Crystal physics and phonon transport in Geant4

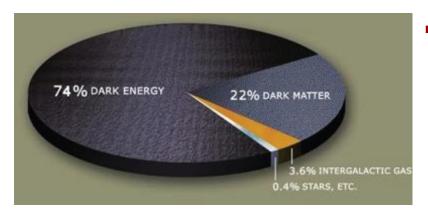
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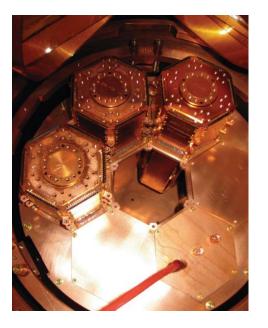


The Cryogenic Dark Matter Search (CDMS)



- CDMS attempting to detect WIMP interactions with crystals
- Deep underground 15kg Ge@50mK
- Pretty much opposite of a space application

- The majority of the mass energy in the Universe is dark matter
 - » No EM interaction
 - » Gravitational interaction
 - » Non-baryonic



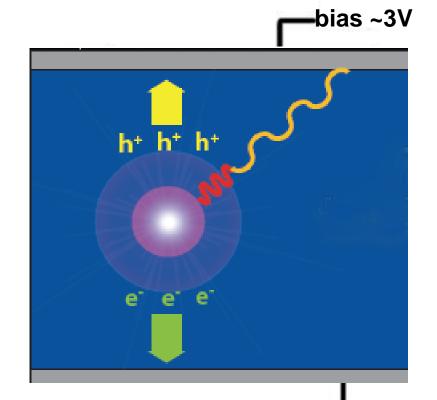






Motivation: SuperCDMS Detector Monte Carlo

- A particle scattering in a crystal will create both phonons and electron-hole pairs
- Electron-hole pairs are collected by a small drift field
- Phonons are collected by Transition Edge Sensors (TESs)



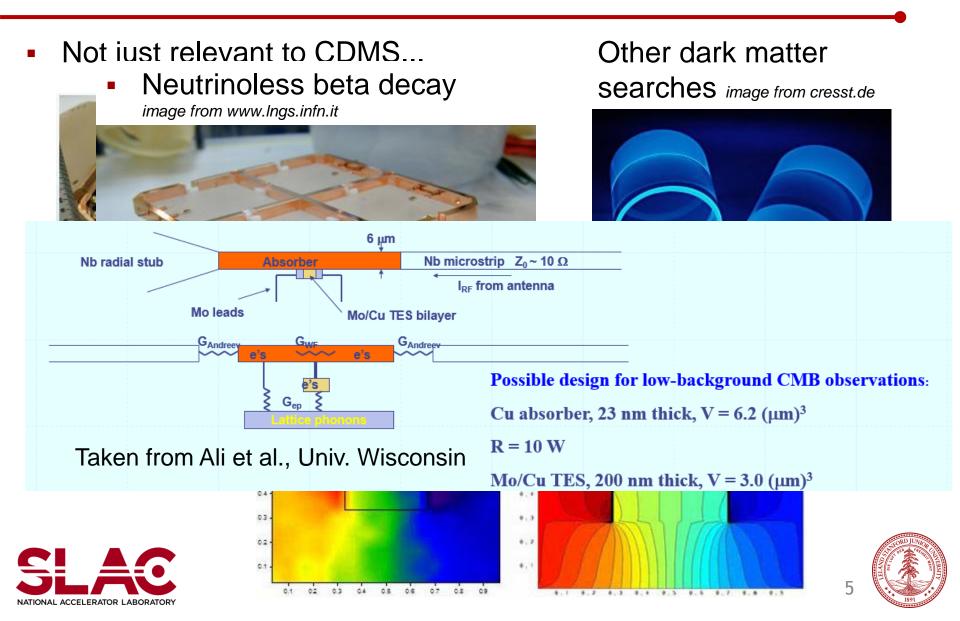
Deposited by WIMP: 10-100 keV e-/hole pair: 3 eV to create Individual phonon: ~80 meV to create



6 March 2013, G4SUW, Barcelona

TELA TELA

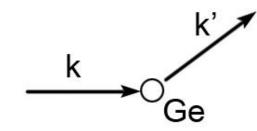
Motivation: Community interest



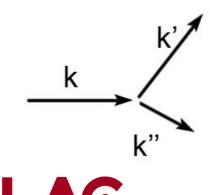
Physics processes

The CDMS detectors operate at 40 mK. Two phonon processes are relevant.

