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UCL DEPARTMENT OF SPACE AND CLIMATE PHYSICS
MULLARD SPACE SCIENCE LABORATORY



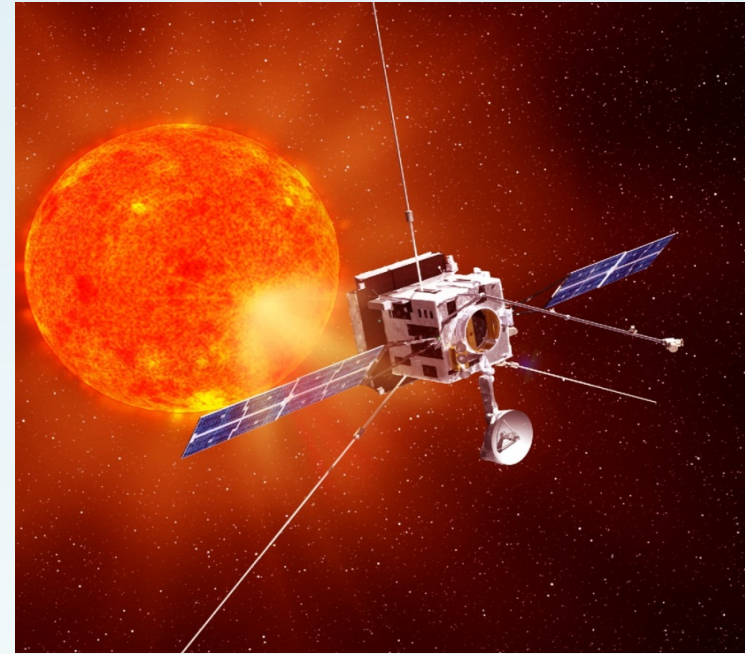
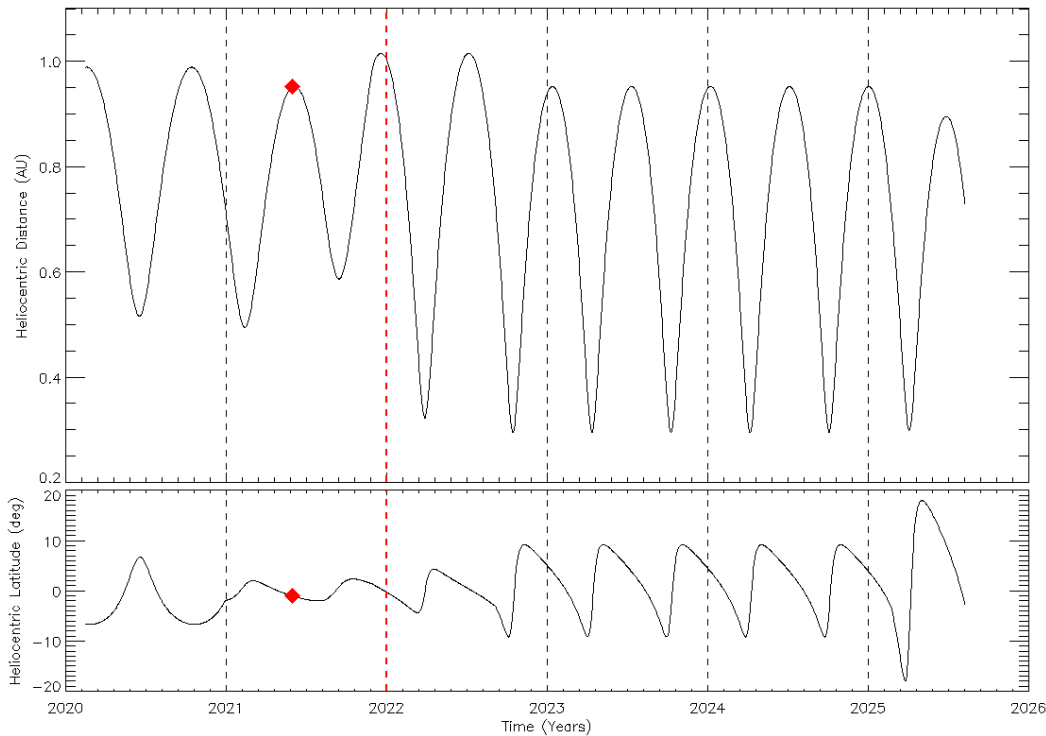
Spacecraft effects on electron measurements - combining SPIS with data from Solar Orbiter's Solar Wind Analyser

G. Lewis¹, R. Wicks², C. Owen¹, D. Kataria¹, S. Reddy¹,
G. Deprez³, D. Verscharen¹, C. Anekallu¹, G. Nicolaou⁴, C. Kelly¹, V.
Fortunata⁵, L. Bercic¹

1. MSSL-UCL, UK
2. University of Northumbria, UK
3. ESA
4. South West Research Institute, San Antonio, Texas, USA
5. Planetek, Bari, Italy

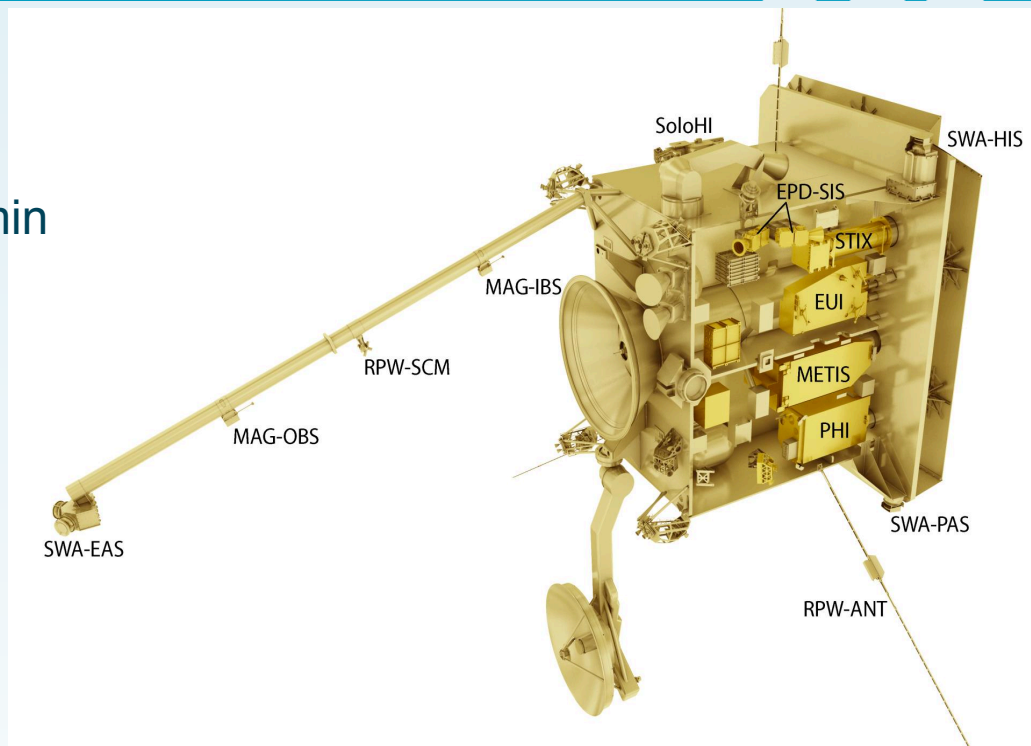
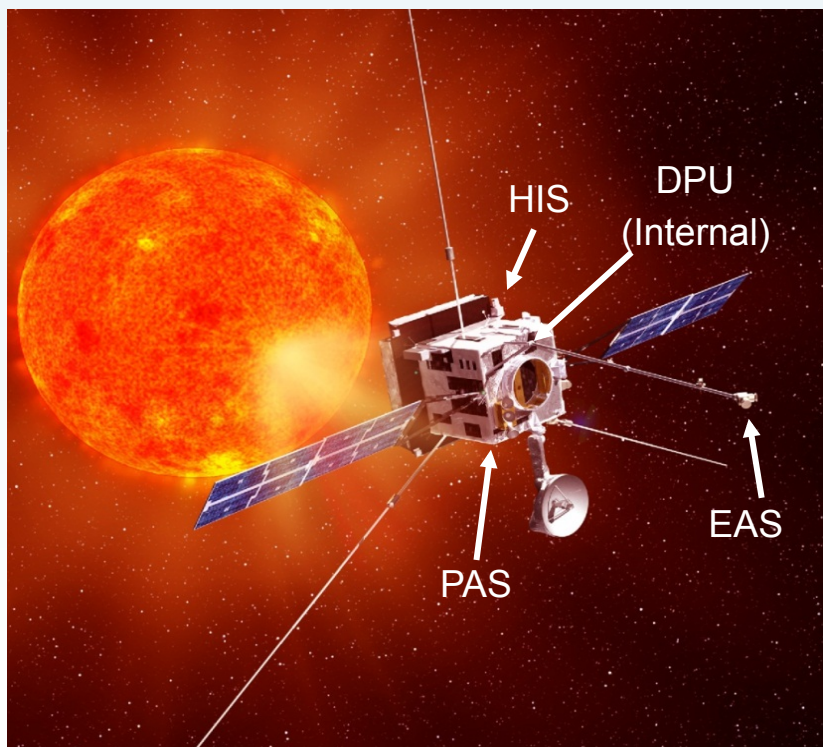
- 1a. Solar Orbiter: Orbits etc
- 1b. Solar Orbiter: Architecture
- 2a. EAS: How does it work
- 2b. EAS: Location, baffle, look-directions
- 3. Measured data
- 4a. SPIS: The model
- 4b. SPIS: Some initial results
- 5. Conclusion

- Solar Orbiter was launched on the 10th February 2020
- It has been in orbit for 16 months
- Has recently completed its second full orbit



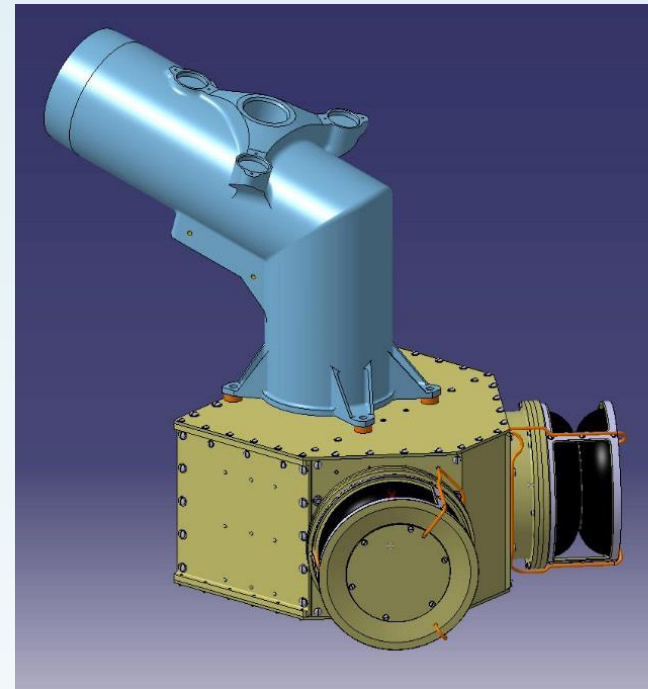
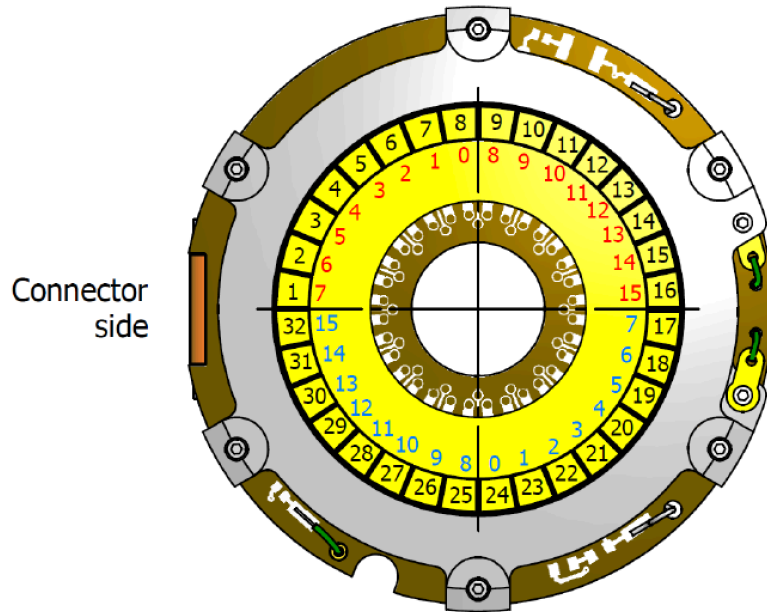
- The SC is still in its Cruise Phase
- The Science Phase begins at the start of 2022

- 10 instruments on board
- *Remote Sensing* instruments will operate around closest approach & min and max latitudes
- *In situ* instruments will operate continuously

