



Simulating the Space Environment for Hardware-in- the-Loop Testing of Space Particle Sensor Signal Chains

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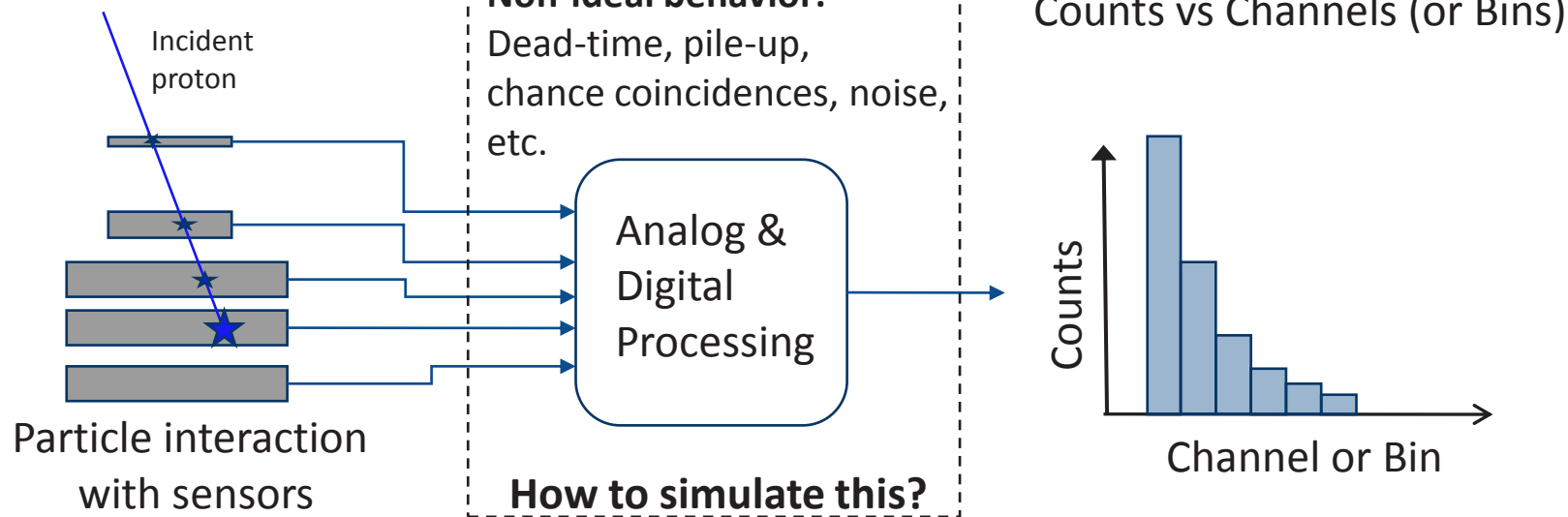
Motivation

