

# *A Bethe-Heitler 5D Polarized Photon-to- $e^+e^-$ Pair Conversion Event Generator*

D. Bernard

LLR, Ecole Polytechnique and CNRS/IN2P3, France

12th Geant4 Space Users' Workshop G4SUW  
10-12 April 2017  
Guildford, Surrey, England

<http://llr.in2p3.fr/~dbernard/polar/harpo-t-p.html>

# *Talk Layout*

- The context: HARPO: high-performance  $\gamma$ -ray astronomy and linear polarimetry with conversion to  $e^+e^-$  pairs with a gas TPC
  - telescope performance studies NIM A 701 (2013) 225
  - polarimeter performance studies NIM A 729 (2013) 765
  - cosmic-rays TPC tracker characterization NIM A 718 (2013) 395
  - polarized  $\gamma$ -ray beam data-taking campaign PoS (ICRC2015) 1016
  - high dilution factor polarimetry on beam demonstrated SPIE (2016) 99052R
  - Recent Summary SciNeGHE2016 arXiv:1702.08429
- Event generator:
  - Past achievements: a VEGAS based generator
    - Exact, 5D, polarized, Event generator NIM A 729 (2013) 765
    - Event generator comparison Astropart. Phys. 88 (2017) 60
  - Present activities: towards VEGAS free generation
  - Perspectives for the future: change it to a private G4EmModel
    - Donation as a public physics model if agreeable by the G4 Coll.

# HARPO: angular resolution and sensitivity

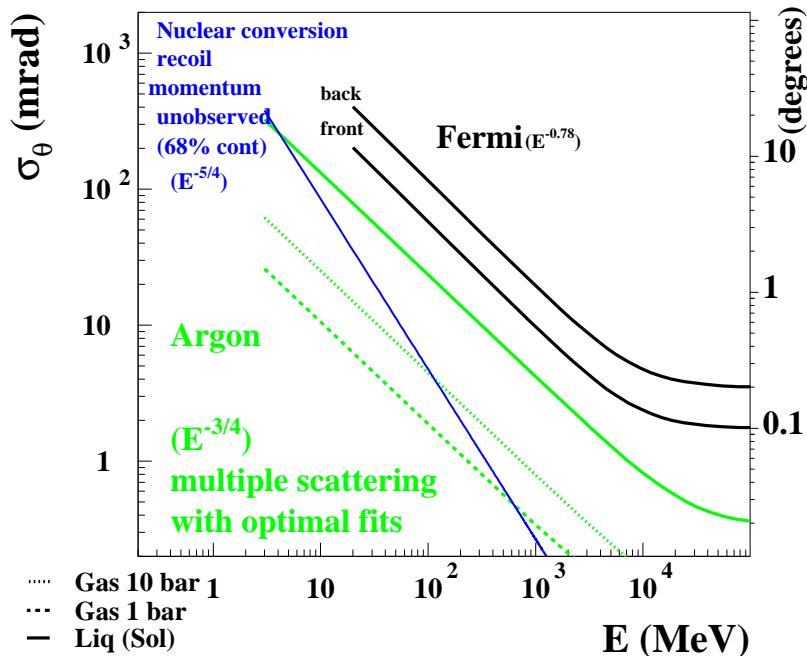
- For  $\gamma \rightarrow e^+e^-$ , the sensitivity wall is mainly an angular resolution wall.

single photon angular resolution

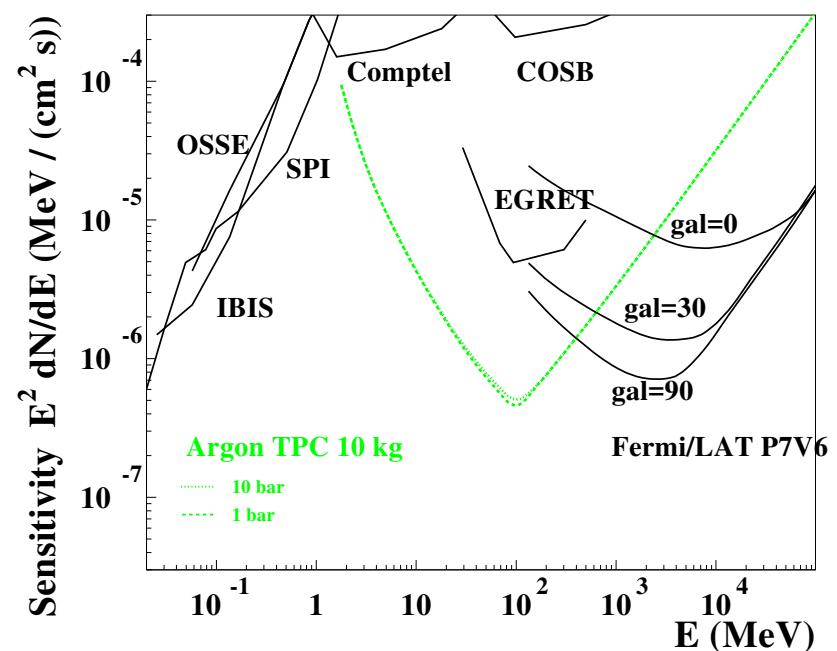
$$0.27 \oplus 0.27 = 0.38^\circ @ 100 \text{ MeV}$$

point-like source sensitivity

(à la Fermi, 3 year,  $5\sigma$ ,  $> 10\gamma ..$ )