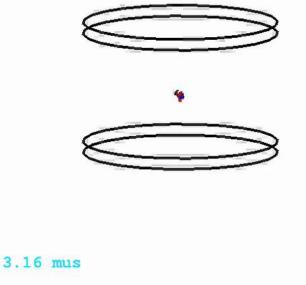
1) Isotope scattering and mode mixing



2) Anharmonic down conversion



ACCELERATOR LABORATO





Physics processes

Isotope scattering and mode mixing

XPhononScatteringProcess : G4VDiscreteProcess

Anharmonic downconversion

XPhononDownconversionProcess : G4VDiscreteProcess

Phonon focusing using singleton lattice manager object

XLogicalLattice(*dynamic constants*) XPhysicalLattice(&XLogicalLattice, *orientation within volume*) XLatticeManager3::registerLattice(&G4VPhysicalVolume, &XPhysicalLattice)

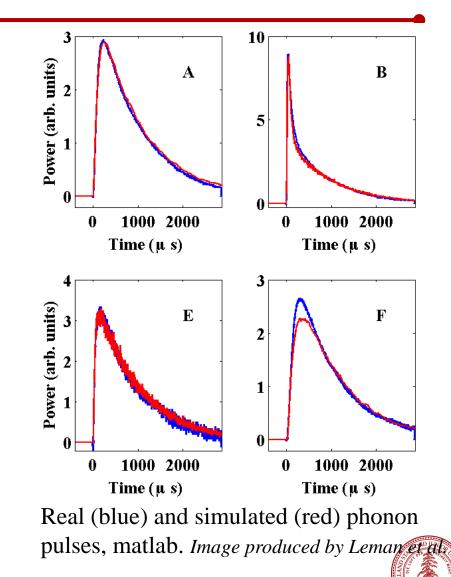






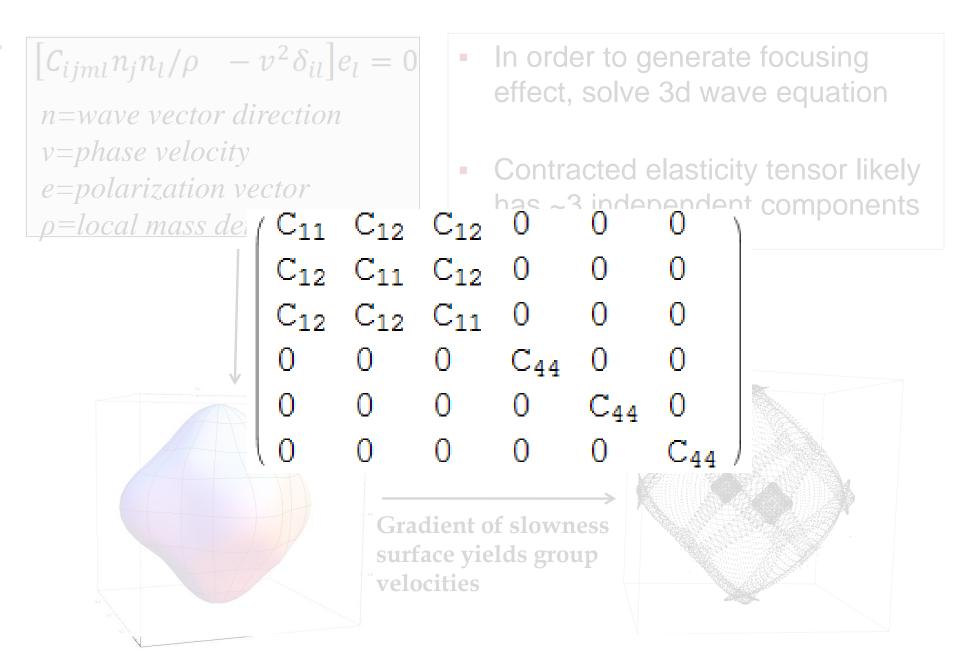
Simulating the CDMS Signal

- Simulated phonon pulse shapes match well with observed pulse shapes
- Pulse shapes shown are simulated TES traces
- TES simulation currently performed in matlab
- Matlab provides output in CDMS-raw format