Charged particle fluxes can vary greatly in magnetosphere

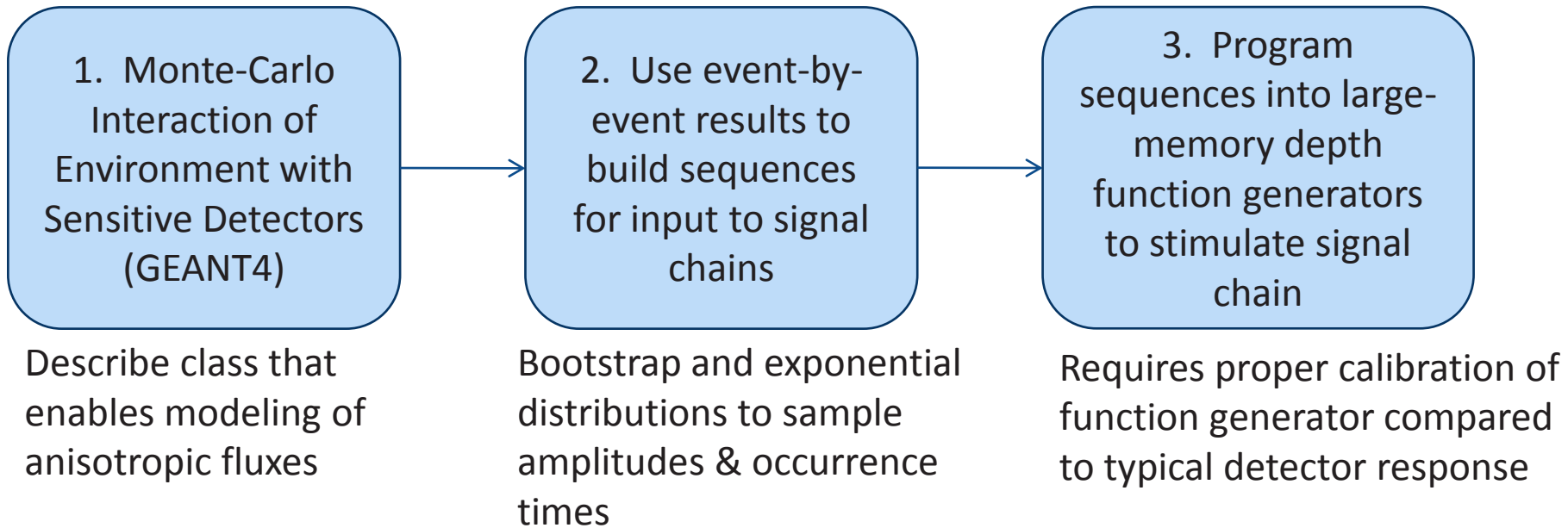
- Wide-dynamic range in particle flux due to location and time
- Relative ratio of charged particle-species (background contamination)
- Anisotropic angular dependence

Places significant demands on charged particle sensors to make measurements over wide range of conditions

Multi-element telescope stack



Simulation Approach



Ultimately goal is to compare reported measurements from the signal chain against particle flux spectrum used to drive the simulation

Modeling of Anisotropic Particle Fluxes

From Sullivan (1971) for a single, time independent particle species for a detector:

$$C_i = \int_0^\infty dE \int_\Omega d\omega \int_S d\vec{\sigma} \cdot \hat{r} \epsilon_i(E, \vec{\sigma}, \omega) j(E, \omega)$$

Labels for the equation above:
 - C_i : Average count rate in i'th channel
 - $\int_0^\infty dE$: Energy domain
 - $\int_\Omega d\omega$: Domain of integration in solid angle
 - $\int_S d\vec{\sigma} \cdot \hat{r}$: Surface element, Solid element dir., Domain over sensor surface
 - $\epsilon_i(E, \vec{\sigma}, \omega)$: Chan. efficiency
 - $j(E, \omega)$: Differential Flux

Make approximation to separate flux into intervals with small variations in angle dist.

$$j(E, \omega) = J(E)F(\omega | E_k)$$

$J(E)$ – Energy dep. amplitude

$F(\omega | E_k)$ – Angular distribution

$\omega - (\theta, \phi)$ spherical polar angles

Can show for anisotropic flux with limited support (in energy) that:

Diff. omni. flux at center energy E_k and interval width ΔE_k

$$C_i \approx \sum_{k=1}^N \int_{E_k}^{E_{k+1}} J_{omni}(E) \tilde{\Gamma}_{ik}(E) dE$$

Labels for the equation above:
 - C_i : Count rate in i'th channel
 - $\tilde{\Gamma}_{ik}(E)$: Normalized gathering power

where:

$$\tilde{\Gamma}_{ik} = \frac{\int_\Omega d\omega \int_S d\vec{\sigma} \cdot \hat{r} \epsilon_i(E, \vec{\sigma}, \omega) F(\omega | E_k)}{\int_\Omega F(\omega | E_k) d\omega}$$

Inputs: Incident flux distribution in magnetic coordinates; Euler angles between magnetic and detector coordinates