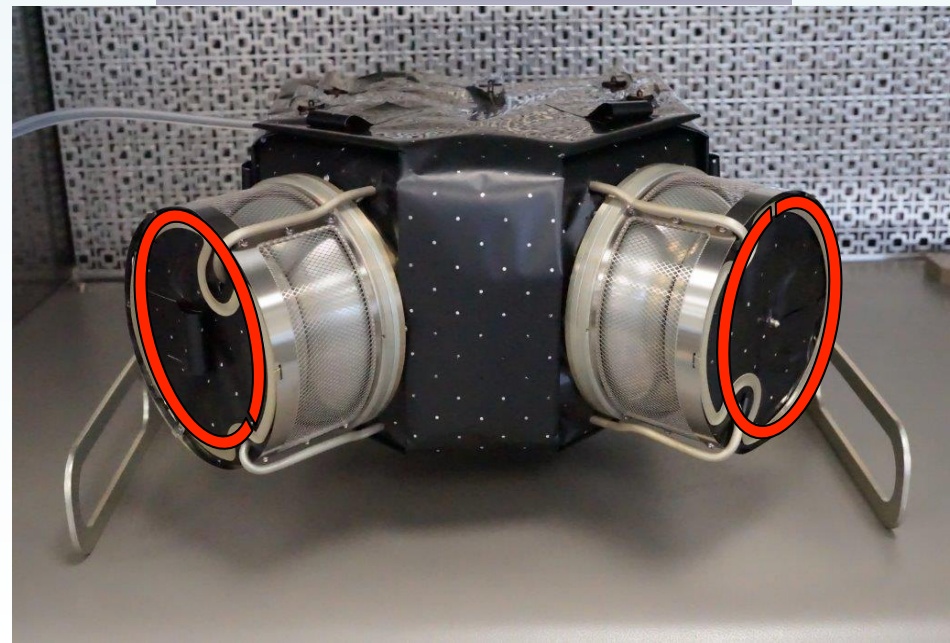
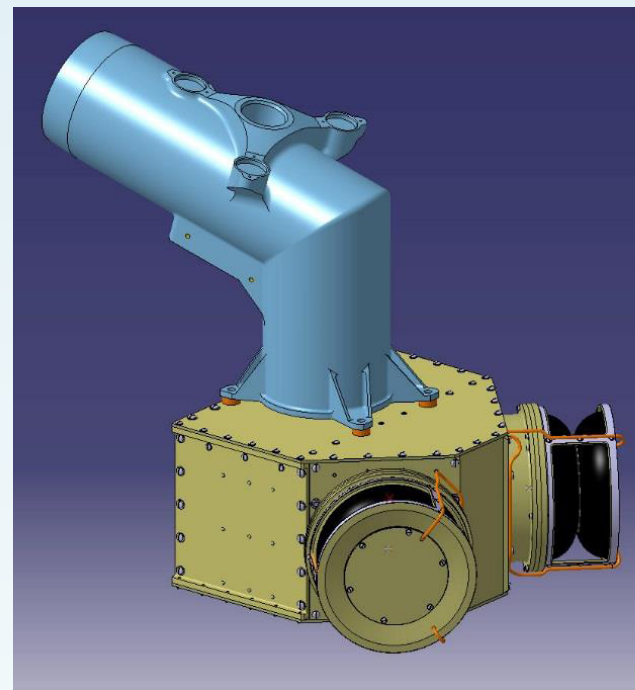
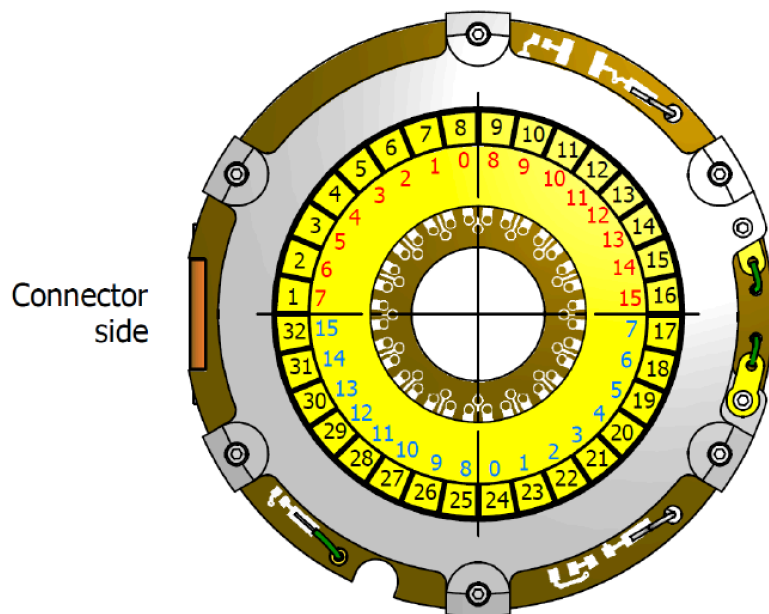


- The **Solar Wind Analyser (SWA)** is made of 3 separate sensors:
- Heavy Ion System (HIS), SwRI & UMich, USA
- Proton Alpha System (PAS), IRAP, France
- **Electron Analyser System (EAS), UCL-MSSL**
- Data Processing Unit DPU, 3 companies, Italy

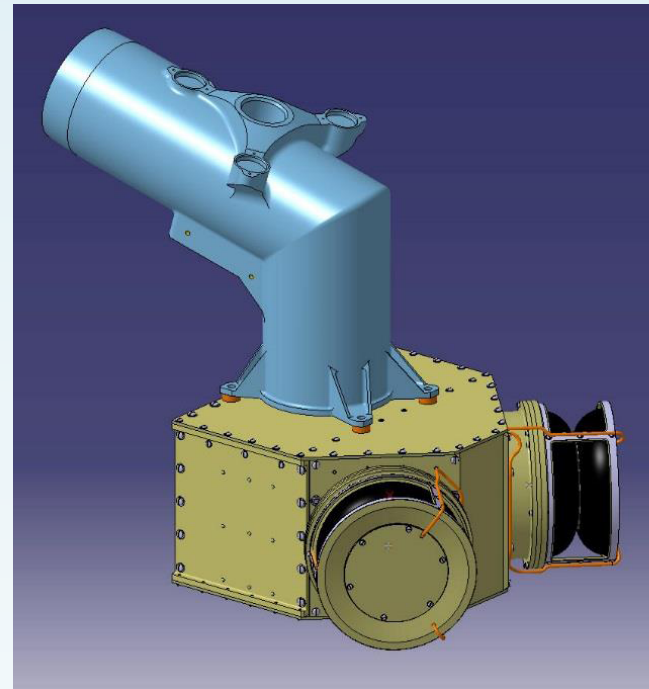
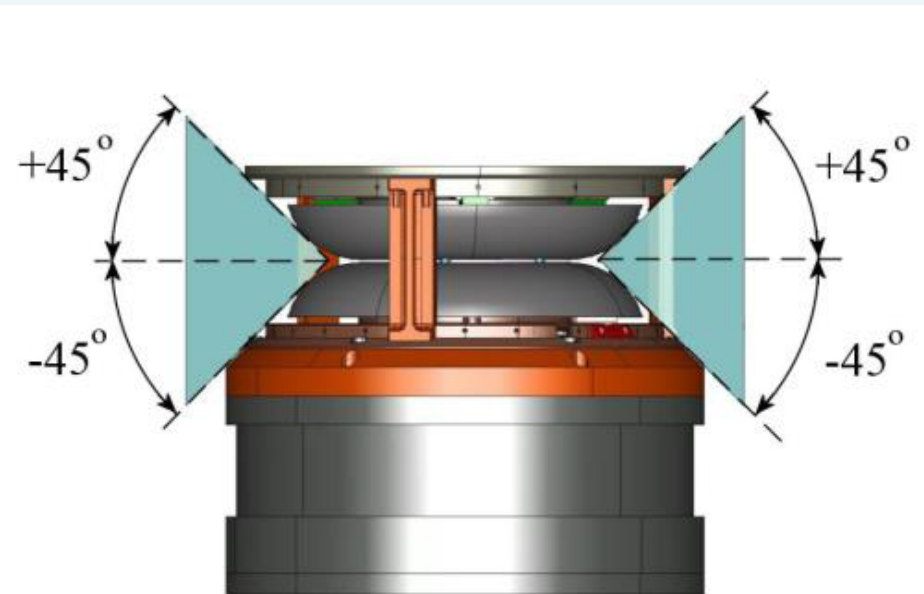
- SWA-EAS is a Top-Hat electrostatic analyser
- Electrons are measured on **32 anodes** that cover 360 degrees