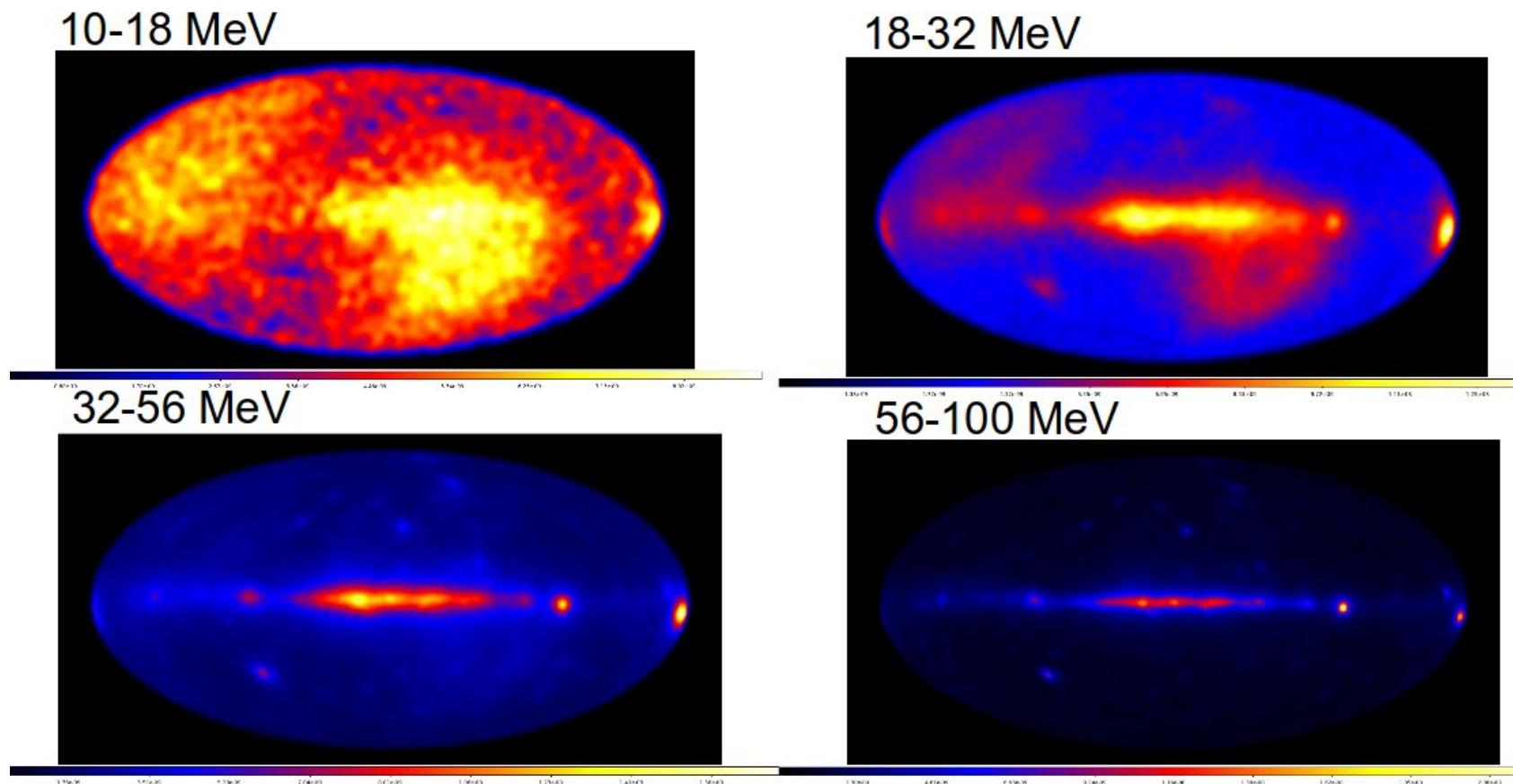


Nucl. Instrum. Meth. A 701 (2013) 225



Nucl. Instrum. Meth. A 701 (2013) 225

# *Angular resolution and sensitivity*



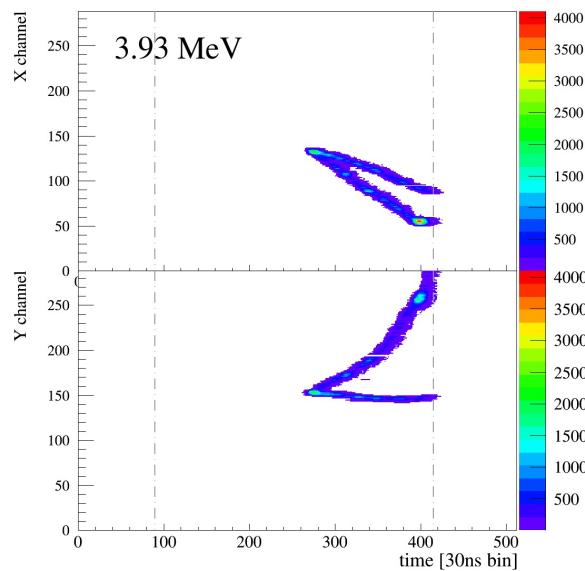
"Fermi-LAT below 100 MeV (Pass8 data)", Julie McEnery,

"e-ASTROGAM workshop: the extreme Universe", Padova Feb-March 2017

# HARPO: Linear polarimetry

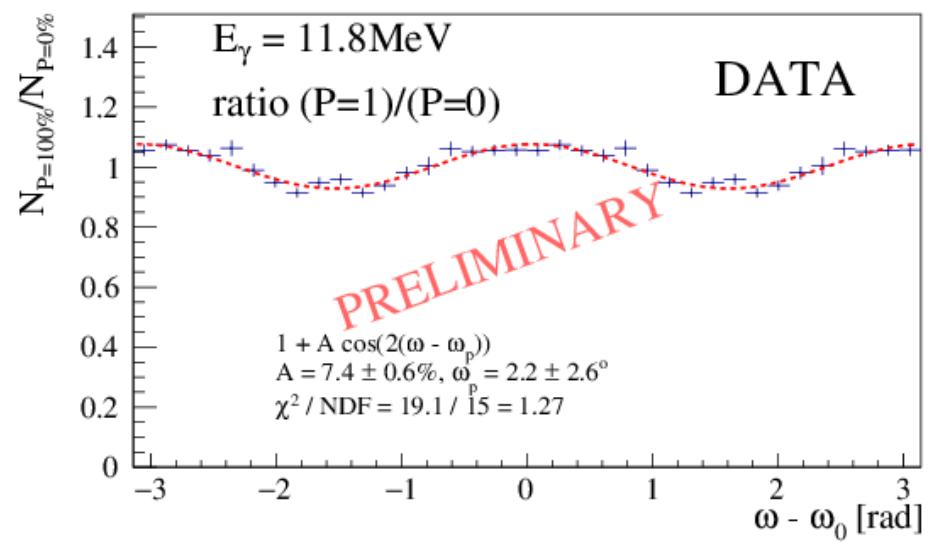
- Polarimetry major diagnostic at low energies (radio - optics - X)
- Missing  $E > 1 \text{ MeV}$
- Would enable decipher leptonic/hadronic Blazar models, point to emission region in pulsars, tag the transition energy for magnetars, extend Lorentz invariance violation (LIV) searches sensitivity ...
- Needs large statistics → needs low energies

$(x, t)$  and  $(y, t)$  signal maps  
4 MeV  $\gamma$  in 2.1 bar Ar-iC4H10 95-5%



SPIE (2016) 99052R

azimuthal angle distribution  
11.8 MeV  $\gamma$  in 2.1 bar Ar-iC4H10 95-5%



SciNeGHE (2016) arXiv:1702.08429

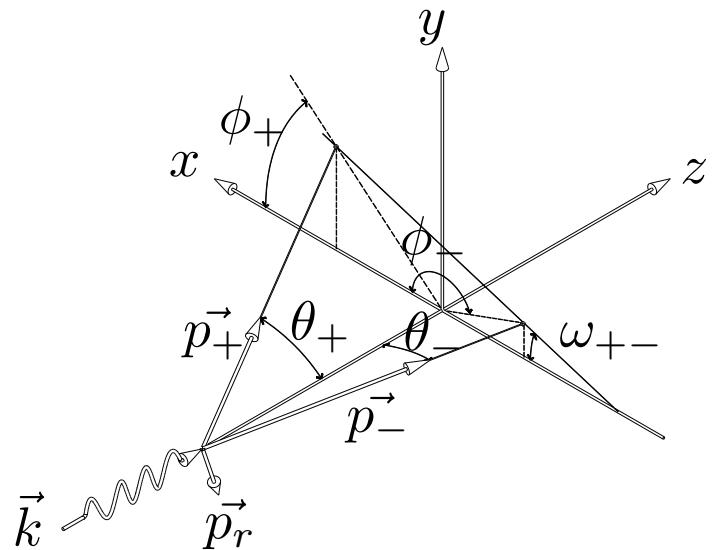
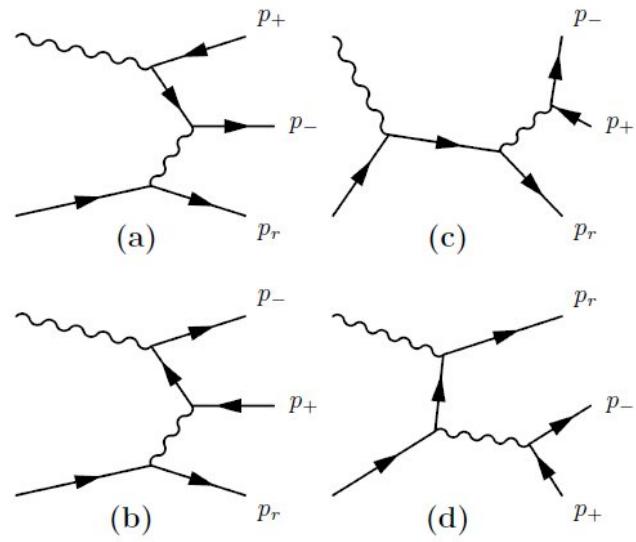
# *An exact, 5D, Polarized Generator*