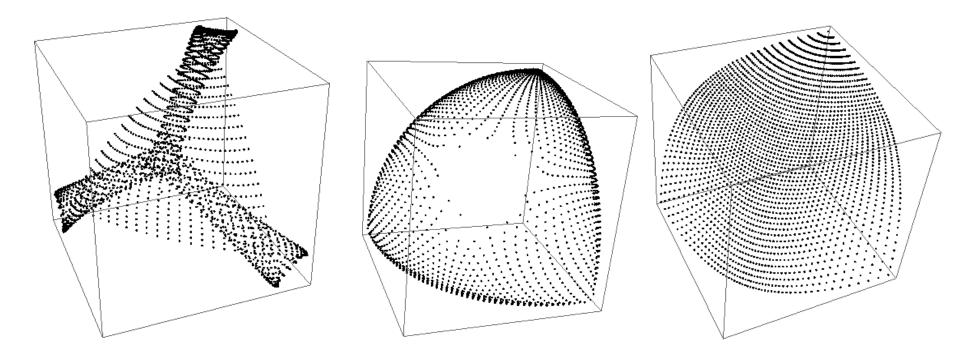




Phonon focusing



Group velocity densities



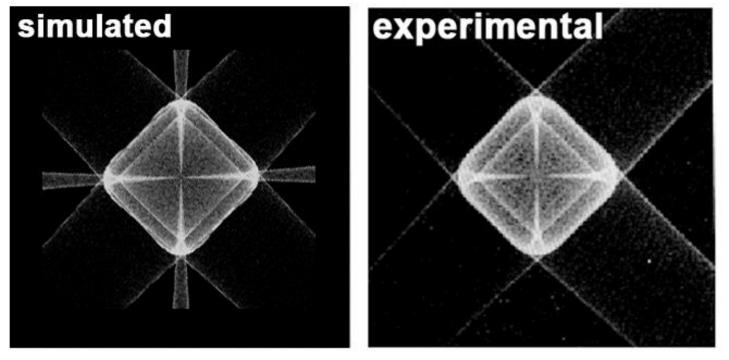
 Densities of group velocity vectors, left to right: Longitudinal, slow transverse, fast transverse





Phonon caustics in Germanium

 Anisotropies in the elasticity tensor cause phonons to be focused onto preferred propagation vectors.



Phonon flux intensity on a Ge crystal face due to a point source at the crystal center. Left: simulated with Geant4 **Right**: as observed by Nothrop and Wolfe





Phonon example in Geant4 v9.6

Phonon transport code available as example since v9.6.0

examples/extended/exoticphysics/phonon

 In its current form the example stores phonon arrival times/eneriges and xy locations in two .ssv files

examples/extended/exoticphysics/phonon/timing.ssv examples/extended/exoticphysics/phonon/caustic.ssv

 The sort of caustics shown on the previous slide can be generated by reducing the initial phonon energy into the ballistic regime and creating a 2D histogram of *caustic.ssv*





Generating caustics using scorers - I

Can use Geant4 scorers to create caustics. To enable scorers add the following to XGeBox.cc:

#include "G4ScoringManager.hh"

• • •

G4ScoringManager scoreMan = G4ScoringManager::GetScoringManager();

Change the phonon energy in XPrimaryGeneratorAction:

void XPrimaryGeneratorAction::GeneratePrimaries(G4Event*)

particleGun->SetParticleEnergy(0.0001*eV);





• Compile, run XGeBox and run the scoring script:

/score/create/boxMesh mesh1 /score/mesh/boxSize 0.01 0.01 0.01 m /score/mesh/nBin 31 31 31 /score/quantity/flatSurfaceCurrent myQuant /score/close

/run/beamOn 10000

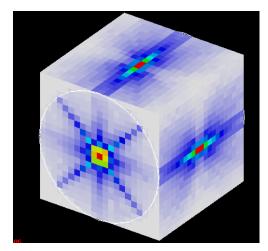
/score/drawProjection mesh1 myQuant

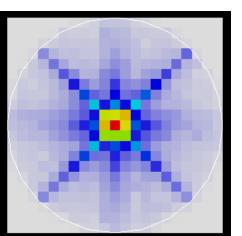


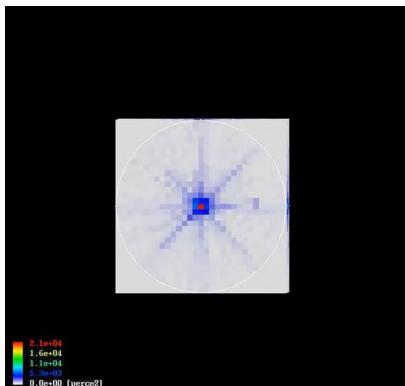


Generating caustics using scorers - III

- These images were generated with the geant4 internal scorer outlined in the previous slide
- May need /scorer/colorMap/setMinMax