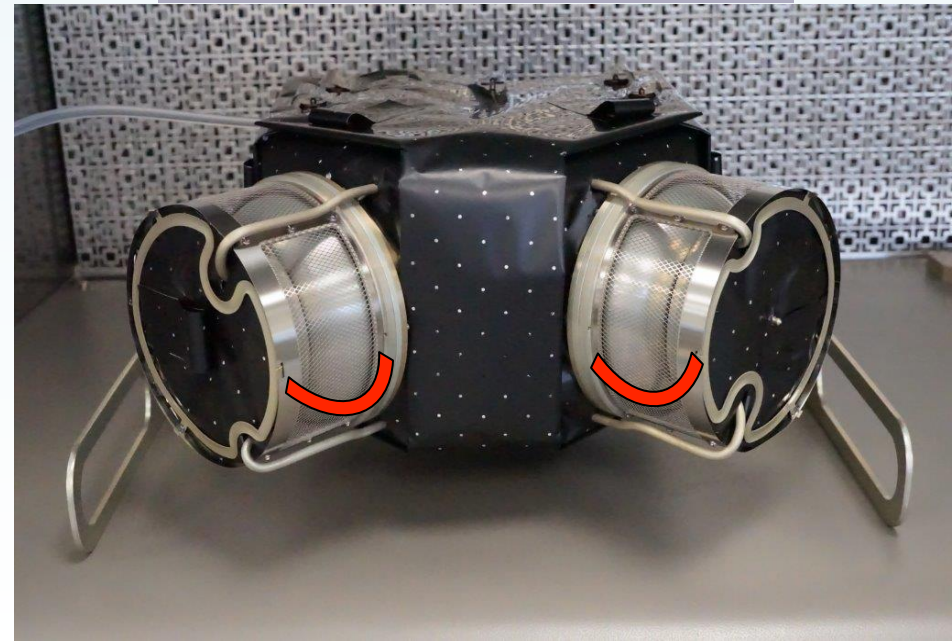
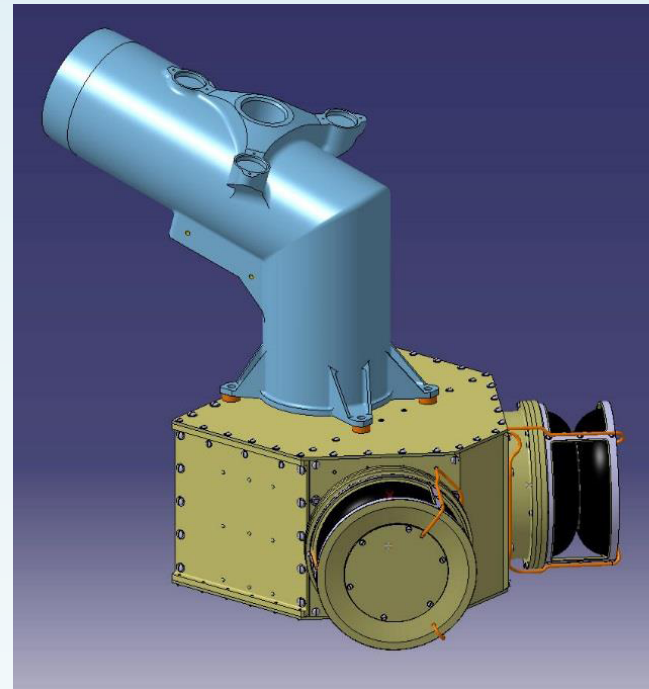
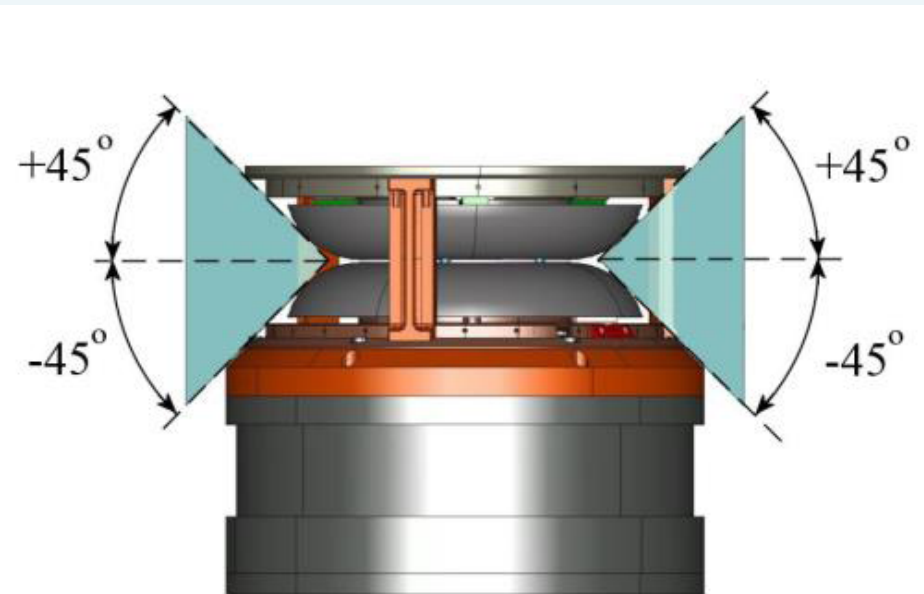
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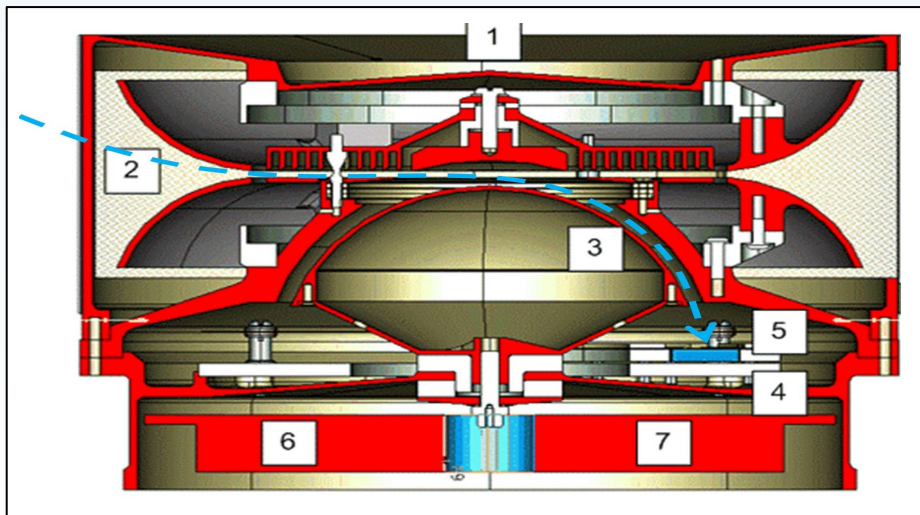
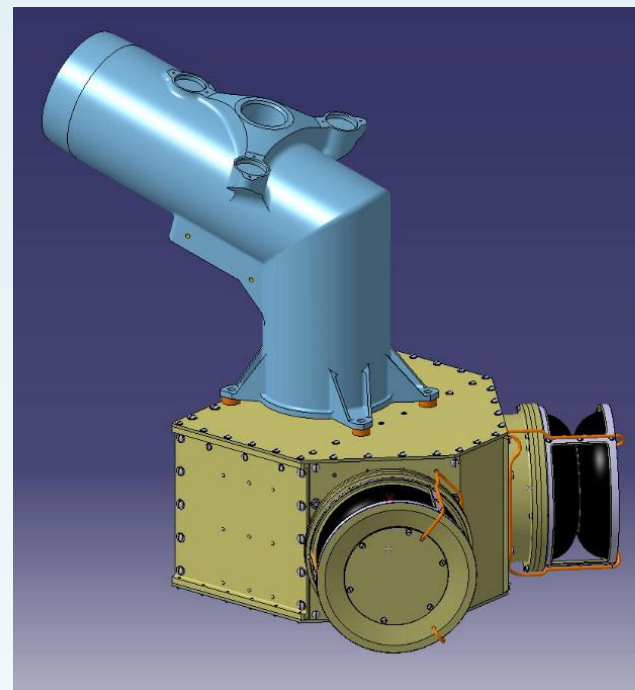
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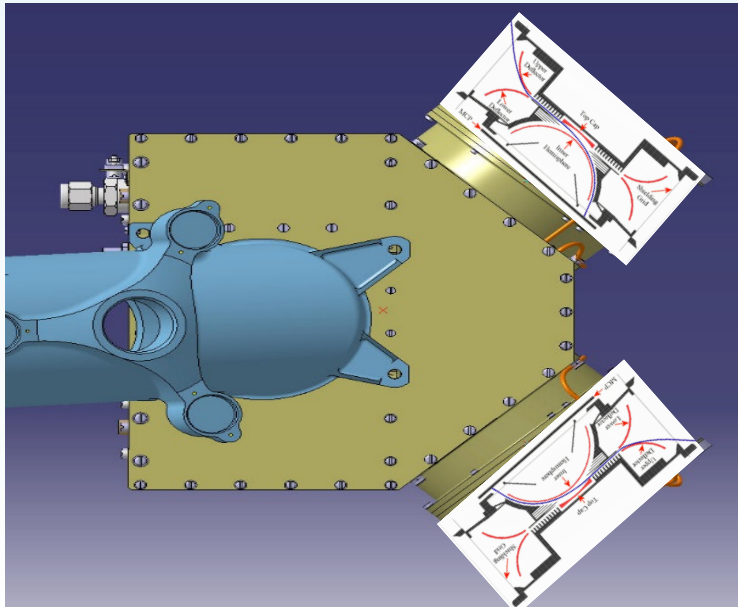
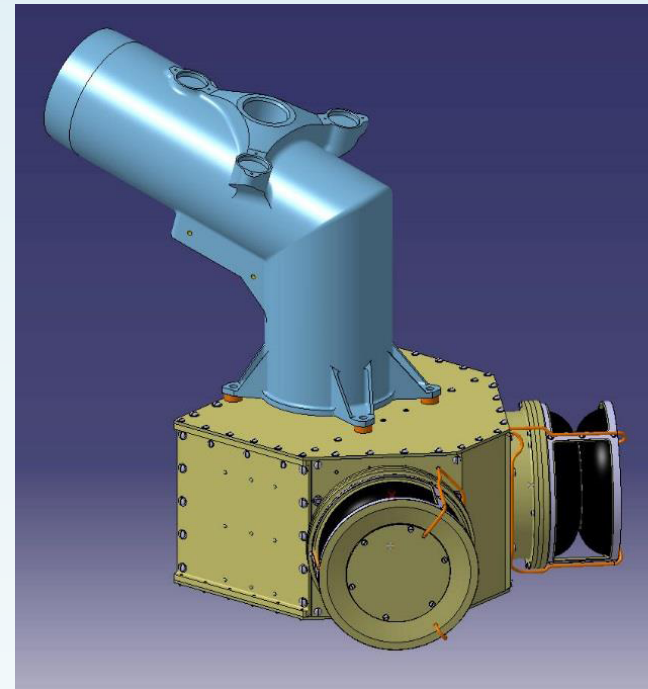
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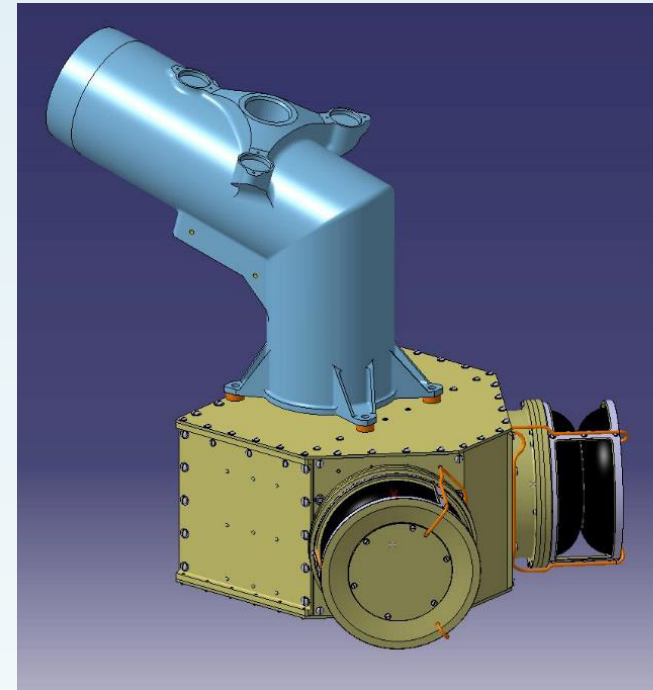
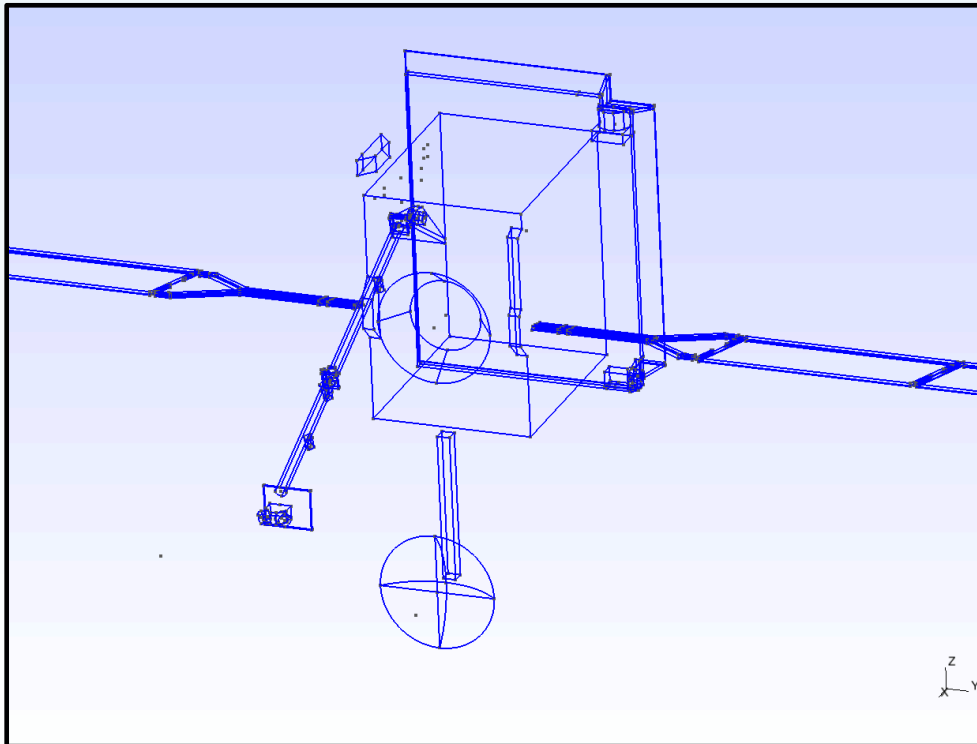
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- Electrons are directed into **16 Elevations** that cover a 90 degree spread
- Electrons are measured over a spectrum of **63 energies** from 0.5eV to 10KeV



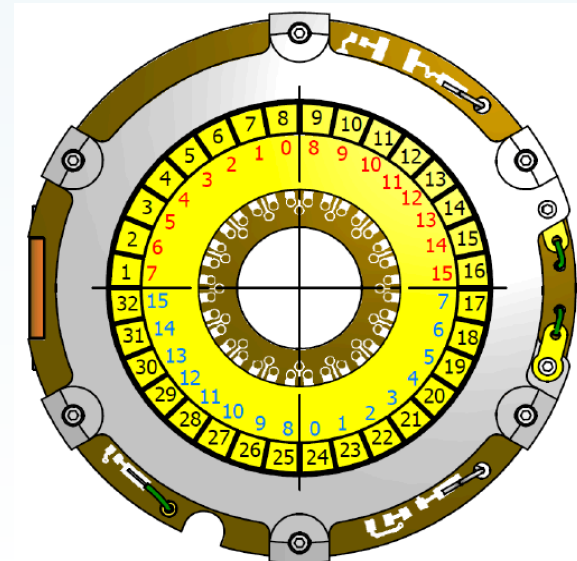
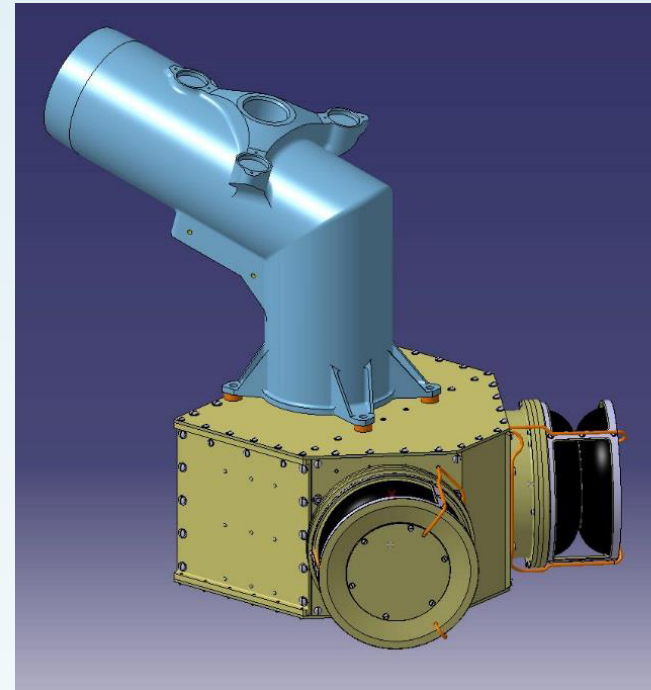
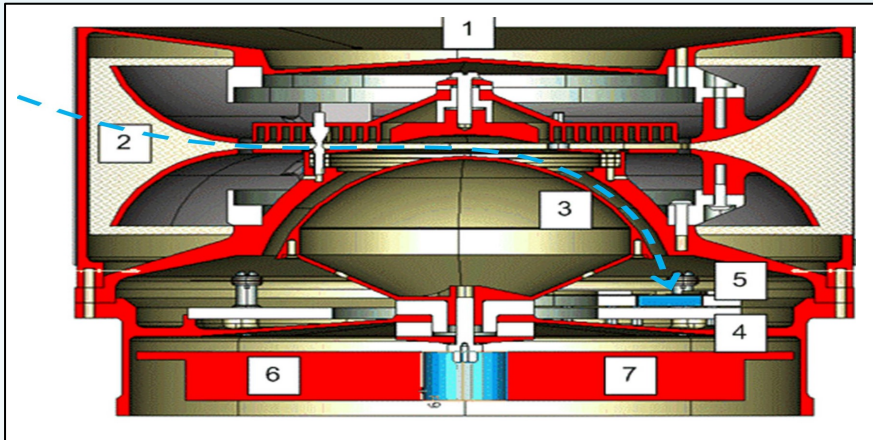
- In **1 second** each EAS covers the 63 energy levels, & the 16 deflectors & the 32 anodes
- The 2 EAS 'Heads' are offset from each other by **90 degrees**
- This means that in 1 second the 2 EAS heads measure electrons over the entire **4π space** around the EAS instrument