Constraints: Source is “flat” surface with constant surface normal

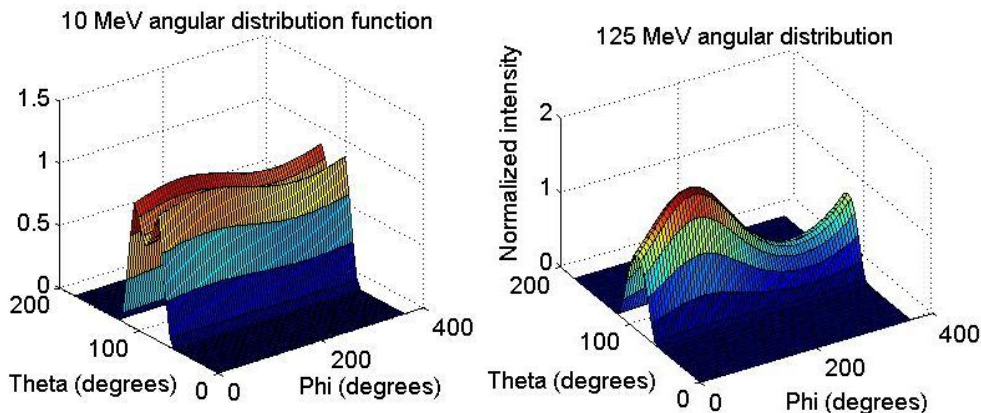
Outputs: (1) Source gathering power (2) Integral of angular dist. (3) Initial momentum direction for each simulated event

$$\Gamma_{ik} = \frac{N_{ik}}{N_{sim}} \Gamma_{source}$$

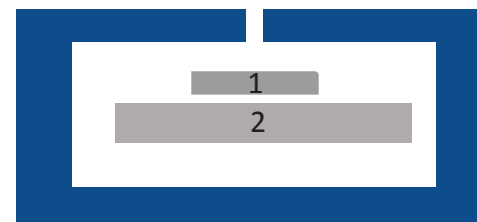
GEANT class allows calculation of normalized gathering power

Example: LEO Proton Response

Angular Distributions

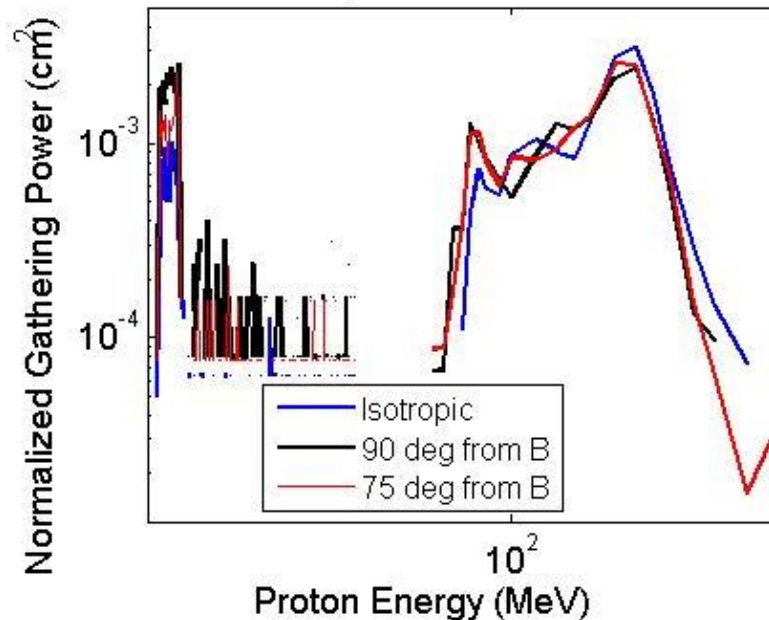
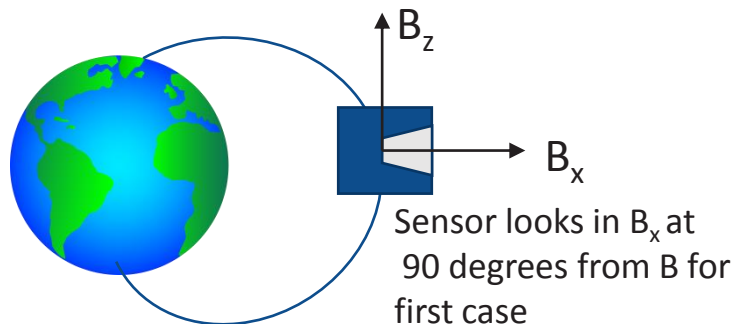


Pinhole type detector that only counts particles that deposit energy in detector 1 (rejects events that deposit 1 & 2)



Lenchek & Singer type model¹ to include East-West effects in low-altitude trapped protons

- Pancake type pitch angle distribution
- Theta is pitch-angle relative to local mag. field; Phi is azimuthal angle
- Theta, phi are spherical polar angles in magnetic frame



1. A.M. Lenchek, S.F. Singer, "Effects of the finite gyroradii of geomagnetically trapped protons", J. Geophys. Res. 67, 4073-4075 (1962).