Conclusions

- Phonon transport at zero temperature implemented
- Phonon transport code available as example since v9.6.0
- Good agreement with experiment
- Can generate caustics using Geant4 scorers
- Can simulate any crystal whose elasticity tensor can be Voigt-contracted to 6x6 tensor
- Working on non-zero temperatures and optical phonons





Future Developments

- Generating phonons from physics processes
- Phonon transport at non-zero temperature
- Physical implementation of boundary interactions (including mode-matching for different crystals)
- Optical phonon processes
- Phonon transport in metals and semi-metals







The SuperCDMS collaboration



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Backup slides



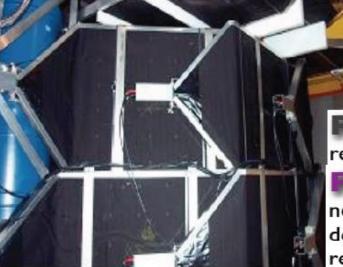


SLAC Annual Budget Briefing for OHEP

2100 mwe underground



rejects events from cosmic rays



Pb: shielding from gammas resulting from radioactivity

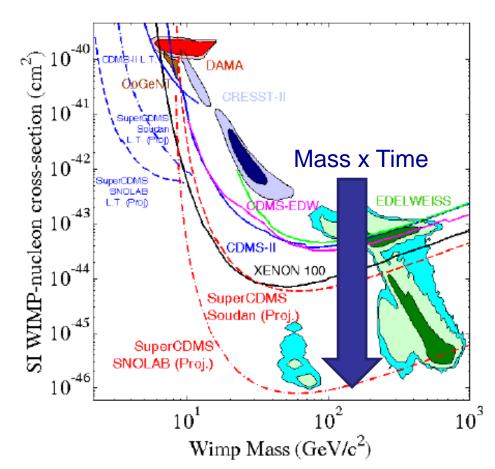
Polyethyene: moderate neutrons produced from fission decays and from (α,n) interactions resulting from U/Th decays

Ancient lead

Passive Shielding

Current and Projected Limits

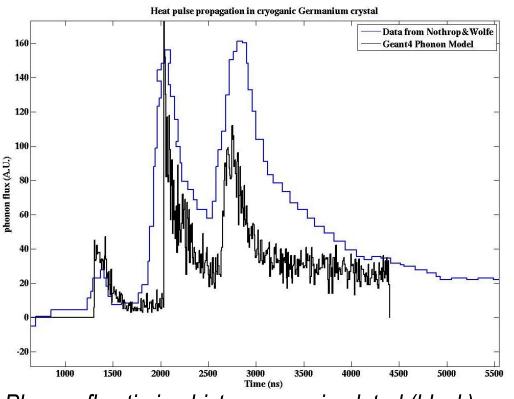
- SuperCDMS Soudan will match current XENON 100 limit
- SuperCDMS Lite will produce world leading low-mass limits
- SuperCDMS SNOLAB to improve limit by two orders of magnitude







Validation of transport code



Phonon flux timing histograms, simulated (black) and observed by Nothrop&Wolfe (blue)

•Heat pulse propagation is a good test of phonon transport model

•Simulated heat pulse reproduces three peaks

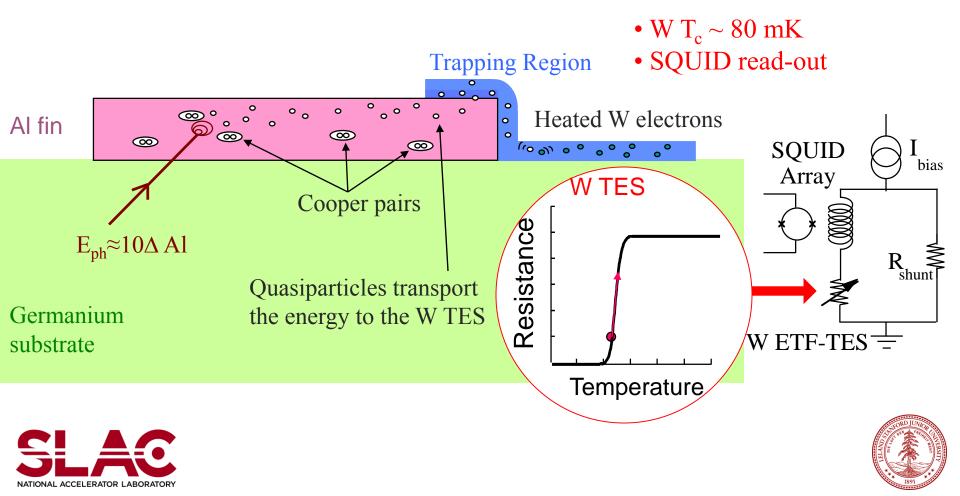
•Simulation yields right branching ratios

