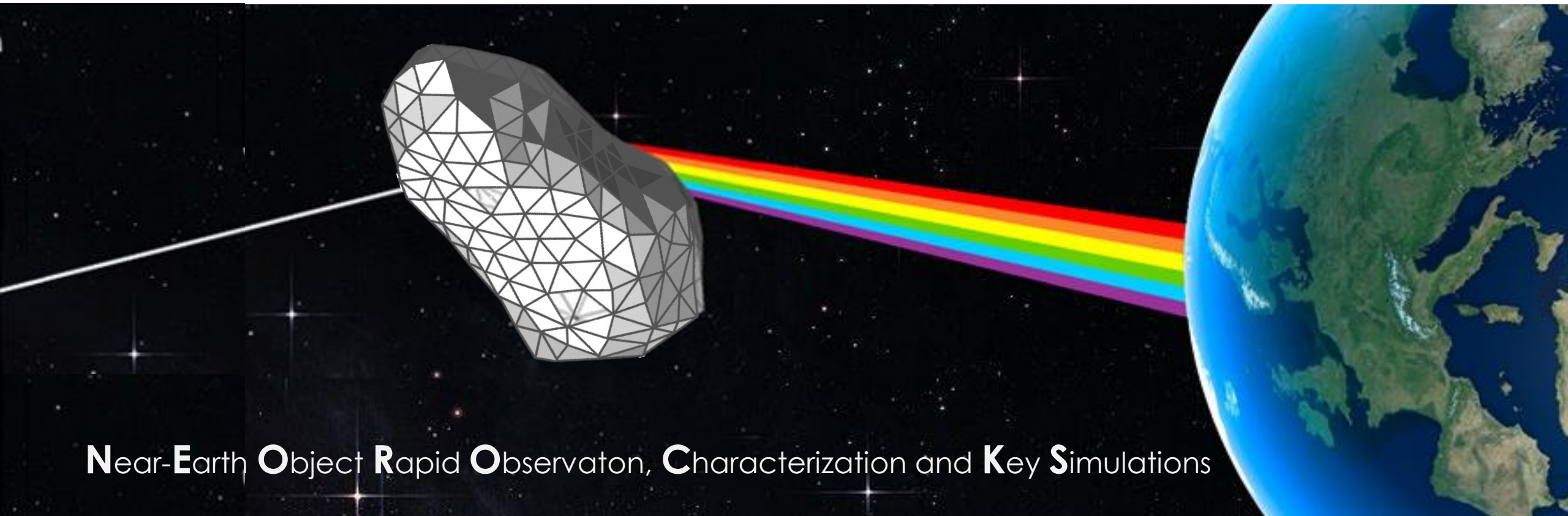


The NEOROCKS Rapid Response Experiment

lessons learned and future developments



Near-**E**arth **O**bject **R**apid **O**bservation, **C**haracterization and **K**ey **S**imulations

Elisabetta Dotto, Ettore Perozzi and the NEOROCKS Team



EU-ESA Workshop on NEO Imminent Impactors Warning Coordination 12-14 December 2022
ESOC - Darmstadt





PROGRAMME:

Horizon 2020 - Work Programme 2018-2020

Leadership in Enabling and Industrial Technologies – Space

Call: SU-SPACE-23-SEC-2019 – Advanced research in Near Earth Objects (NEOs) and new payload technologies for planetary defence.

European Commission Decision C(2018)4708 of 24 July 2018

TIMELINE:

Start: 1st January 2020

KOM: 20th January 2020

End: June 2023

Participant organisation name	Country
Istituto Nazionale di Astrofisica (coordinator)	Italy
Agenzia Spaziale Italiana	Italy
University of Padova	Italy
LESIA-Observatoire de Paris	France
Observatoire de la Cote d’Azur	France
University of Edinburgh	UK
Astron. Inst. of Czech Academy of Sciences	Czech Rep.
Instituto de Astrofisica de Canarias	Spain
SpaceDyS s.r.l.	Italy
DEIMOS Space s.l.u.	Spain
DEIMOS Space s.r.l.	Romania
DEIMOS Castilla La Mancha	Spain
NeoSpace sp z.o.o	Poland
Resolvo Srl	Italy

optimize observational activities, enhance modelling and simulation tasks, foster international coordination and speed-up response times