Simulated Detector & Hardware-in-Loop Test Setup

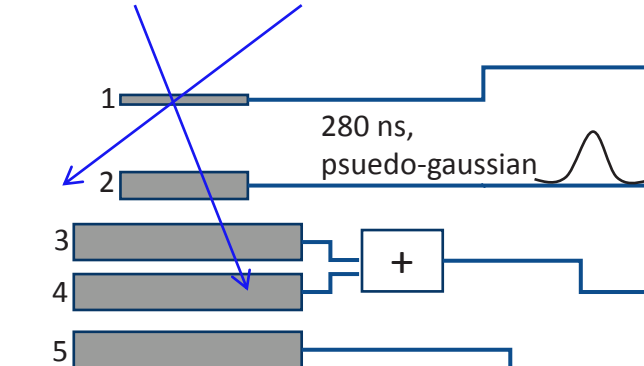
Simulated Detector

- 5 element detector (solid-state)
- Requires coincidence between 1 & 2 for detection of event
- Pulse-height prop. to energy deposited

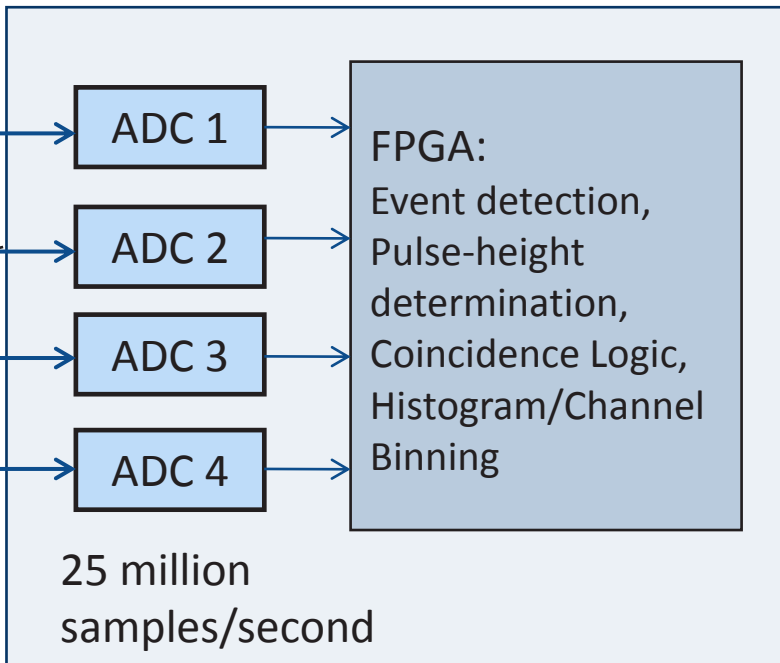
Analog to Digital Conversion

Event Processing

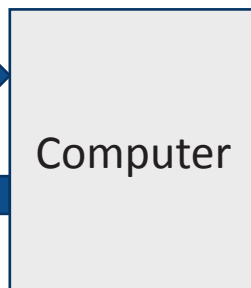
counted not counted



Spectrum-Instrumentation Corp. 4 output, 125 MSPS, 1 Gsample, 14-bit, D/A card used to simulate detector & analog chain



Serial UART (RS232)

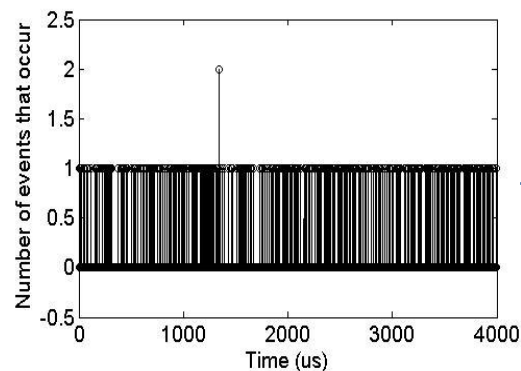


25 million samples/second

Events for electrons binned based on sum of energy deposited in detectors 1 through 4 (Energy deposited in Det. 2 < 1.8 MeV)