- 5D final state:  $\varphi_+, \theta_+, \varphi_-, \theta_-, x_+ \equiv E_+/E_\gamma$
- BASES / SPRING version of the VEGAS method, [S. Kawabata, Comp. Phys. Comm. 88, 309 \(1995\)](#).
- Differential cross section either from:
  - Feynman diagrams (important for low energy triplet) HELAS, [H. Murayama, KEK-91-11](#).
  - Bethe-Heitler approximation (2 dominant diagrams only)
- Exact:
  - no low energy approximation
  - no small angle approximation
  - 5D differential Xsection sampled, no product of 1D differential Xsections
  - strict energy-momentum conservation
- Polarized:
  - linearly polarized photons
  - unpolarized photons
  - partially polarized photons
- Target: nucleus or electron (triplet conversion)
- Atomic electron field screening: form factor  $F(q^2)$ , coherent (nuclear), incoherent (triplet)
- (2012 - 2013): extensive validations for high-energy photons,  $E = 4 - 400$ , MeV

[NIM A 729 \(2013\) 765](#)

# A full (5D) exact (down to threshold) polarized evt generator

- Variables: azimuthal ( $\phi_+$ ,  $\phi_-$ ) and polar ( $\theta_+$ ,  $\theta_-$ ) angles of  $e^+$  and  $e^-$ , and  $x_+ \equiv E_+/E$



- Diagrams

- (a), (b) dominant, either for nuclear, or triplet at high-energy.
- in addition, for triplet, 4 additional “exchange” diagrams.

# Bethe-Heitler differential cross section: 1

- Linearly polarized gamma rays:  $(c = 1, \hbar = 1, E = \hbar\omega = \omega)$

$$\begin{aligned} d\sigma = & \frac{-2\alpha Z^2 r_0^2 m^2}{(2\pi)^2 \omega^3} dE_+ d\Omega_+ d\Omega_- \frac{|p_-||p_+|}{|\vec{q}|^4} \left[ \left( 2E_+ \frac{p_- \sin \theta_- \cos (\psi + \phi)}{E_- - p_- \cos \theta_-} + 2E_- \frac{p_+ \sin \theta_+ \cos \psi}{E_+ - p_+ \cos \theta_+} \right)^2 \right. \\ & - q^2 \left( \frac{p_- \sin \theta_- \cos (\psi + \phi)}{E_- - p_- \cos \theta_-} - \frac{p_+ \sin \theta_+ \cos \psi}{E_+ - p_+ \cos \theta_+} \right)^2 \\ & \left. - \omega^2 \frac{(p_+ \sin \theta_+)^2 + (p_- \sin \theta_-)^2 + 2p_+ p_- \sin \theta_+ \sin \theta_- \cos \phi}{(E_- - p_- \cos \theta_-)(E_+ - p_+ \cos \theta_+)} \right] \end{aligned}$$

with:  $|\vec{q}|^2 = |\vec{p}_+ + \vec{p}_- - \vec{k}|^2$ ,  $\psi \equiv \varphi^+$  and  $\psi + \phi \equiv \varphi^-$

"On the Polarization of High Energy Bremsstrahlung and of High Energy Pairs", M. May, Phys. Rev. 84, 265 (1951).  
Correction by a factor of 2: Jauch and Rohrlich, *The theory of photons and electrons* (Springer Verlag, 1976).

- Non polarized gamma rays:

$$\begin{aligned} & \frac{-\alpha Z^2 r_0^2 m^2}{(2\pi)^2 \omega^3} dE_+ d\Omega_+ d\Omega_- \frac{|p_-||p_+|}{|\vec{q}|^4} \left[ \left( \frac{p_+ \sin \theta_+}{E_+ - p_+ \cos \theta_+} \right)^2 (4E_-^2 - q^2) + \left( \frac{p_- \sin \theta_-}{E_- - p_- \cos \theta_-} \right)^2 (4E_+^2 - q^2) + \right. \\ & \left. \frac{2p_+ p_- \sin \theta_+ \sin \theta_- \cos \phi}{(E_- - p_- \cos \theta_-)(E_+ - p_+ \cos \theta_+)} (4E_+ E_- + q^2 - 2\omega^2) - 2\omega^2 \frac{(p_+ \sin \theta_+)^2 + (p_- \sin \theta_-)^2}{(E_+ - p_+ \cos \theta_+)(E_- - p_- \cos \theta_-)} \right] \end{aligned}$$

"The quantum theory of radiation", W. Heitler, 1954.

# Bethe-Heitler differential cross section: 2

- That is:  $d\sigma = \Phi(X_u + P \times X_p) dE_+ d\Omega_+ d\Omega_-$  with  $\Phi = \frac{-\alpha Z^2 r_0^2 m^2}{(2\pi)^2 \omega^3} \frac{|p_-||p_+|}{|\vec{q}|^4}$
- $P$  the polarization fraction of the incident photon,
- And:

$$X_u = \left[ \left( \frac{p_+ \sin \theta_+}{E_+ - p_+ \cos \theta_+} \right)^2 (4E_-^2 - q^2) + \left( \frac{p_- \sin \theta_-}{E_- - p_- \cos \theta_-} \right)^2 (4E_+^2 - q^2) + \frac{2p_+ p_- \sin \theta_+ \sin \theta_- \cos \phi}{(E_- - p_- \cos \theta_-)(E_+ - p_+ \cos \theta_+)} (4E_+ E_- + q^2 - 2\omega^2) - 2\omega^2 \frac{(p_+ \sin \theta_+)^2 + (p_- \sin \theta_-)^2}{(E_+ - p_+ \cos \theta_+)(E_- - p_- \cos \theta_-)} \right]$$

$$X_p = \cos 2(\phi + \psi) (4E_+^2 - q^2) \left( \frac{p_- \sin \theta_-}{E_- - p_- \cos \theta_-} \right)^2 + \cos 2\psi (4E_-^2 - q^2) \left( \frac{p_+ \sin \theta_+}{E_+ - p_+ \cos \theta_+} \right)^2 + 2 \cos(\phi + 2\psi) (4E_+ E_- + q^2) \frac{p_- \sin \theta_- p_+ \sin \theta_+}{(E_- - p_- \cos \theta_-)(E_+ - p_+ \cos \theta_+)}$$

- Partial screening effect of (other) atomic electrons parametrized by form factor (fn of  $q$ )

# Gears

- Variables that are actually generated:

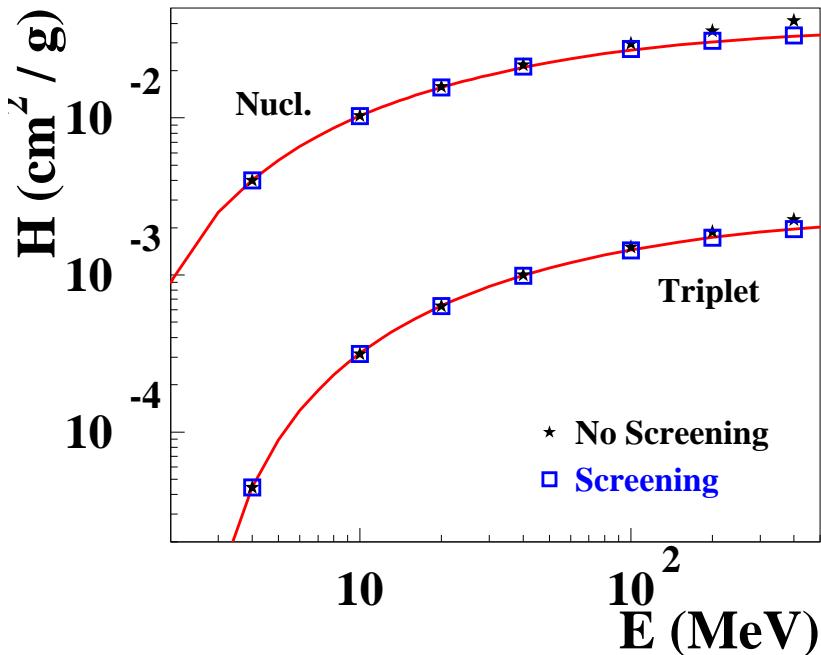
variable	related to	
$\lambda(1)$	recoil polar angle	in the CMS
$\lambda(2)$	recoil azimuthal angle	
$\lambda(3)$	pair Invariant Mass(MeV)	
$\lambda(4)$	positron polar angle	in the pair CMS frame
$\lambda(5)$	positron azimuthal angle	in the pair CMS frame

- Output:

- $e^+$ ,  $e^-$  and recoil 4-vectors

CMS : center of mass system

## Benchmark: 0: cross sections (argon)



- Comparison of the total mass attenuation coefficients (argon; with and without screening) with the data from NIST
- NIST data include all known effects (screening, Coulomb corrections ..)

J. H. Hubbell et al., J. Phys. Chem. Ref. Data 9, 1023 (1980).

NIM A 729 (2013) 765

# Benchmark: 1: triplet / nuclear $q$ -distribution ratio

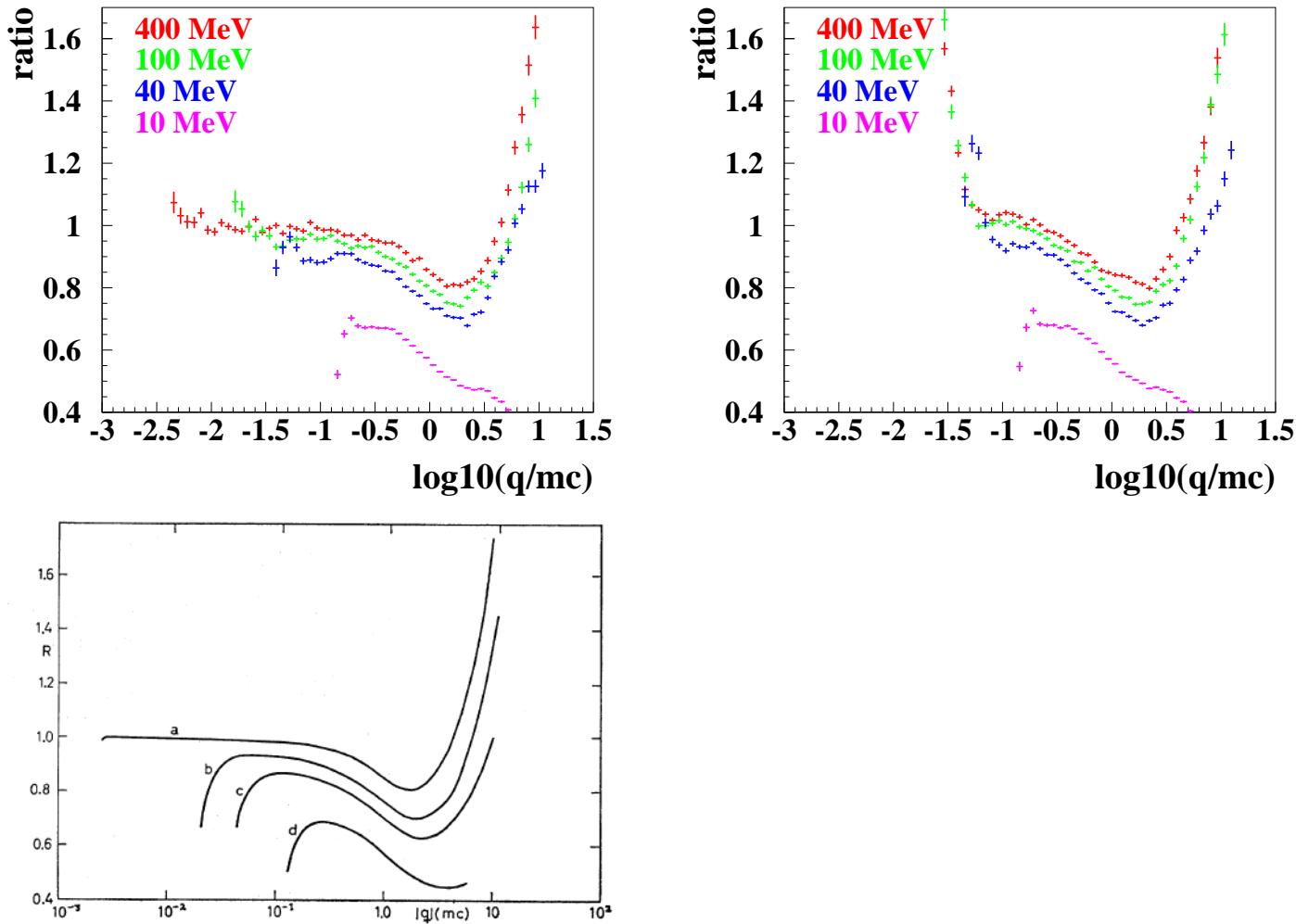


FIG. 7. The ratio  $R$  between the recoil distributions for pair production by electrons and by a very heavy nucleus. The curves  $a$ ,  $b$ ,  $c$ , and  $d$  refer to lab photon energies  $\omega = 1000$ ,  $100$ ,  $50$ , and  $20 mc^2$ , respectively.