Ambition 1: Networking large aperture telescopes

Ambition 2: Advancing NEO physical properties modelling and simulations

Ambition 3: Improving the orbit determination process

Ambition 4: Addressing the imminent impactors monitoring

Ambition 5: Establishing a NEO physical properties data centre

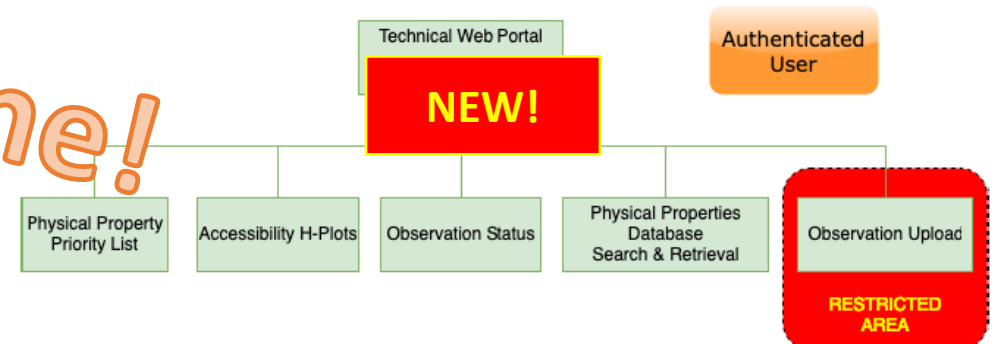
Ambition 6: Fostering international cooperation for follow-up

Ambition 7: Raise the public awareness on NEO and impact hazard



Done!

Done!





SUMMARY

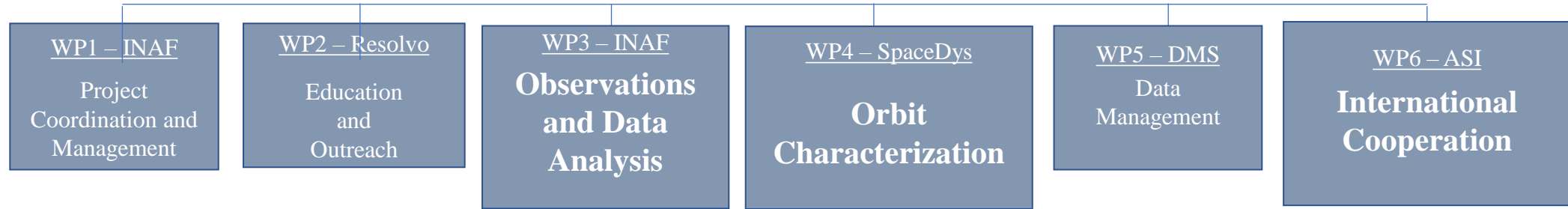
To plan and execute breakthrough experiments foreseeing the remote tasking of highly automatized robotic telescopes in order to provide a proof-of-concept rapid response system

AMBITION 4: ADDRESSING THE IMMINENT IMPACTORS MONITORING

To prepare for this scenario by prioritizing and coordinating quick follow-up observations either astrometric or physical. To this end, **an innovative tasking experiment of a rapid-response system will be organized** profiting of the experience gained by the DEIMOS Sky Survey (DeSS) in providing space debris observations in an operational environment

EXPECTED IMPACTS

To pave the way for the engineering developments needed to prototype a rapid response system able to task ground based telescopes, linking together for the first time astrometric and physical characterization follow-up observations.



WP3 – carry out physical characterization upon short notice through the project observational network



WP4 – implement a basic **rapid response system** scenario based on existing robotic telescopes through developing the related procedures and tasking SW



WP6 – analyze outcome and issue recommendations: *Final Report on international cooperation for NEO physical characterization and rapid response systems (D6.4)*

SCHEDULING TOOL



- Monitoring NEOScan Priority List https://newton.spacedys.com/neodys2/NEOScan/index_nspl.html
- Runs in loop mode
- Configuration variables
- Generates files ingested by sensor control (very particular to each sensor/control SW)

monitoring

Variable poll : for next period -> now to 60 min def
 Variable outputs 10 min def . intervals along next 60 min (6 lines per file)
 Created new files overwrite for same objects