Test Sequence Creation

Step 1: Sample poisson distribution to determine when events occur



$$\lambda = (\text{Sample Interval}) \times (\text{Highest Detector Rate})$$

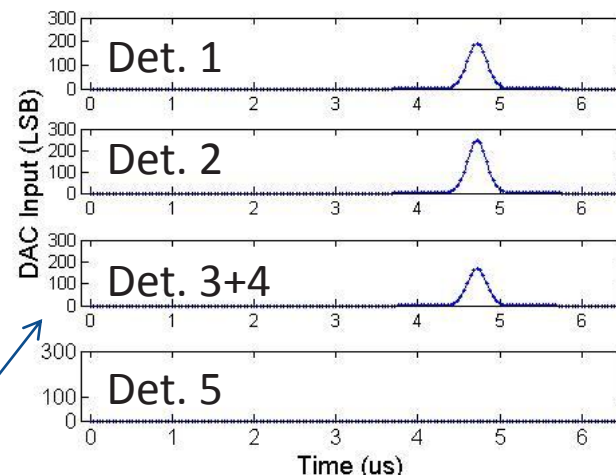
Currently implemented using Matlab functions

Step 2: For each non-zero time, sample the output of Monte Carlo simulation to get peak amplitudes

Event	Det. 1 Energy dep.	Det. 2 Energy dep.	Det. 3 Energy dep.	Det. 4 Energy dep.	Det. 5 Energy dep.
1	0.61	0.00	0.00	0.00	0.00
2	0.17	0.00	0.00	0.00	0.00
3	0.18	0.00	0.00	0.00	0.00
4	0.17	0.00	0.00	0.00	0.00
5	0.16	0.22	0.85	0.00	0.00
6	0.55	0.00	0.00	0.00	0.00
7	0.14	0.00	0.00	0.00	0.00
8	0.15	0.65	1.03	0.00	0.00
9	0.14	0.00	0.00	0.00	0.00
10	0.43	0.67	0.46	0.00	0.00
11	0.35	0.72	0.00	0.00	0.00
12	0.17	0.00	0.00	0.00	0.00
13	0.45	0.00	0.00	0.00	0.00
14	0.17	0.00	0.00	0.00	0.00
15	0.24	0.00	0.00	0.00	0.00
16	0.14	0.64	0.00	0.00	0.00
17	0.16	0.54	0.00	0.00	0.00
18	0.00	0.70	0.33	0.00	0.00

Use uniform (or appropriate) distribution to select particular event with replacement (bootstrap method)

Step 3: Create pulses for event and add to existing sequence



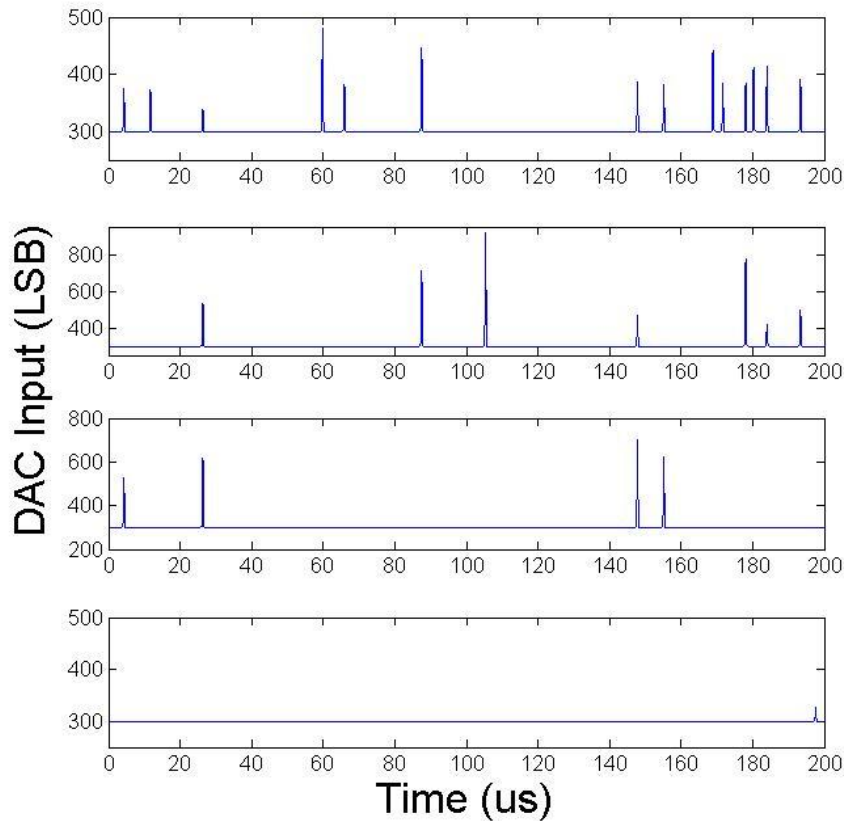
Requires that calibration exists to convert from MeV to DAC input