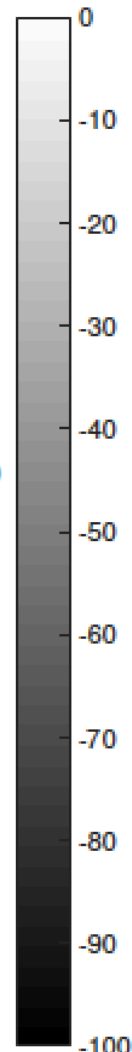
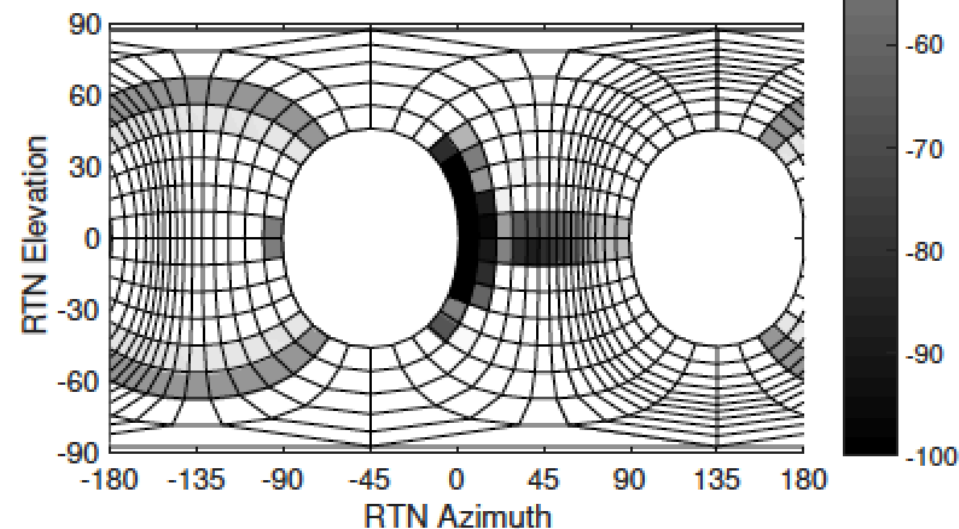
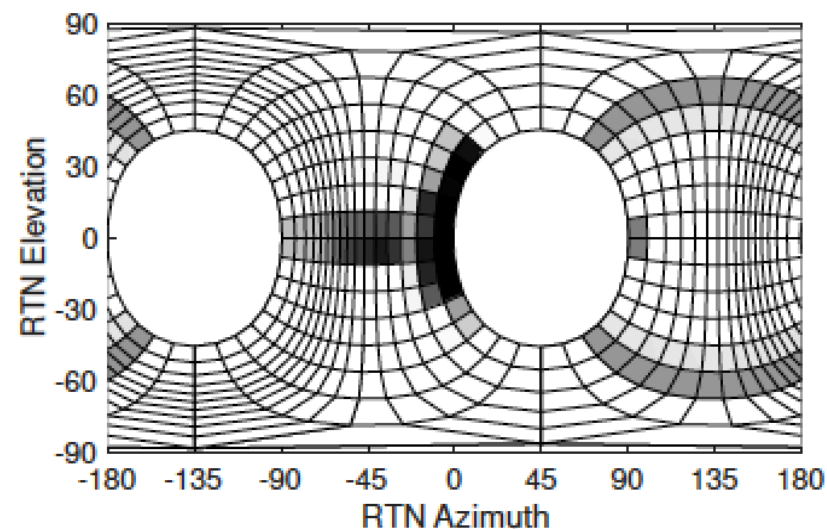
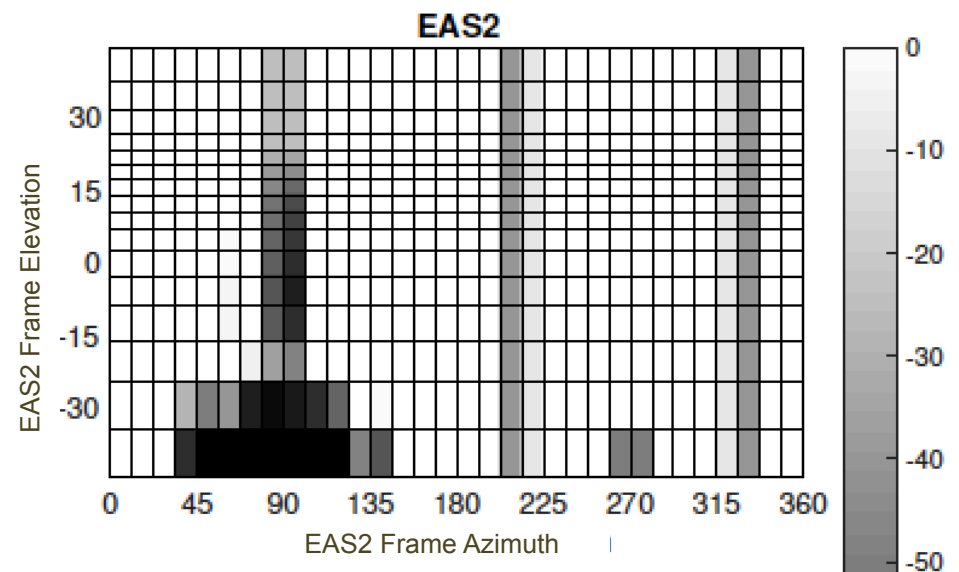
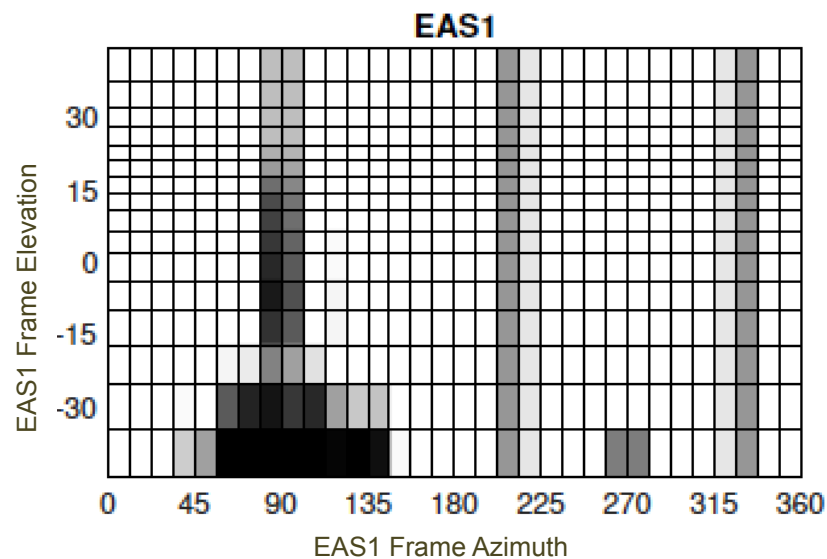
- SWA-EAS has a **4π FOV**
- The SC has many features that can interfere with EAS FoV
- Solar Arrays, SC body, Baffle plate, HGA etc



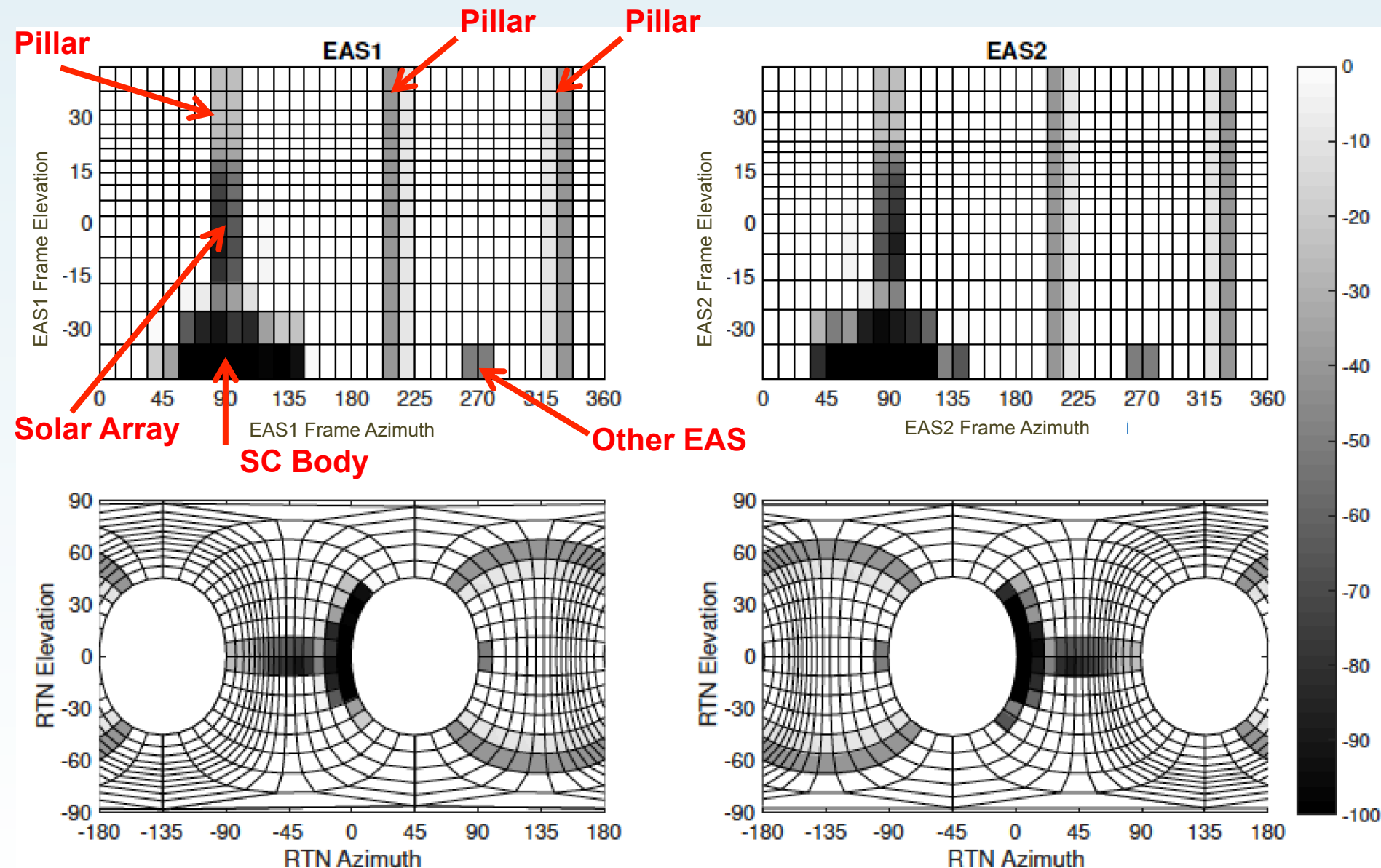
- SWA-EAS has a **4π FOV**
- The instrument also has features that can interfere with EAS FoV