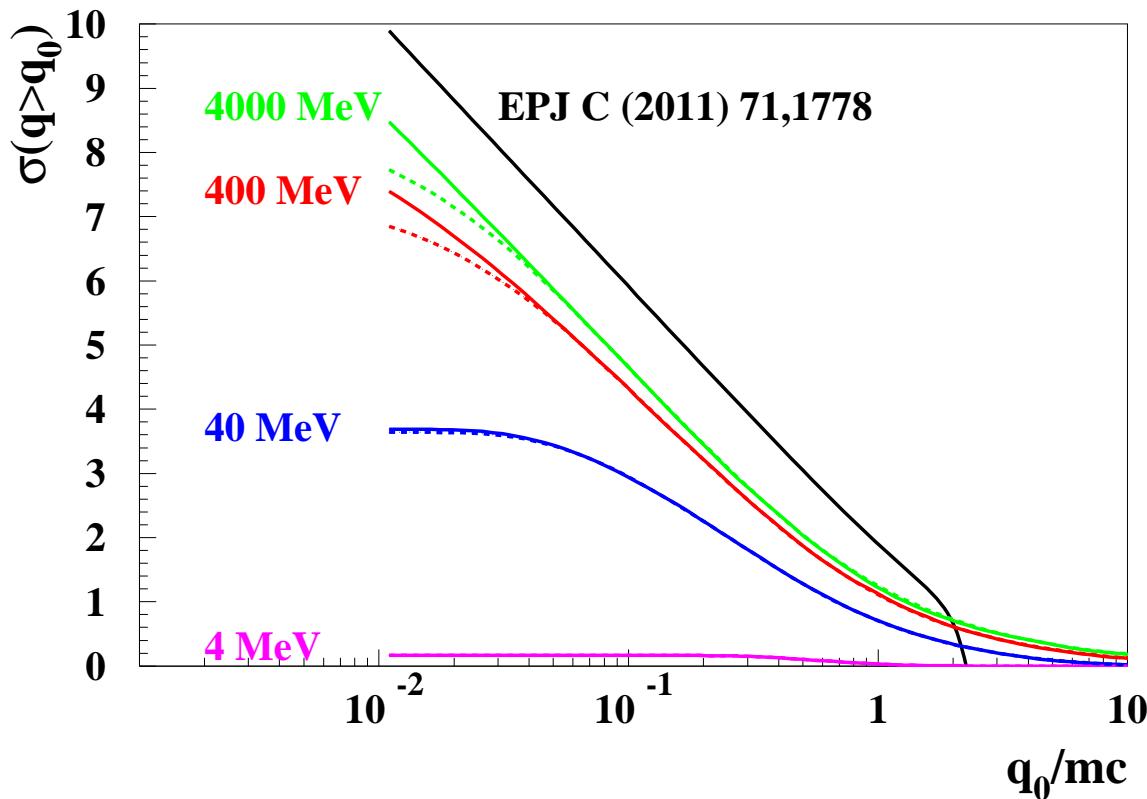
Without screening

with screening

"Pair Production by Photons on Electrons", K. J. Mork, Phys. Rev. 160, 1065 - 1071 (1967).

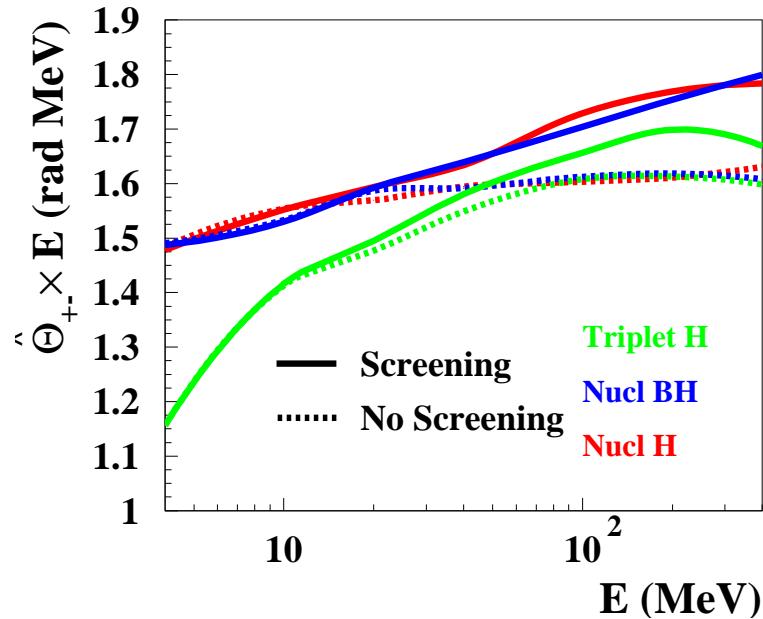
## Benchmark: 2: triplet cross section above threshold $q_0$

- Triplet conversion: cross section for recoil electron momentum larger than  $q_0$ ,  $\sigma(q > q_0)$ , as a function of  $q_0/mc$ , for various photon energies  $E$ ; (Dashed is with form factor applied).
- High photon energy asymptotic expression by M. L. Iparraguirre and G. O. Depaola, Eur. Phys. J. C 71, 1778 (2011).