Process is repeated until all events have been created

Examples of Sequences

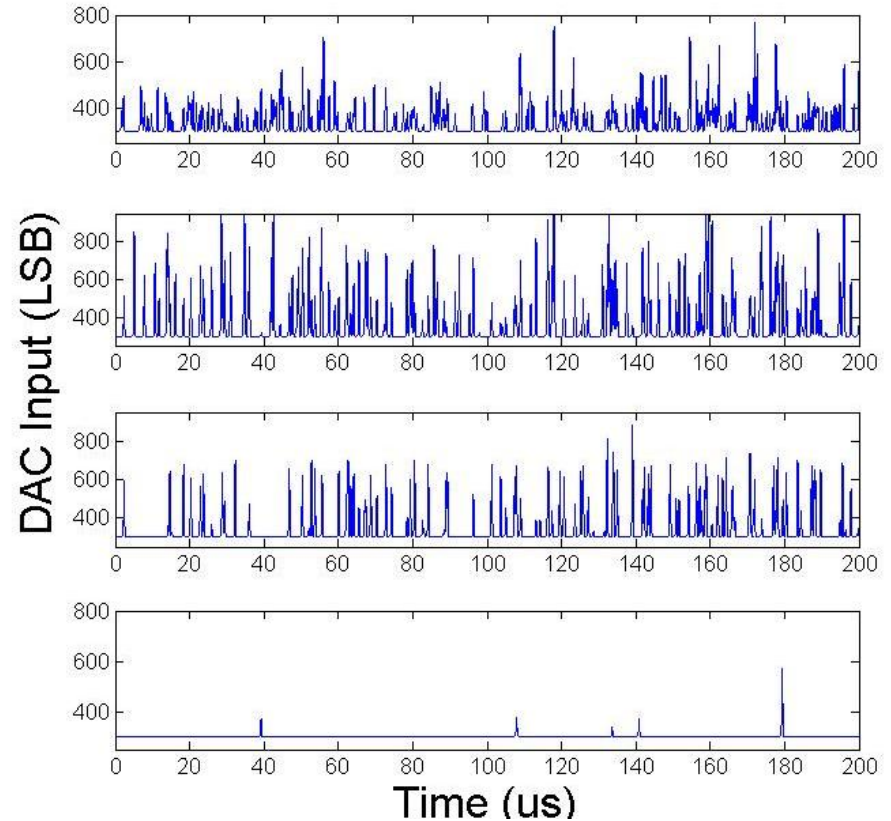
A 3 MeV electron beam was used to illuminate the sensor:

(~ 80 kHz)