- Percentage of pixel blockage in sensor frame & in a RTN frame

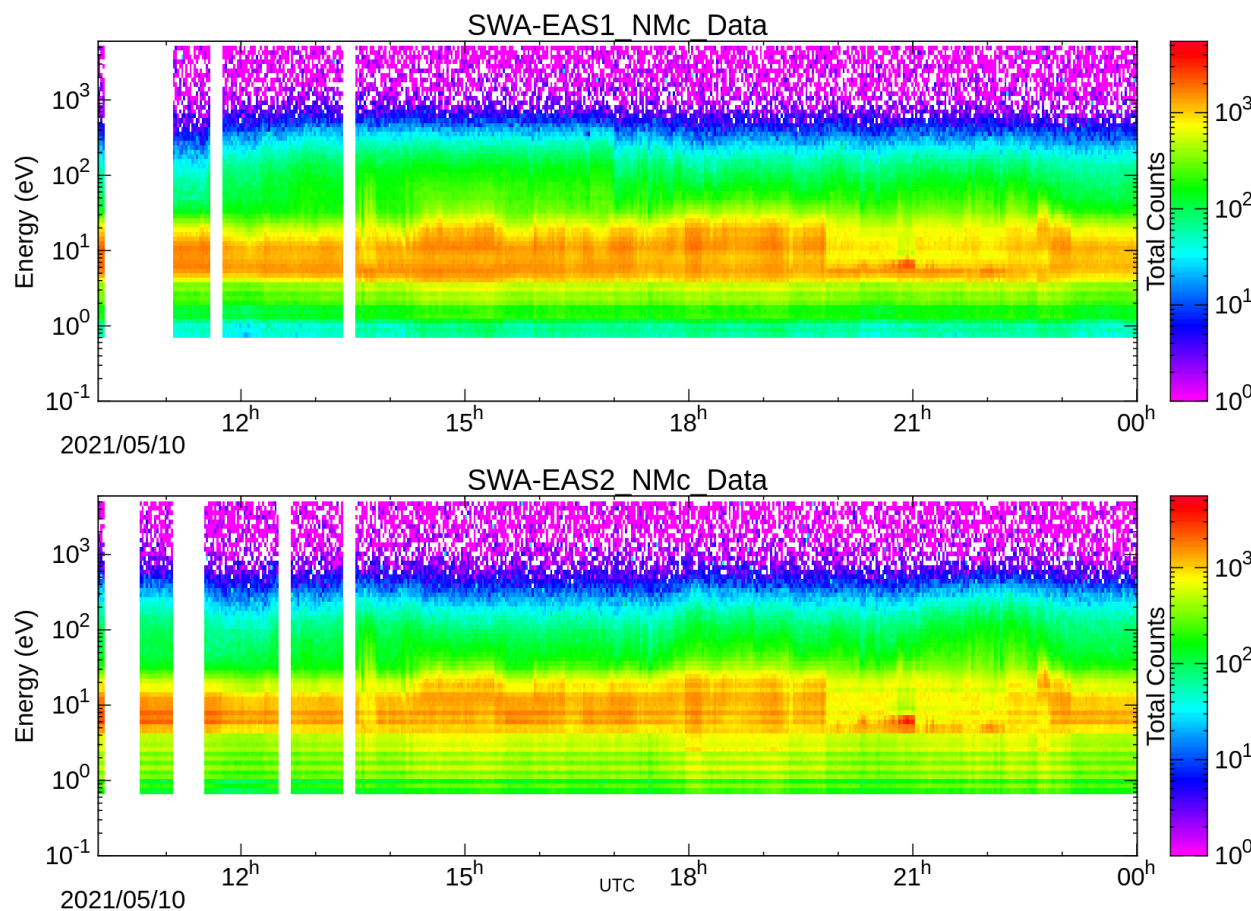


- Percentage of pixel blockage in sensor frame & in a RTN frame





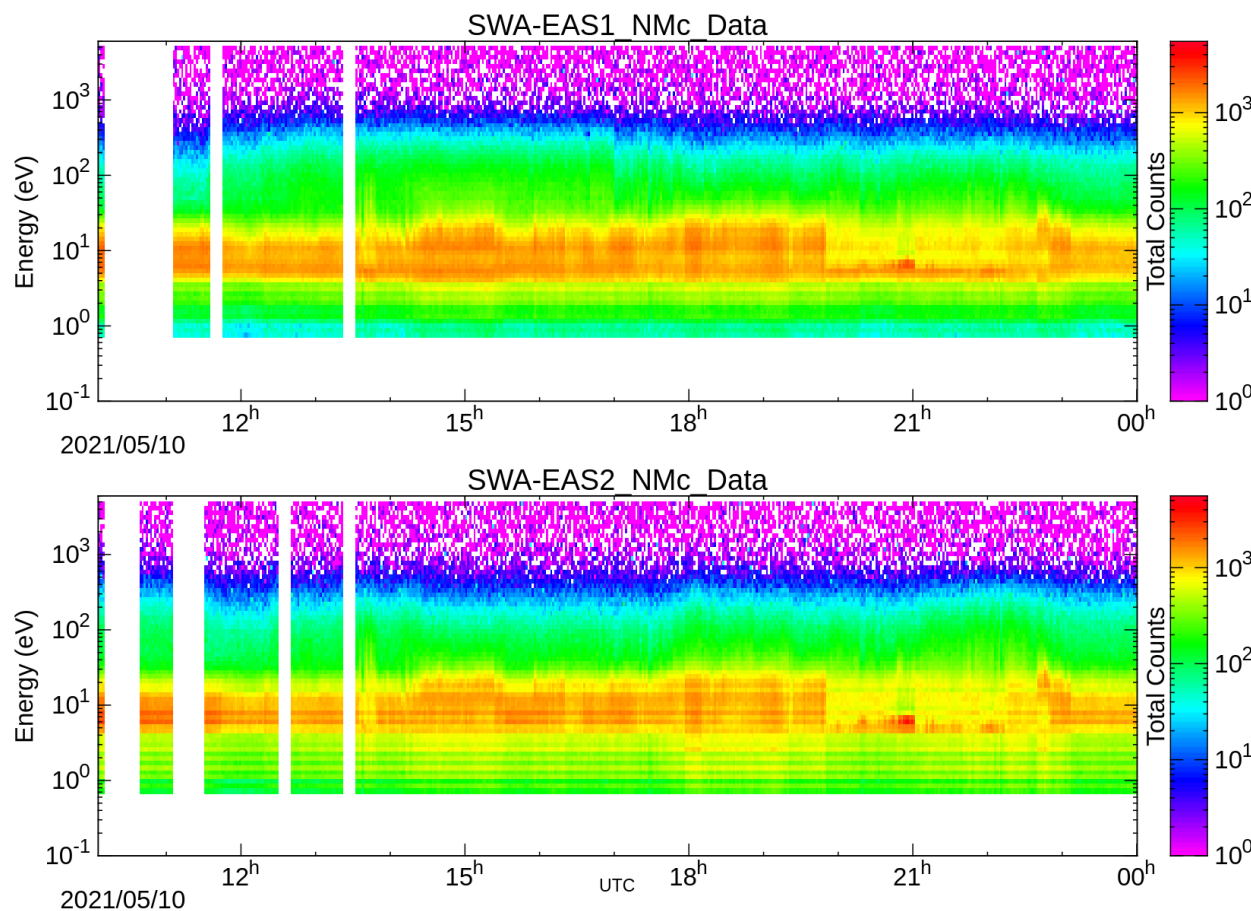
- The EAS data is a [time, 63, 16, 32] array per day
- It is usually displayed as a time, energy spectrogram for each 32 anodes and 16 deflector
- Or with averaging or summing over the anodes & deflectors



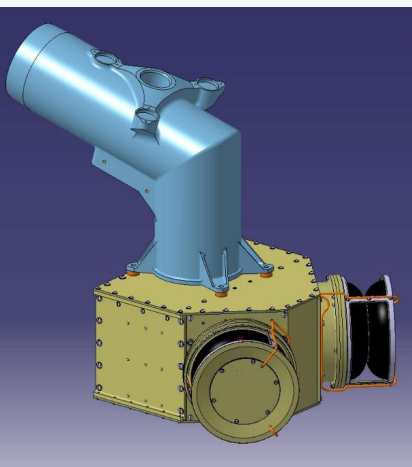
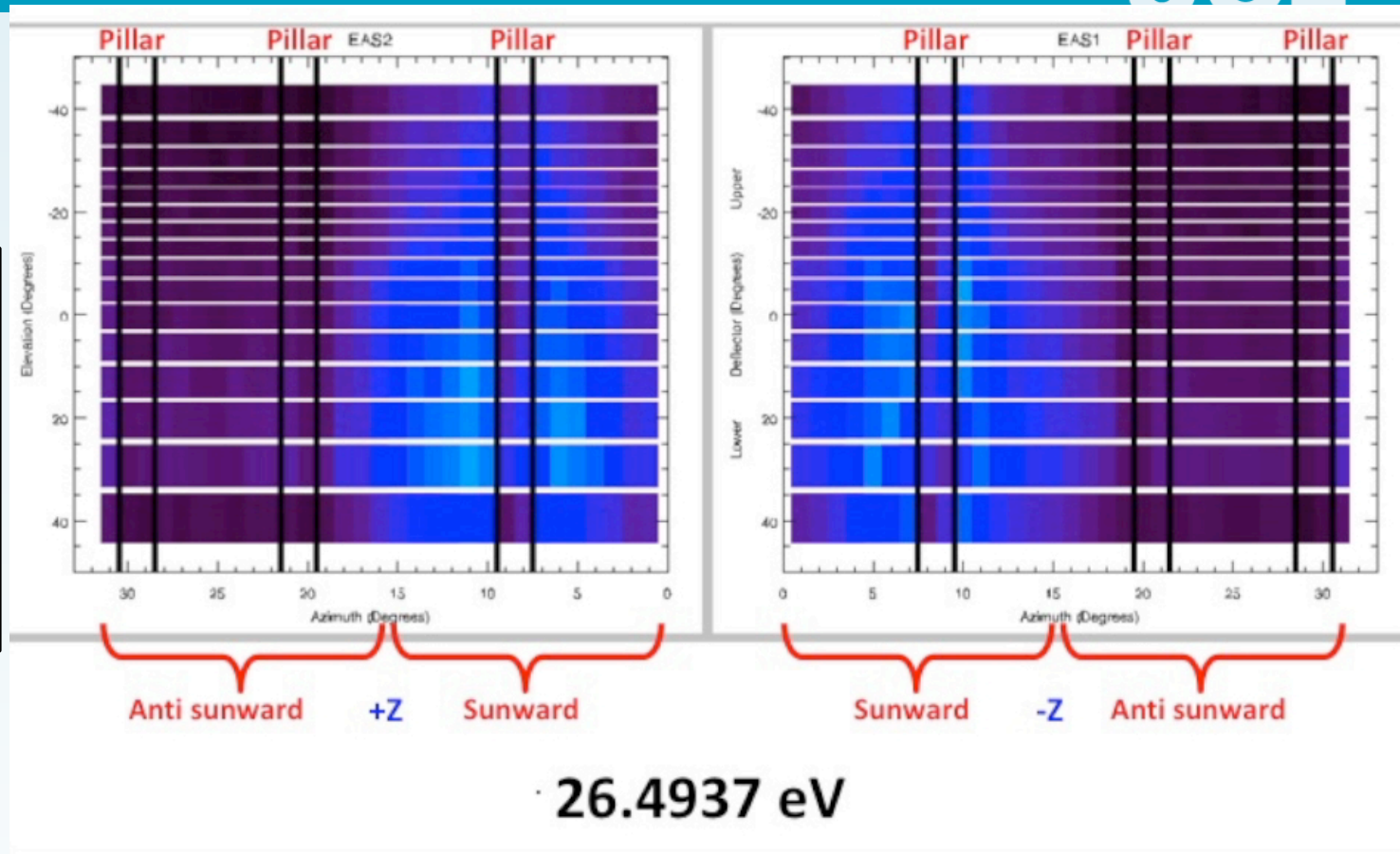
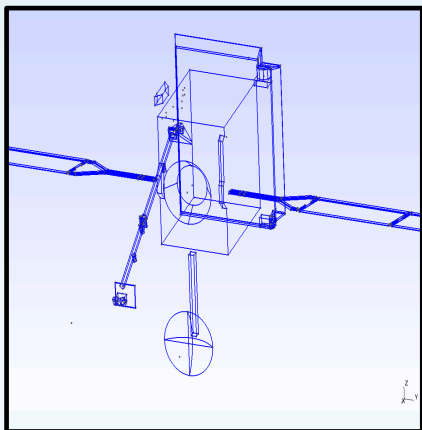
- Notice the distribution peaks at ~10 eV