•Discrepancies in onset time and Slow Transverse fall off are due to laser pulse shape and e⁻/h⁺ recombination

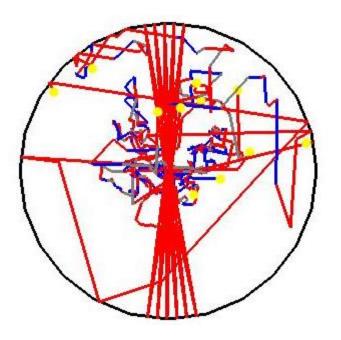


Phonon detection in CDMS

Recoil event occurs in Germanium substrates, 76 mm diameter, 25 mm thick
Aluminum fins 300 nm thick absorb phonons
Fins connect to Tungsten transition edge sensors (W TESs)



Quasi-diffuse Phonon Propagation



•Phonons of different energies have vastly different mean free paths

•Down conversion causes phonons to change from diffuse to ballistic propagation

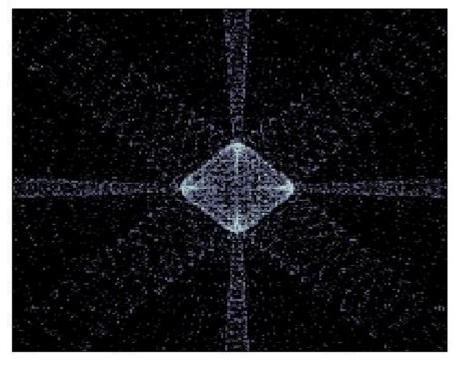
•This kind of quasi-diffuse propagation is difficult to model analytically

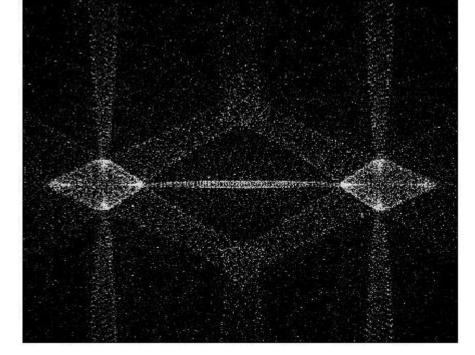




Phonon Focusing

Intensity patterns from a point source of phonons at the crystal center





Ge crystal in [001] orientation

Ge crystal in [111] orientation

28 June 2011 SC



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Custom focusing maps

• Generate the Christoffel Tensor:

Solve[{Det[Dij[{1, 0, 0}] - va] == 0, $v \neq 0$ }, v]

• Find phase velocity:

Build group velocity

$$\left\{ \left\{ \mathbf{v} \rightarrow -\frac{\sqrt{C_{11}}}{\sqrt{\rho}} \right\}, \left\{ \mathbf{v} \rightarrow \frac{\sqrt{C_{11}}}{\sqrt{\rho}} \right\}, \left\{ \mathbf{v} \rightarrow -\frac{\sqrt{C_{44}}}{\sqrt{\rho}} \right\}, \\ \left\{ \mathbf{v} \rightarrow -\frac{\sqrt{C_{44}}}{\sqrt{\rho}} \right\}, \left\{ \mathbf{v} \rightarrow \frac{\sqrt{C_{44}}}{\sqrt{\rho}} \right\}, \left\{ \mathbf{v} \rightarrow \frac{\sqrt{C_{44}}}{\sqrt{\rho}} \right\} \right\}$$

Needs["NumericalCalculus`"]
vgroup[index_, {x_, y_, z_}]:=
 {ND[v[index, {X, y, z}], X, x],
 ND[v[index, {x, Y, z}], Y, y],
 ND[v[index, {x, y, Z}], Z, z]}



vectors:



Motivation: SuperCDMS Detector Monte Carlo

- SuperCDMS background rejection: energy partitioning between phonon and charge channels
- Requires excellent understanding of the phonon signal resulting from real background distributions → Geant4