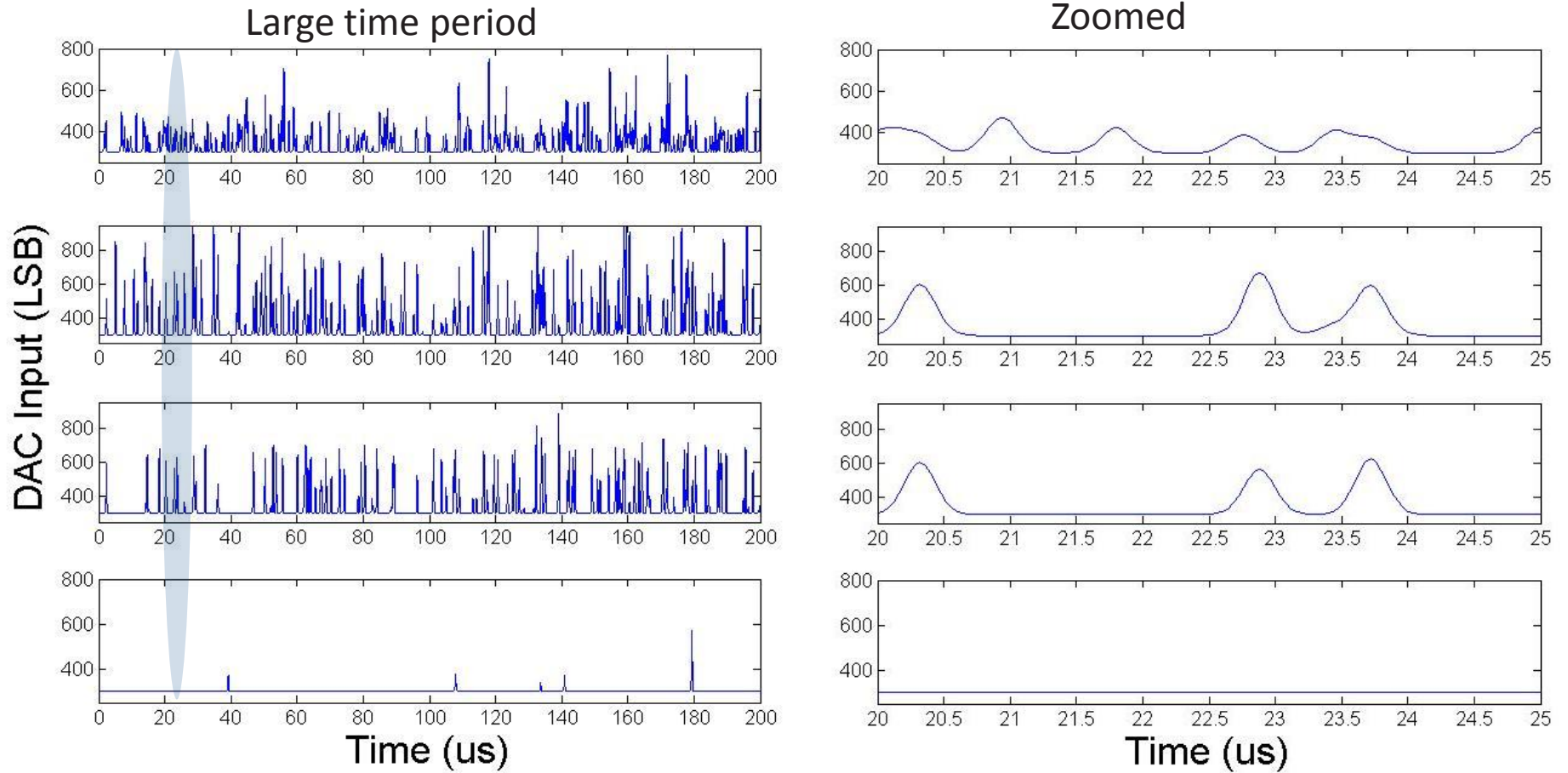
Full length of tested sequence = 0.101 s

(~ 1297 kHz)

