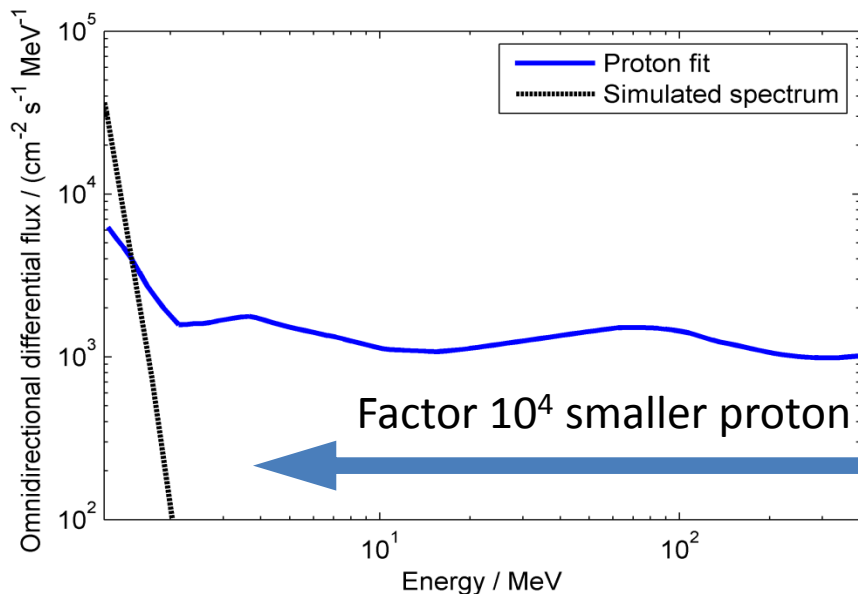
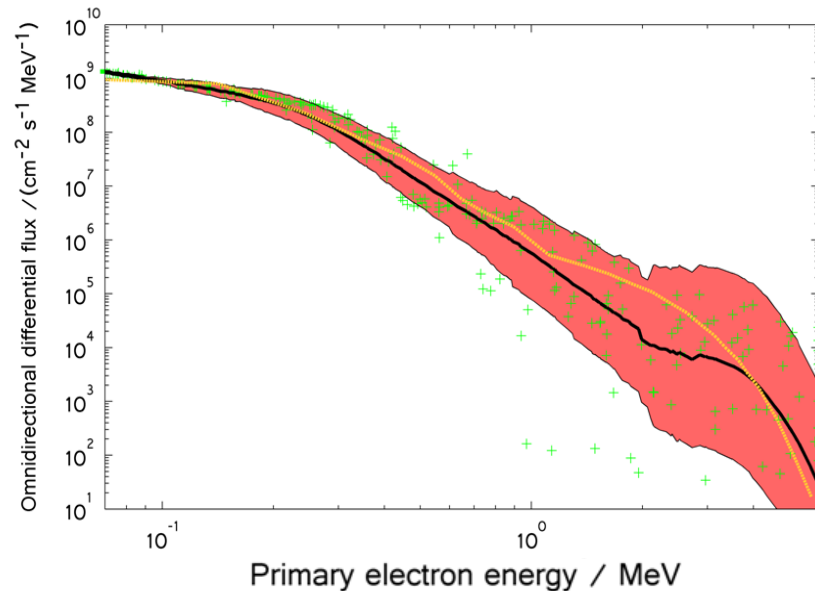


# Spectrum reconstruction: initial tests

60 minutes in 23,222 km, 56° inc. MEO



10 seconds in 10,000 km, 0° inc. MEO



# HMRM collaboration

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  - D. Griffin
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  - N. Guerrini
  - O. Poyntz-Wright
  - S. Woodward
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  - E. Mitchell
- ESA
  - A. Menicucci

## References

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3. F. S. Goulding and B. G. Harvey, Ann. Rev. Nucl. Sci. **25** (1975) 167 and references therein.
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6. G. R. Massoumi et al., Phys. Rev. B **47** (1993) 11007-11018
7. SPENVIS: <http://www.spervis.oma.be/>