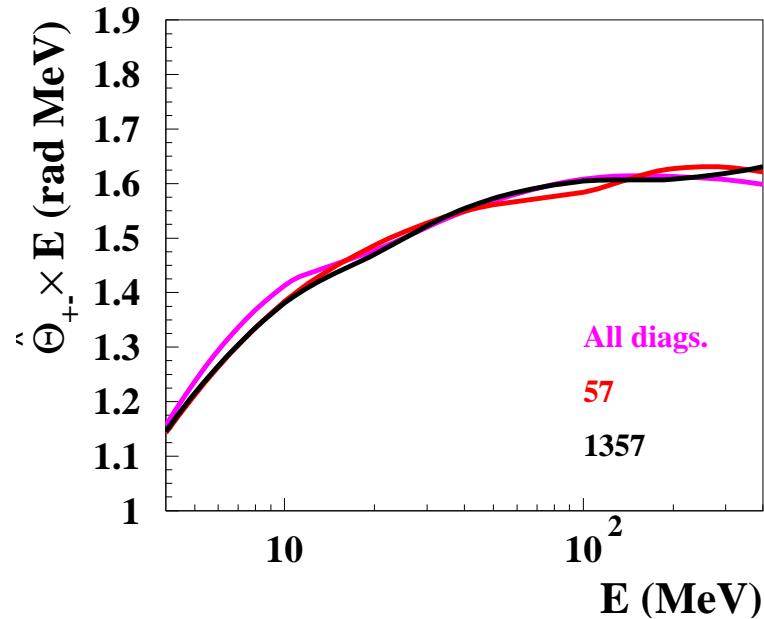


D.B. NIM A 729 (2013) 765

# Benchmark: 3: Most probable opening angle $\hat{\theta}_{+-}$



nuclear and triplet  
with/without screening  
for the H/BH pdf



triplet

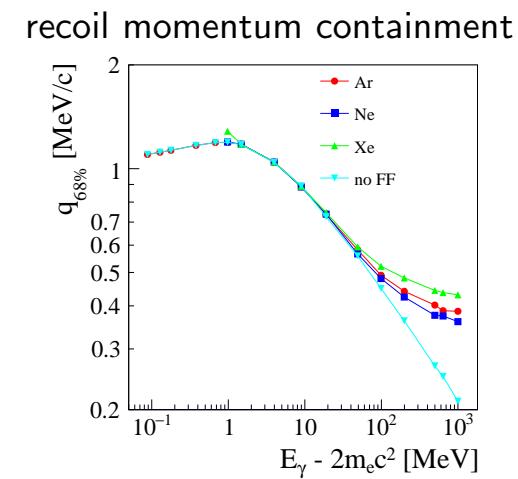
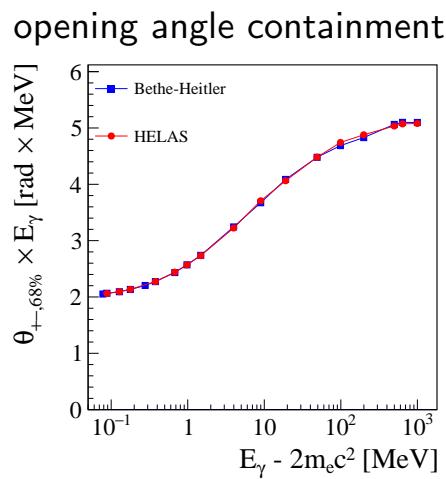
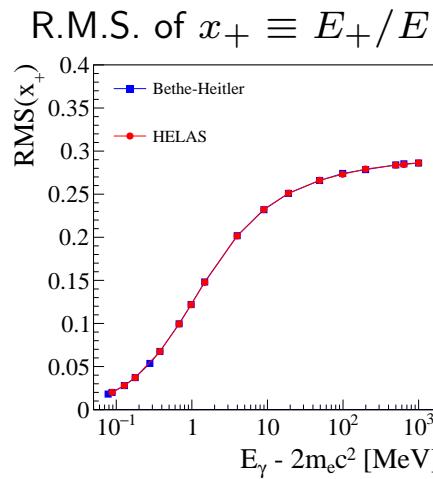
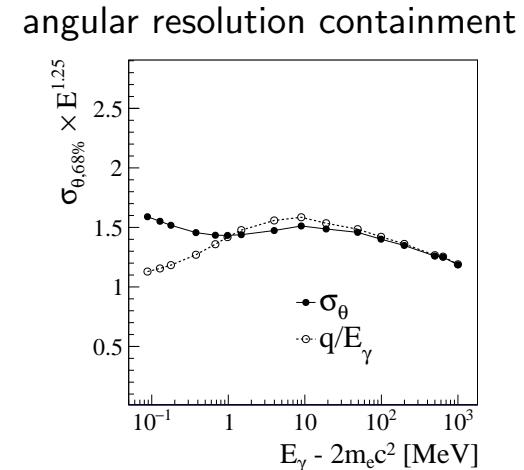
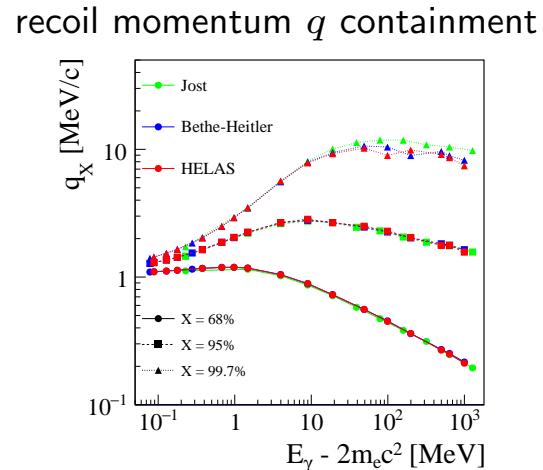
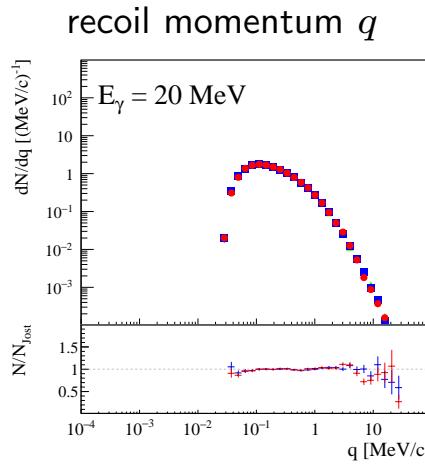
with all diagrams  
without exchange diagrams (1357)  
with BH diagrams only (57)

$$\text{Olsen computes } \hat{\theta}_{+-} \approx \frac{1.6 \text{ MeV}}{E}$$

“Opening Angles of Electron-Positron Pairs”, Haakon Olsen, Phys. Rev. 131, 406 - 415 (1963)

# VEGAS-based generator: more on validations

- Extend to 1.1 MeV - 1 GeV



R. Jost, "Distribution of Recoil Nucleus in Pair Production by Photons," Phys. Rev. 80, 189 (1950).

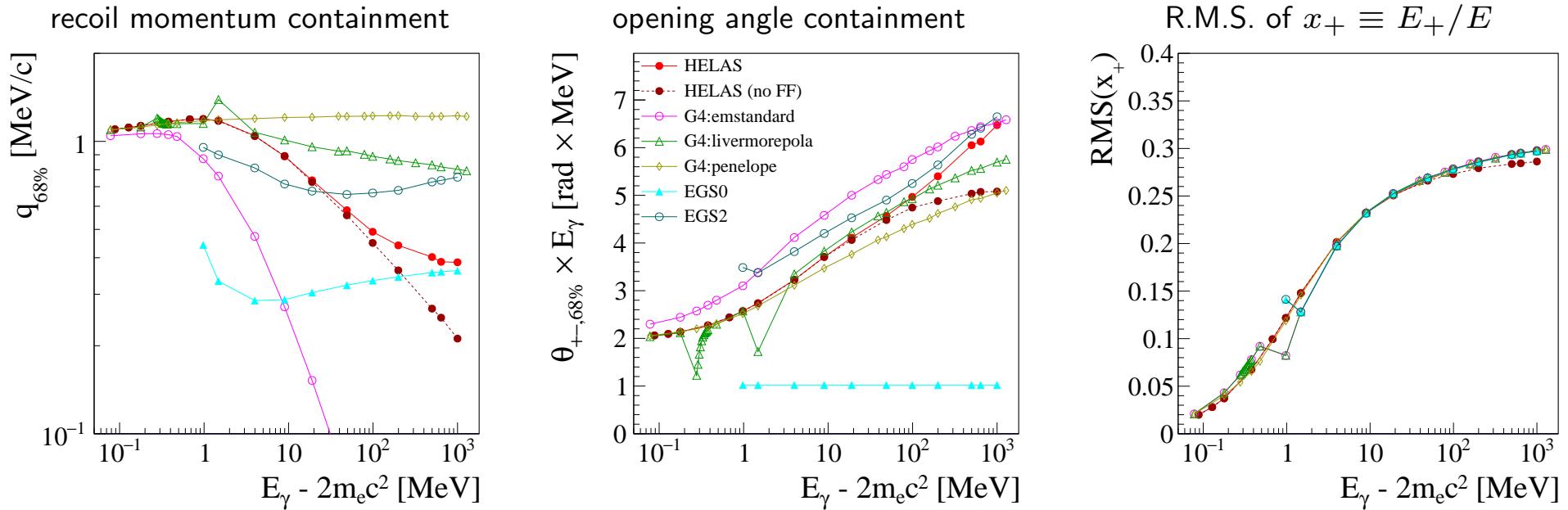
Astropart. Phys. 88 (2017) 60

# *Conversion Event Generators: A Comparison*

Name	Model	Generator
<i>HELAS</i>	HELAS Feynman amplitudes	BASES/SPRING
<i>Bethe-Heitler</i>	Bethe-Heitler	BASES/SPRING
<i>G4:emstandard</i>	G4BetheHeitler	Geant4 10.02.01
<i>G4:livermorepola</i>	G4LivermorePolarizedGammaConversion	Geant4 10.02.01
<i>G4:penelope</i>	G4PenelopeGammaConversion	Geant4 10.02.01
<i>EGS0</i>	egs5, IPRDST= 0	egs5 1.0.6
<i>EGS2</i>	egs5, IPRDST= 2	egs5 1.0.6