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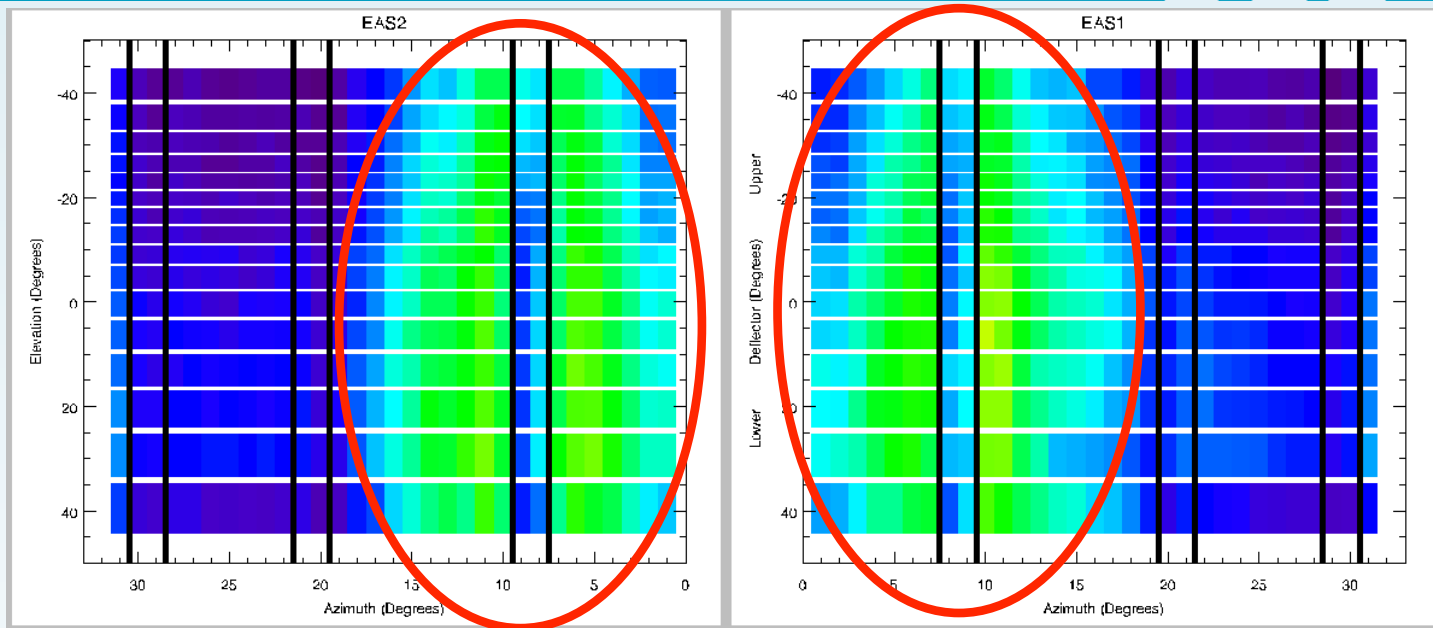


- Notice the distribution peaks at ~ 10 eV
- In the following plot (movie??) I have summed over the whole time array rather than the anodes and deflectors

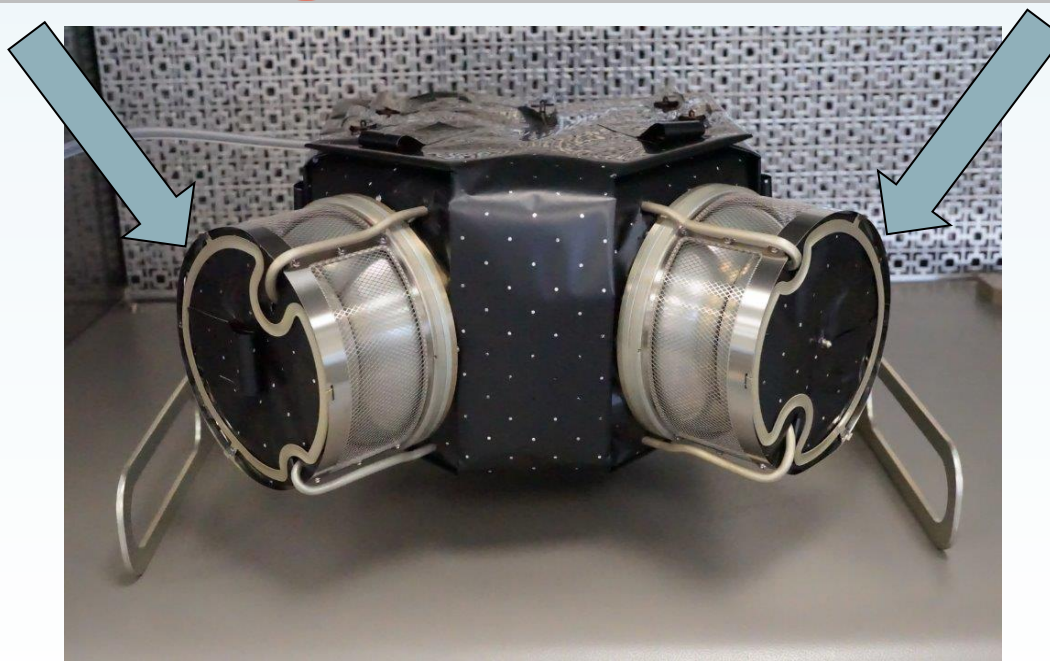


- Lots of direction & structure in the measured plasma
- Not all of it is real Solar Wind
- Lots of SC effects, enhancements, blockages
 - Pillars, Boom/Baffle etc

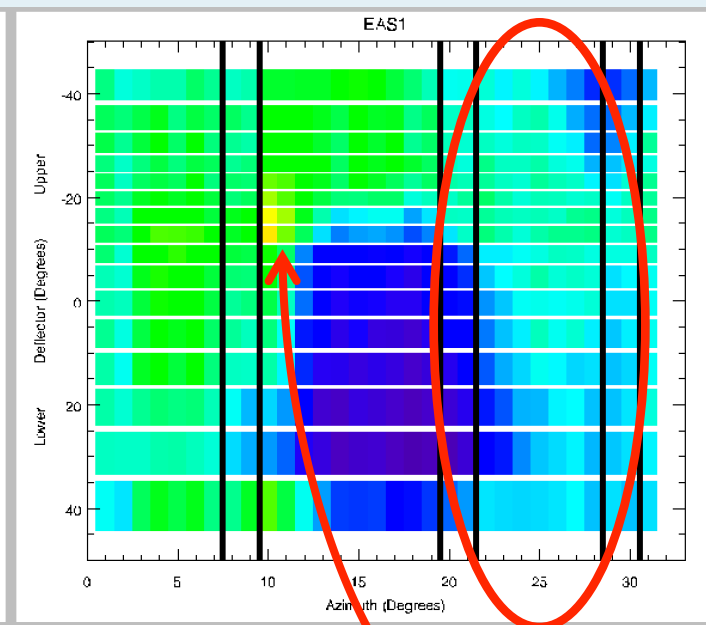
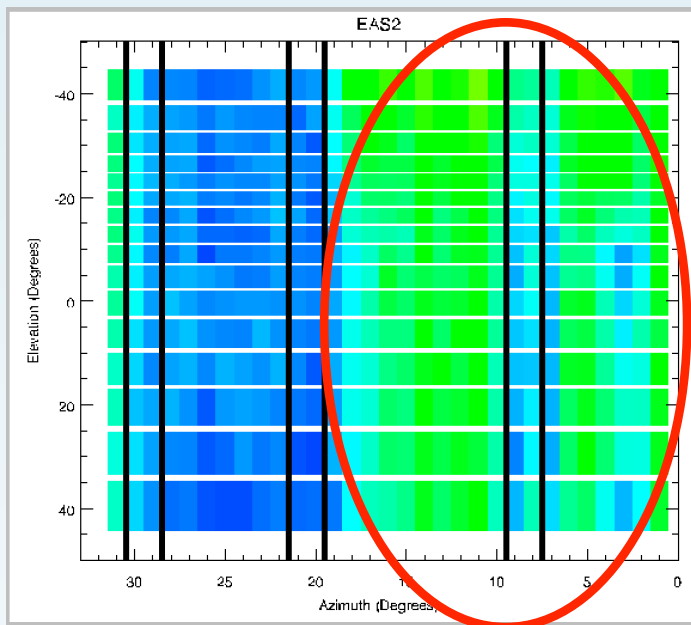
13.1970 eV



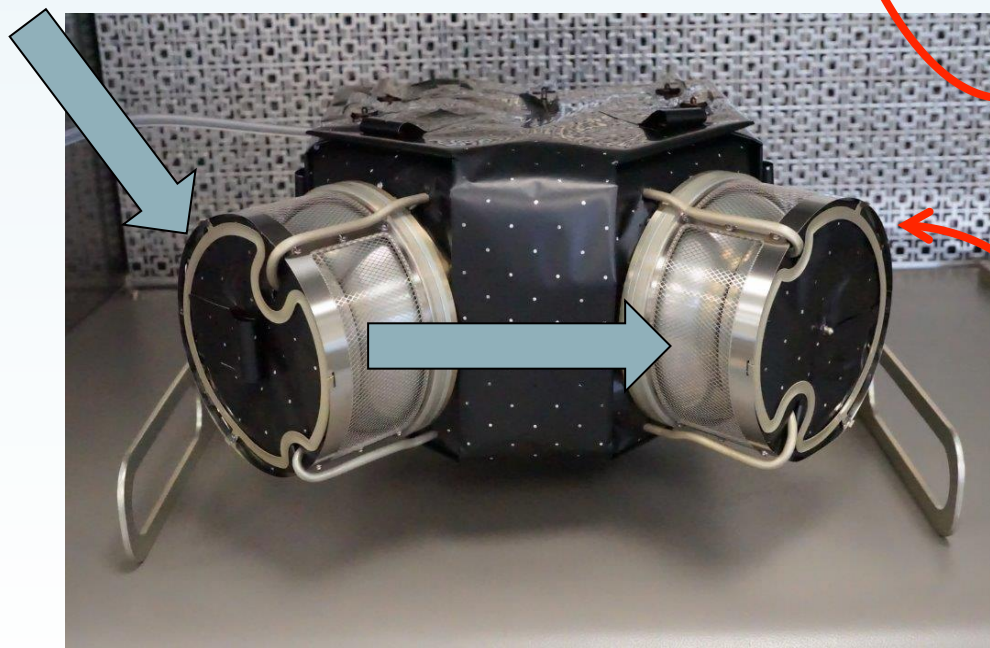
- Solar wind plasma coming around the SC body
- Being detected by both EAS heads



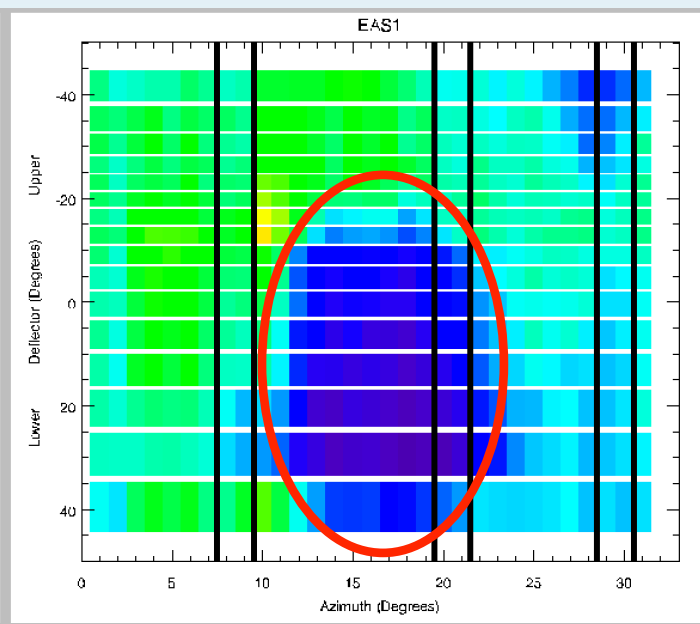
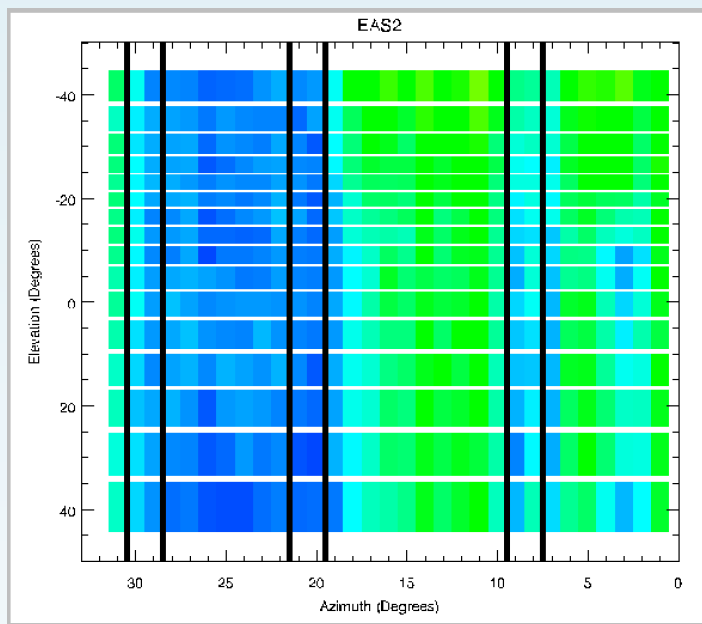
5.7253 eV