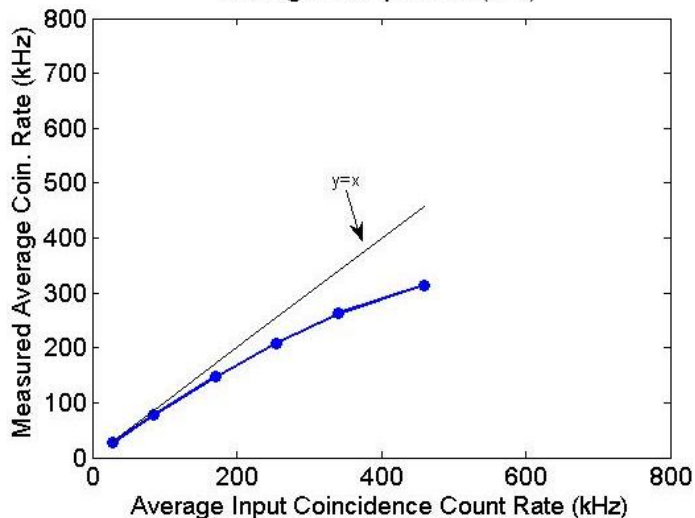
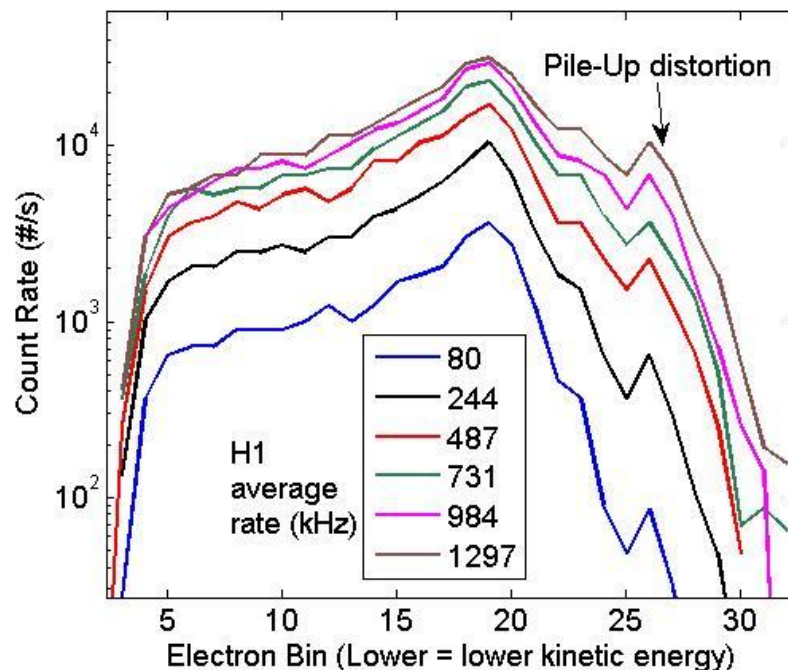
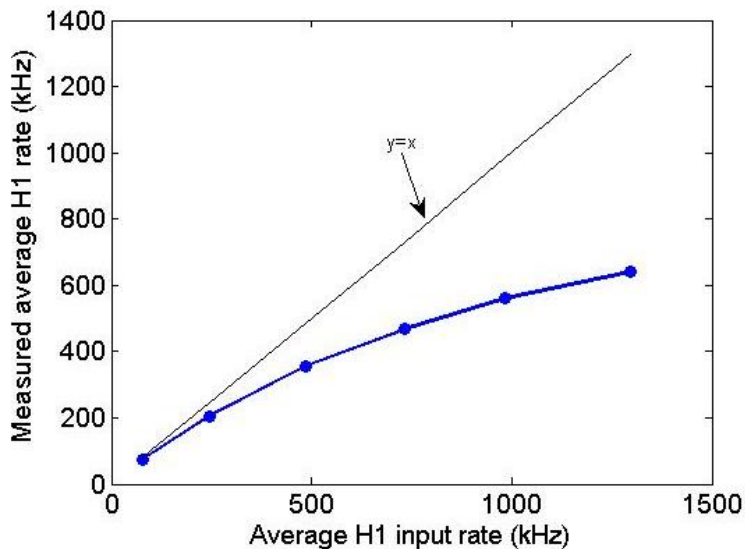


Full length of tested sequence = 0.021 s

Multiple time-scales for 1400 kHz example



Initial Results



- Although approximately 50% of H1 events being dropped at 1.3 MHz more than 70% of coincidence events still being processed
- Note the build-up of the shoulder representative of pile-up in the count spectrum

Conclusions

Methods to investigate misbehavior of the sensor due to anisotropy & high count rates have been proposed

- Often hard to simulate using particle sources on the ground
- Simulating and modeling of anisotropic fluxes on response
- Creating simulated test sequences to investigate dead-time, pile-up, and signal processing artifacts for hardware-in-the loop testing

Questions?