- *G4:livermore* has same kinematics as *G4:emstandard* : not shown here

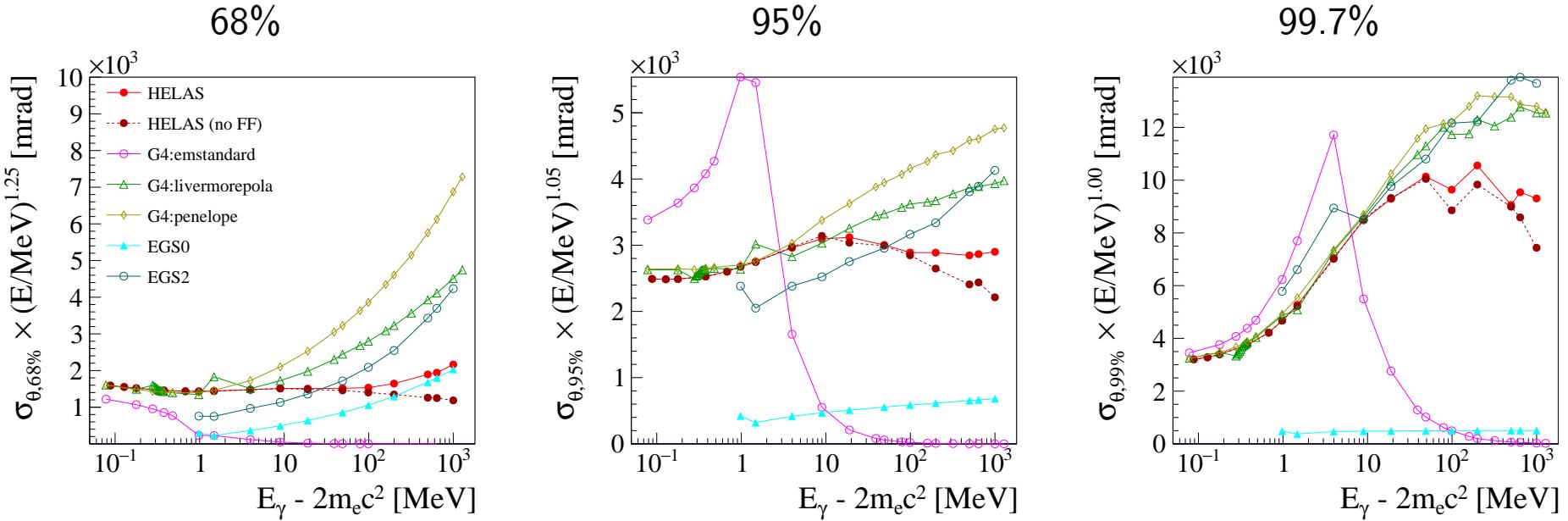
# Event Generator Comparison: Kinetic Variables



- Recoil momentum **incorrectly** simulated by G4 and EGS5 physics models
- Inspection of code and/or documentation shows that they (most often) don't even conserve energy-momentum.

Astropart. Phys. 88 (2017) 60

# Event Generator Comparison: Angular Resolution

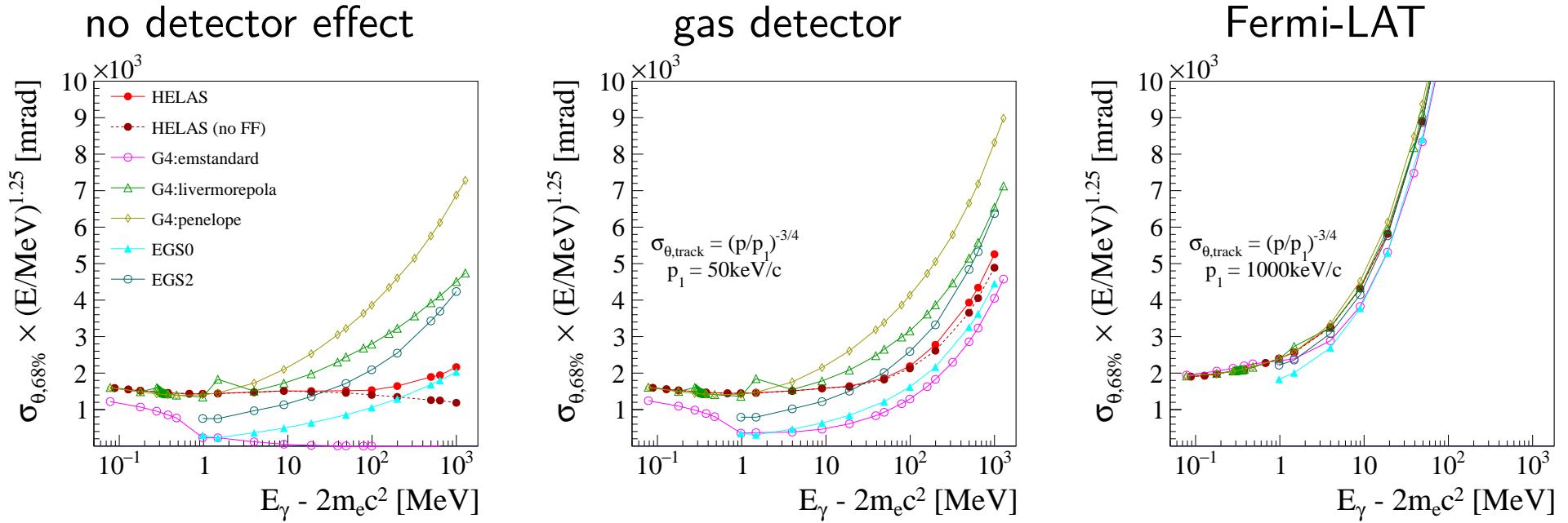


- HELAS: Form Factor (argon): an effect for  $E > 100\text{MeV}$
- HELAS: parametrization of  $\sigma_\theta$ :
 

68%	$1.5 \text{ rad } (E/\text{MeV})^{-1.25}$
95%	$2.9 \text{ rad } (E/\text{MeV})^{-1.05}$
99.7%	$4 - 9 \text{ rad } (E/\text{MeV})^{-1.00}$
- G4 and EGS5 models have 68% angular resolution **crazy**.
  - Most often,  $e^+$  and  $e^-$  generated back-to-back and polar angles generated independently : does not conserve energy-momentum.