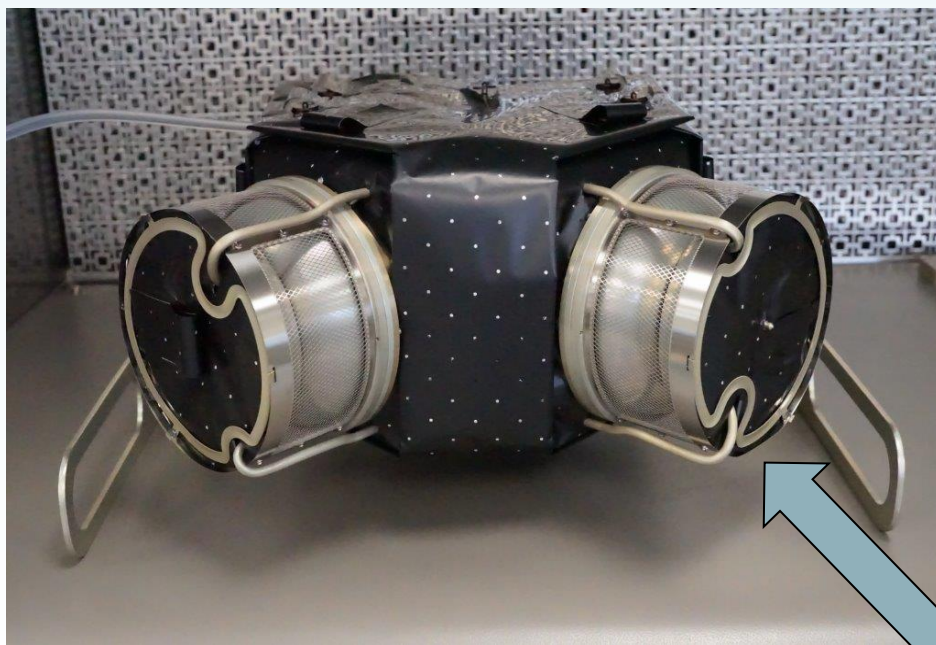
- Photoelectrons measured by EAS2 is continuing on into EAS1
- An enhanced measurement from the Solar array photoelectrons



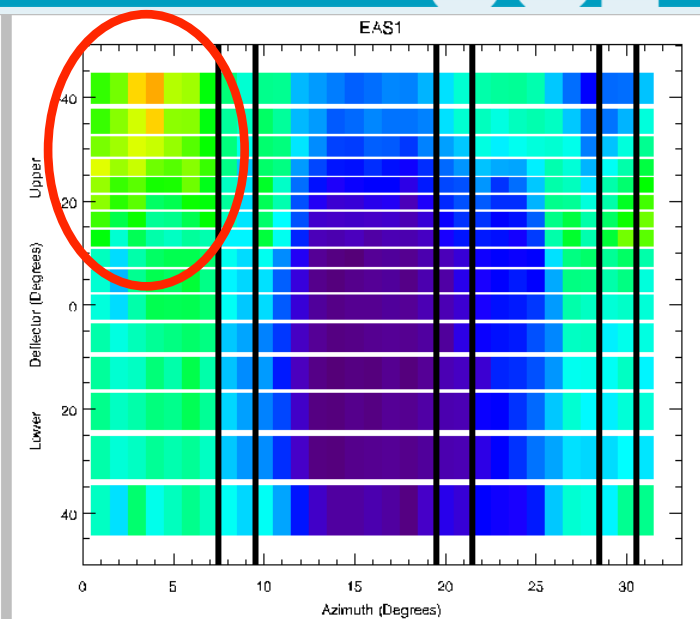
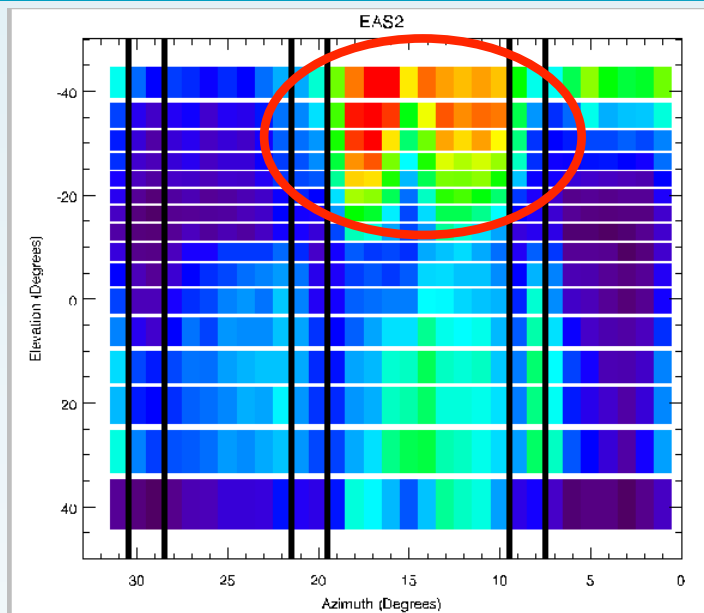
5.7253 eV



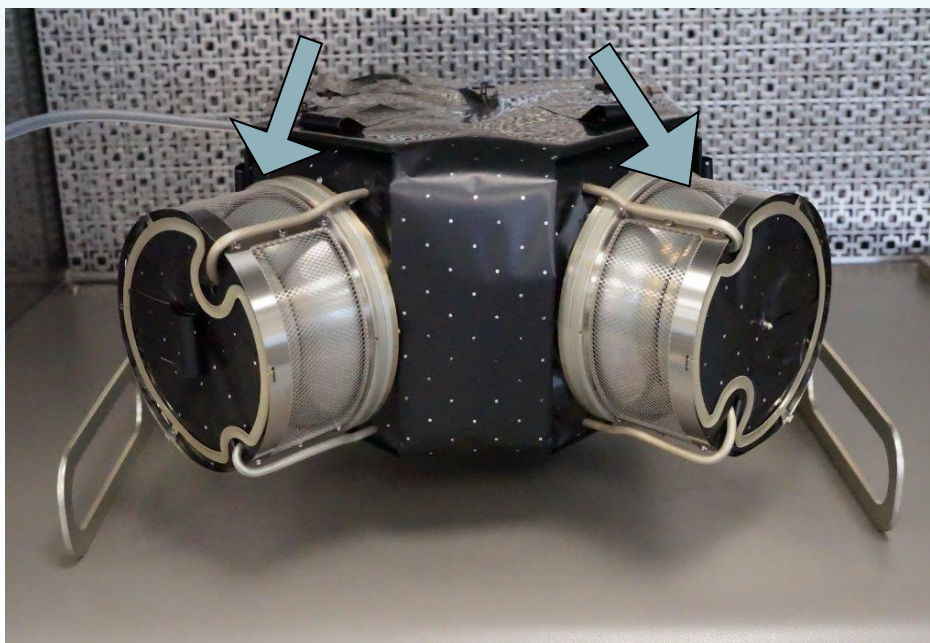
- An interesting depletion on EAS1 is always seen



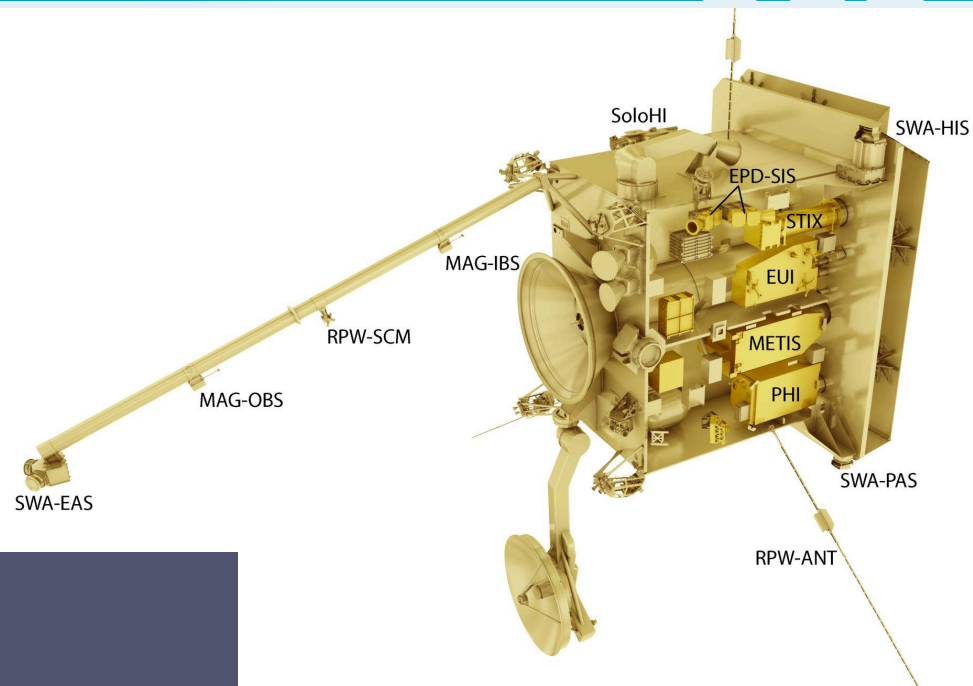
4.3283 eV



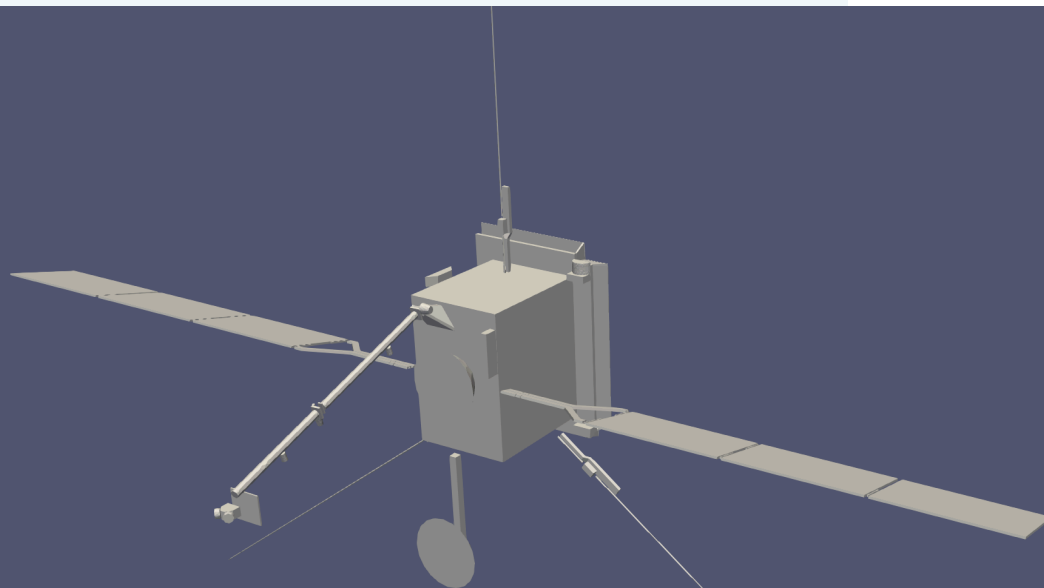
- Photoelectrons measured by EAS2 & EAS1 from the boom or baffle