Astropart. Phys. 88 (2017) 60

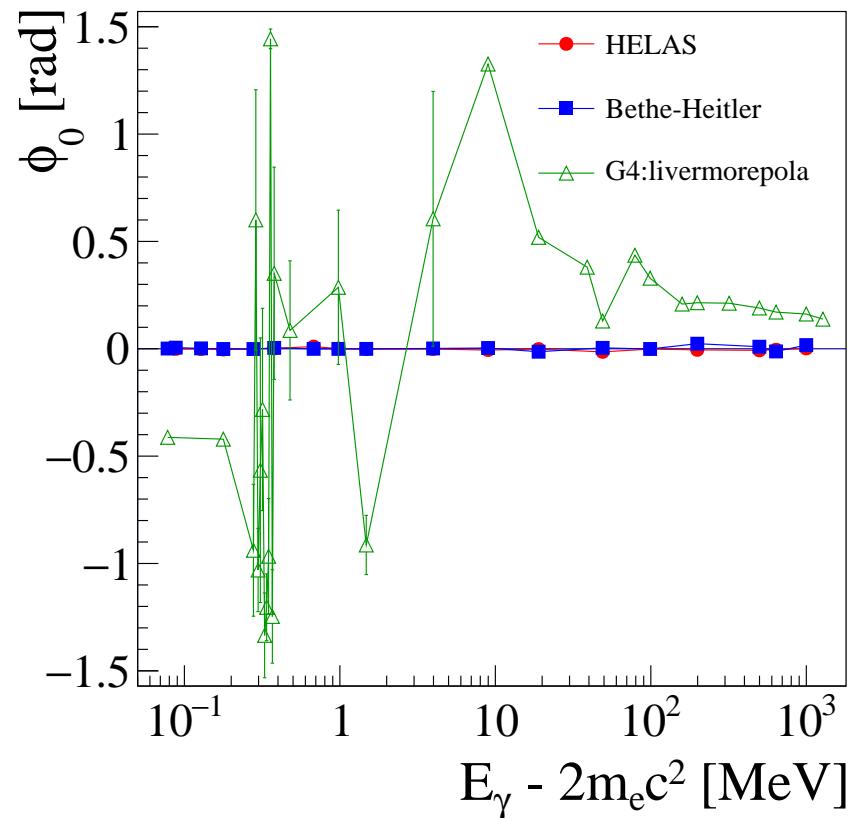
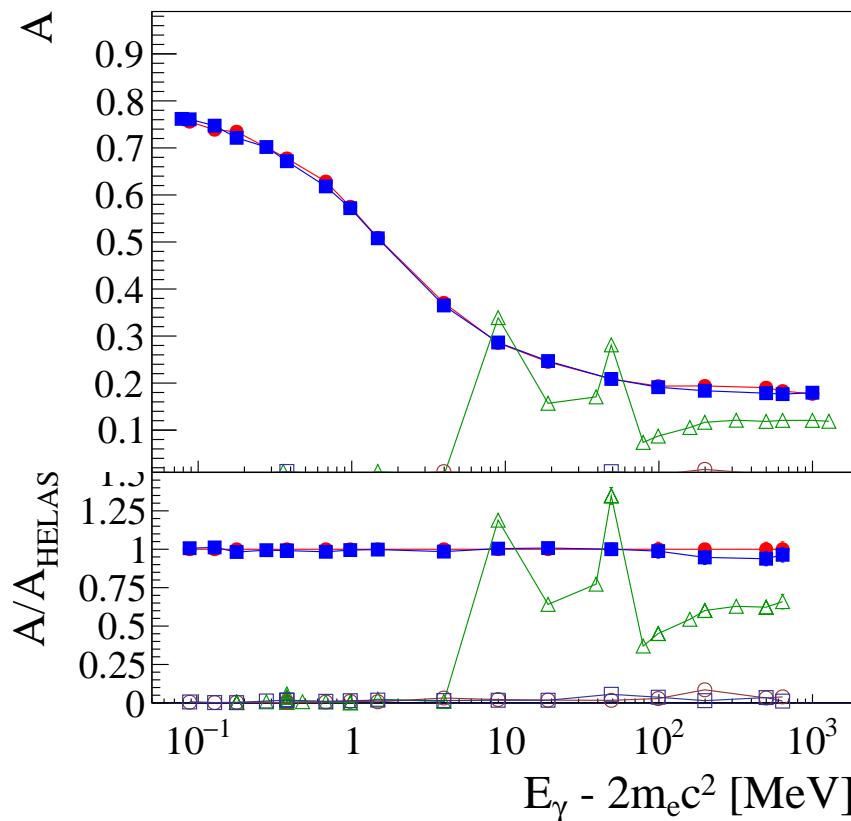
# Event Generators and Telescopes



- For dense, high- $Z$  telescopes, the multiple scattering washes out track correlations and therefore the event generator differences

Astropart. Phys. 88 (2017) 60

# Event Generators and Photon Linear Polarization



- Simulation of polarized  $\gamma$  conversion to pairs by G4:livermorepol **surprising** (polarisation asymmetry and polarisation angle)

$$\text{@ } 100 \text{ MeV, } \frac{\mathcal{A}_{(\text{HELAS or BH})}}{\mathcal{A}_{\text{G4:livermorepol}}} = \frac{(19.1 \pm 0.4)\%}{(8.7 \pm 0.6)\%} \approx 2.2$$

[Astropart. Phys. 88 \(2017\) 60](#)

# *Multi-dimensional integration and event generation with package BASES/SPRING*

- Method: The VEGAS algorithm, due to G. P. Lepage ([J. Comput. Phys. 27 \(1978\) 192](#))
- Integration:
  - $n$ -D hyper-volume first segmented in identical cells ( $n$ -dimensional bins);
  - pdf and its variance evaluated in each cell, by shooting a number of points;
  - hyper-volume segmentation then tuned iteratively, so as to minimize the variance;
  - upon convergence, an optimal tabulation of the pdf is obtained.
- Event generation:
  - an “exact” generation from  $f(\lambda)$  is then obtained easily from the tabulated  $f_0(\lambda)$  the acceptance-rejection method
- The pdf depends on photon energy  $E$  (and target nature through  $F(q^2)$ )
  - the integration step is time consuming (few seconds)
  - generation of batches of millions events with same energy easy.
- Presently moving to a parametrized( $E$ ) version for each variable.

No sign of (Bethe-Heitler)-HELAS difference: **HELAS dropped**

# *Perspectives for the future: a Geant4 Physics Model*

- Fortran → C++ → private G4EmModel
  - G4 experts expertise will be welcome.
- Donation to the Geant4 Collaboration if agreeable by the G4 Coll.

# *Conclusion*

- Presently available physical models on the market inappropriate:
  - for simulation of high-resolution  $\gamma$  telescopes with pairs
  - for simulation of  $\gamma$  polarimeters with pairs
- Electron/ positron energy share distribution found OK.
- VEGAS-based exact, 5D, polarized event photon conversion event generator built, validated, published.
- De-VEGAS-ification in progress → (several)- $\mu$ s single-photon generation validation already extended to  $E = 1 \text{ PeV}$
- Geant4 Physics Model later this year.

# *Back-up slides*

## Form Factors

- Form factors must be taken into account as the fraction of the total cross section lost increase with energy :

$$\frac{\sigma_{\text{noFF}}}{\sigma_{\text{FF}}} = \frac{41 \log(2E/m) - 109}{41 \log(183Z^{-1/3}) - 1} \quad (1)$$

- for 2 elements  $Z_1$  and  $Z_2$ ;

noting that  $(1 - F(q^2, Z_1)) < (1 - F(q^2, Z_2))$  for  $Z_1 > Z_2$ ,

- use  $Z_2$  mock-up to shoot  $Z_1$
- The efficiency loss would be asymptotically

$$\frac{\sigma_{\text{FF},Z_1}}{\sigma_{\text{FF},Z_2}} = \frac{41 \log(183Z_1^{-1/3}) - 1}{41 \log(183Z_2^{-1/3}) - 1}, \quad (2)$$

For example,  $\frac{\sigma_{\text{FF},Ar}}{\sigma_{\text{FF},Xe}} \approx 1.095$ . (after  $Z^2$  dependence rescaled).

# Polarisation asymmetry : asymptotic expressions

- Low energy

$$A = \frac{\pi}{4}.$$

Astropart. Phys. 88 (2017) 30

- High energy

$$A \approx \frac{\frac{4}{9} \ln 2E - \frac{20}{28}}{\frac{28}{9} \ln 2E - \frac{218}{27}}$$

V. F. Boldyshev et al., Yad. Fiz. 14 (1971) 1027, (Sov.J.Nucl.Phys. 14 (1972) 576).