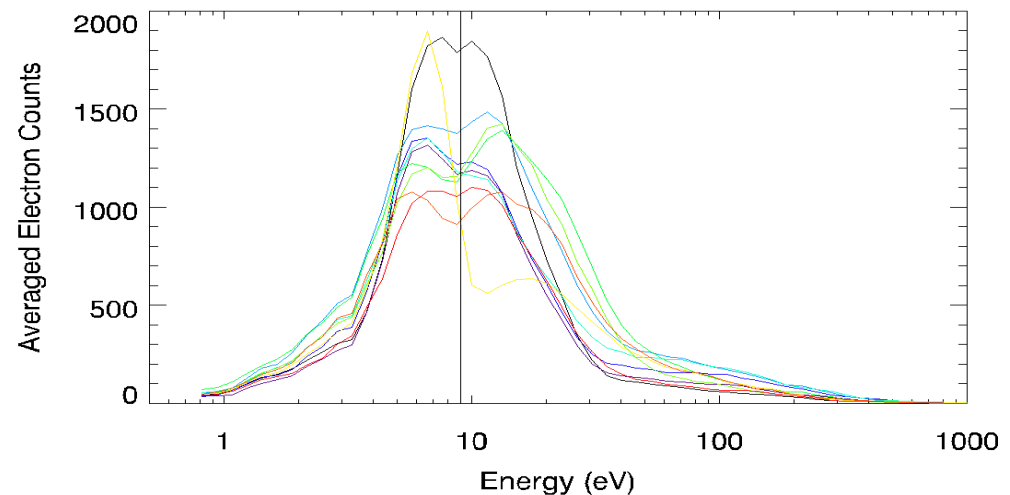
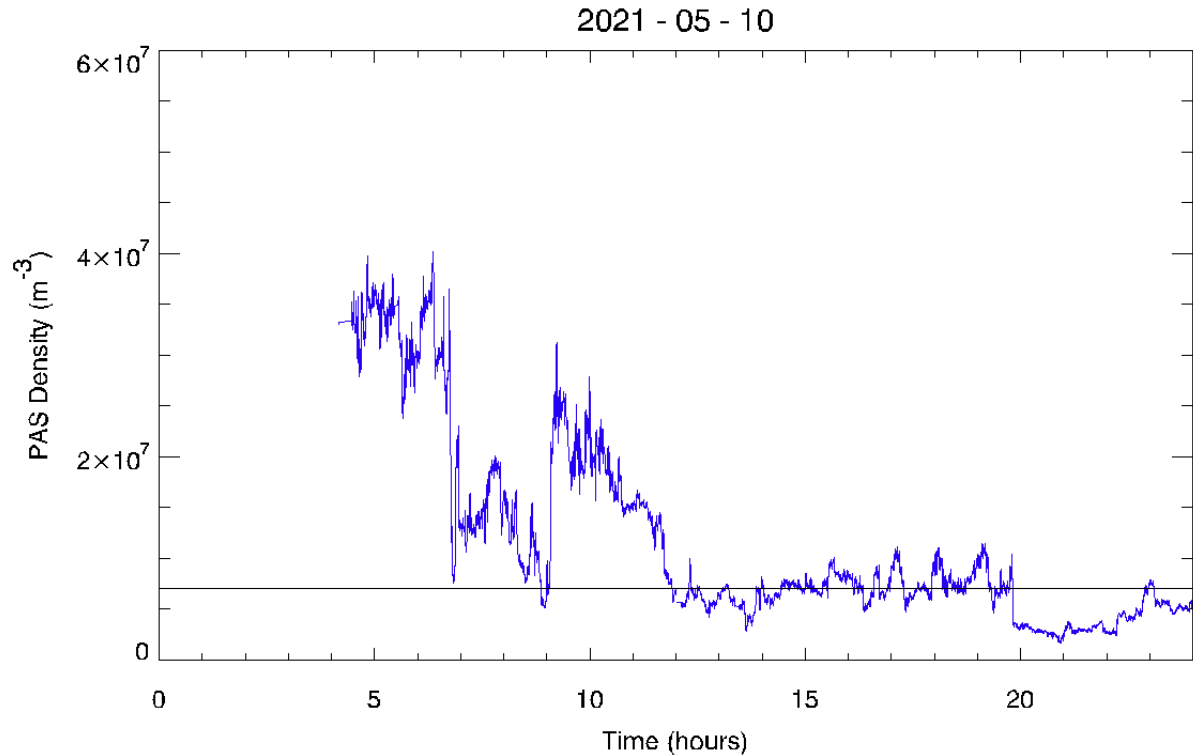
- SO SPIS model provided by ESA
 - Solar Arrays
 - HGA
 - Boom
 - EAS, Baffle
 - RPW
 - Heat shield
 - Radiators



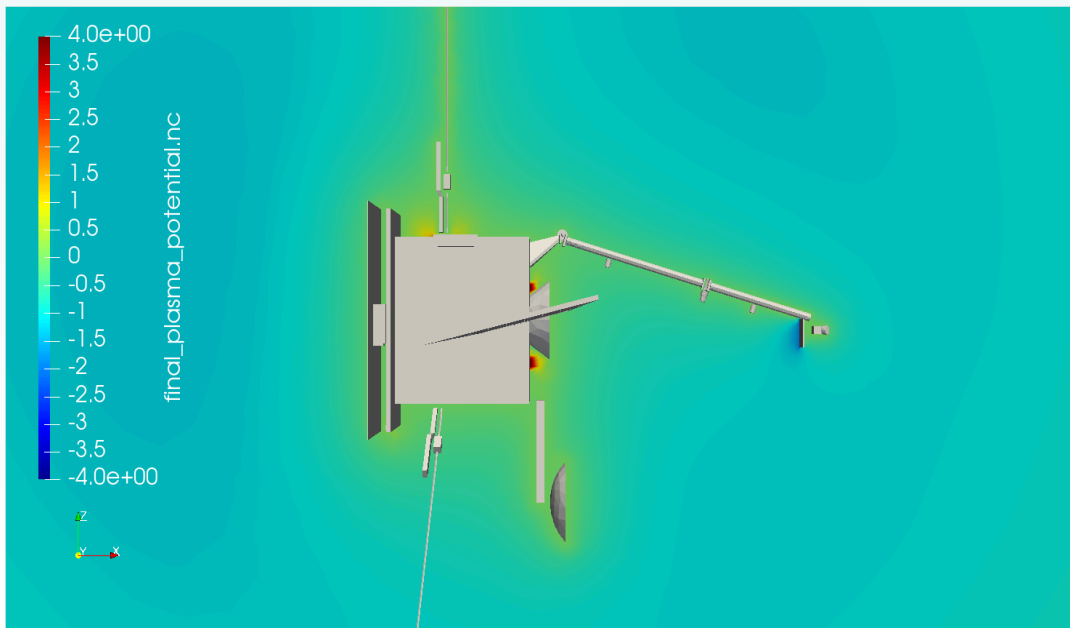
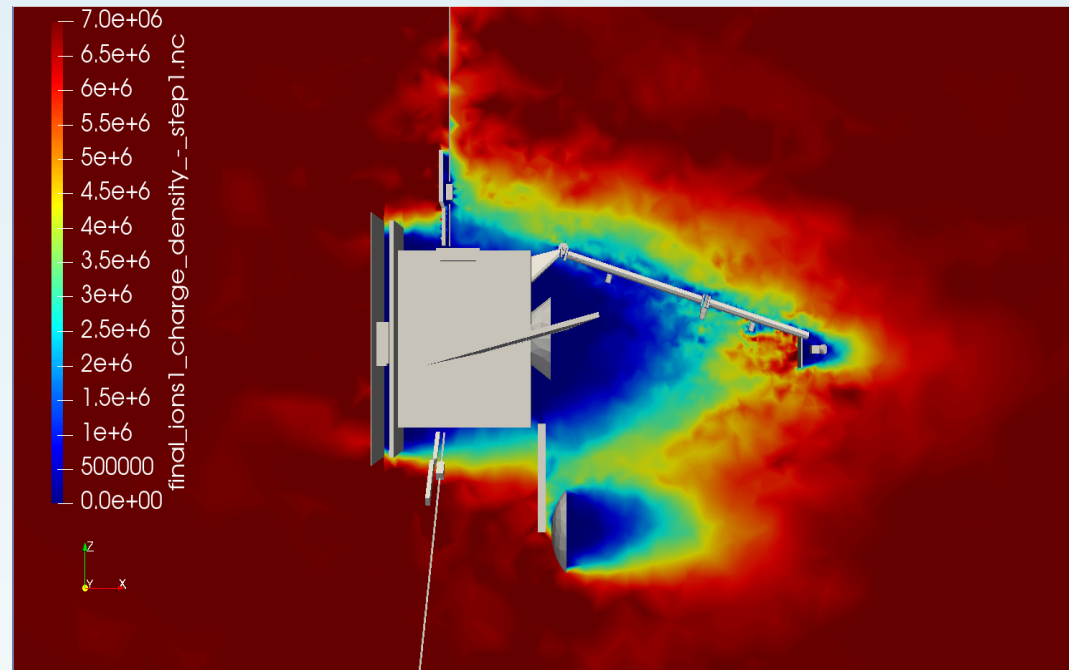
- 43 groups
- Group materials & properties, not precise
- Runs on a 64 thread, 32 core machine
- Run time is about 40 minutes



- SWA data measured on 10th May 2021
- Due to thruster firing, EAS does not operate until 10:00
- PAS 'on-board' ion density measures average $7 \times 10^6 \text{ m}^{-3}$
- On average EAS shows a potential $\sim 9 \text{ eV}$

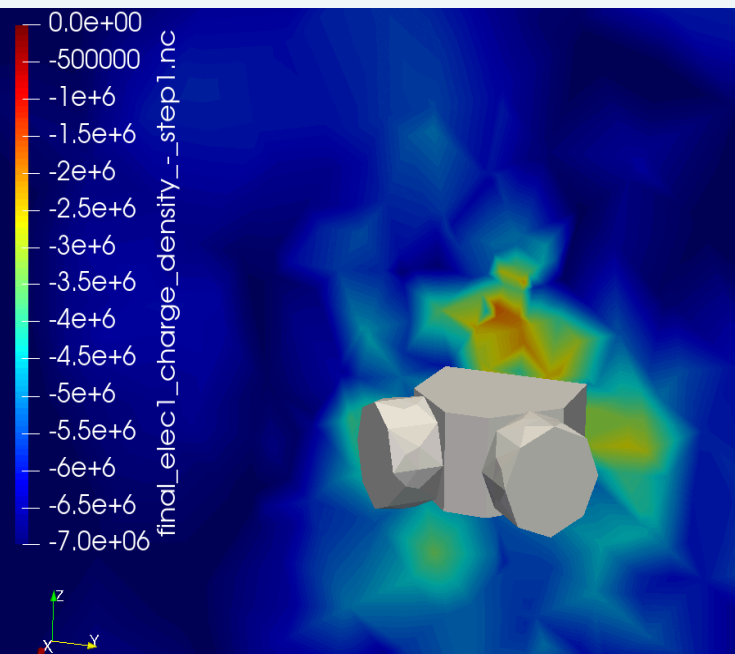
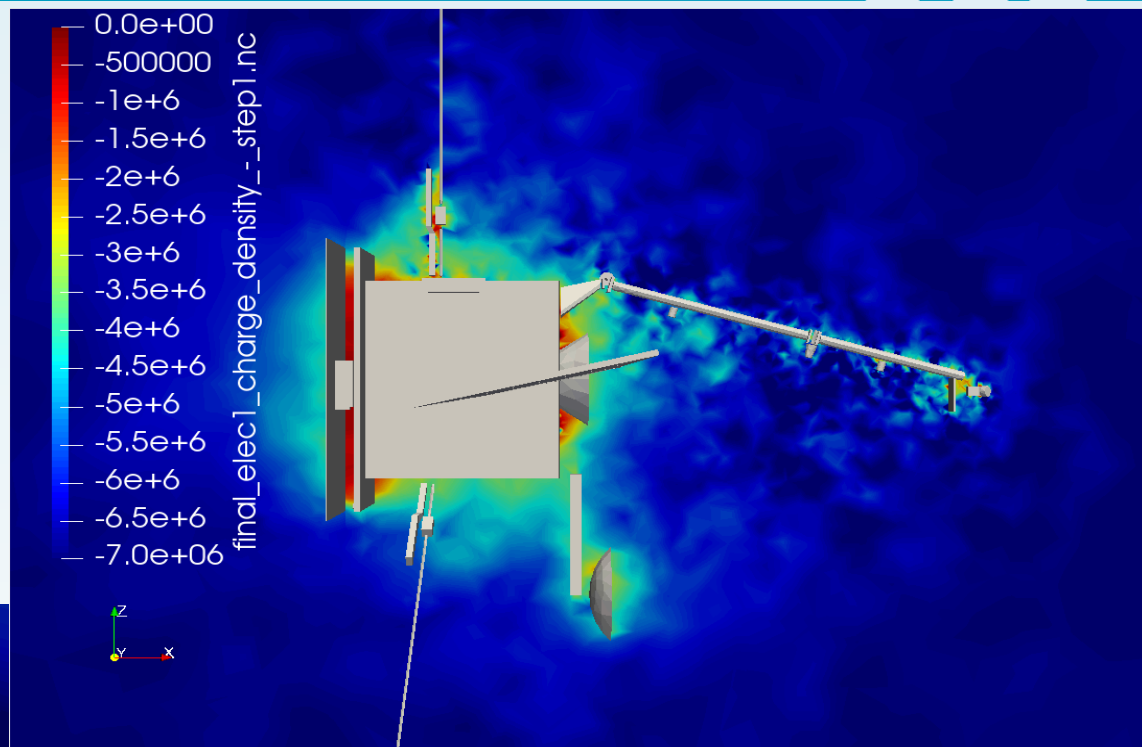


- To compare with data measured on the 10th May.....
- Version 6.0.0
- Ion & Electron input density = $7e6 \text{ m}^{-3}$
- Initial potential = +9 eV
- Sun distance ~ 1AU



- Both ions & electron density is as expected
- Ion density shows a 'nice' wake as expected
- EAS is sitting in an ion void
- The resulting potential seems low compared to the observed values

- There appears to be some structure in electrons around EAS
- Probably due to the boom and/or the baffle



- These results will need to be compared with the measured data
- This is work in progress and the model needs more work

- The Solar Orbiter mission offers an exciting laboratory for Spacecraft charging
 - Thruster firing -> varying boom charging etc
 - Variable solar distance
- SWA-EAS is in prime position
- In the two orbits of the mission so far EAS has already measured ~144 days of data
- Our SPIS model is now being developed further to assist in the full analysis of the electron data set
- Next steps:
 - Better material and electrical properties
 - Deployment of 'instruments'
 - Separate EAS1 & EAS2

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• **THANK YOU FOR LISTENING**

Further details from: g.lewis@ucl.ac.uk