From MPC site Z66 def

visibility

Above > Elev^o
 Below < magnitude

outputs




ASCII files to sensor (interpolated at precise observing time) (1 per NEOCP target)
 Filenames with provisional IDs and sorted by priority, 1 -> VU, 2 -> U, 3 -> N
 Provides PA and Angular Speed (arcsec/min)
 Provides suggested max expo times and n^o of frames required (at tasking level)

Track and stack

settings

Monitoring -> polling intervals, n^o outputs/gap
 Visibility -> MPC code, min elev,, max magnitude
 Sensor -> scale pix resolution, sensitivity: exp/mag factor x 4 : 15mag/0.5sec

<http://www.astrometrica.at/Papers/PointSources.pdf>

NEOScan Sponsored by    

Near Earth Objects - Dynamic Site

Last Update: 2022-03-01 19:06 UTC

NEOCP name	Priority class	Priority value	RA (hh:mm)	DEC (deg)	V mag.	ΔV mag.	Uncertainty (deg)	Sun elong. (deg)	Moon elong. (deg)	Gal. lat. (deg)
C09M2Q5	VERY URGENT	41.887	08:19	65°11'	21.60	4.83	0.67389	-116.5	-126.6	33.6
C7CN7Q2	VERY URGENT	34.760	12:48	11°37'	21.23	1.53	2.27370	150.5	143.6	74.5
C7CL7W2	VERY URGENT	32.117	11:58	7°42'	20.81	0.48	1.34208	163.1	154.5	66.9
C7CM312	VERY URGENT	31.748	12:09	15°32'	21.05	1.10	0.82168	158.8	153.4	74.9
C7CLJC2	VERY URGENT	31.678	12:19	10°25'	21.22	0.38	0.64130	157.8	150.3	71.7
C7CJN22	VERY URGENT	31.322	11:34	12°50'	22.02	0.72	0.16788	167.8	161.5	67.0
C7CNXB2	VERY URGENT	29.523	13:24	10°35'	20.34	0.26	0.36478	141.7	134.7	71.7
P21s58d	VERY URGENT	29.401	08:44	-5°01'	21.18	0.09	0.05123	-146.3	-148.6	22.3
P21rKwH	VERY URGENT	28.352	14:23	-22°59'	21.65	0.07	0.00635	119.7	108.9	35.2
C7CP042	VERY URGENT	27.170	13:29	11°01'	21.52	0.32	0.45345	140.6	133.7	71.6
C09M2R5	VERY URGENT	26.829	12:05	40°38'	22.51	0.19	0.02476	142.8	146.2	73.5
P21s7YS	URGENT	20.119	13:18	-1°11'	20.60	0.16	0.18079	141.9	132.5	60.9
P21s6qA	URGENT	18.804	13:12	-22°21'	21.34	0.19	0.03555	134.1	123.0	40.3
C09N205	URGENT	15.818	13:16	54°25'	22.17	0.19	0.20855	124.6	128.5	62.3
C7CPMQ2	URGENT	12.536	16:17	11°48'	19.87	0.55	0.65590	99.4	93.6	39.4
C7C3862	URGENT	11.463	11:28	14°16'	21.79	0.24	0.29335	168.2	163.1	66.9
C09M105	NECESSARY	8.974	08:40	64°12'	22.21	0.02	0.00427	-118.8	-128.8	36.1
C09LEN5	NECESSARY	8.958	16:24	15°46'	22.06	0.01	0.00181	98.1	93.0	39.5
P21rKyC	NECESSARY	8.875	13:17	-14°31'	22.11	0.02	0.00720	137.3	126.5	47.9
C7CGKG2	NECESSARY	8.509	10:50	9°47'	21.69	0.09	0.04297	-177.6	170.3	56.7
C09M3H5	NECESSARY	8.331	12:07	39°28'	22.48	0.02	0.00132	143.5	146.6	74.6
P21rLPs	NECESSARY	7.879	13:16	32°13'	21.95	0.02	0.00311	137.8	136.6	82.5
P21s7YT	NECESSARY	7.851	13:19	-1°52'	20.07	0.11	0.13925	141.5	132.0	60.2
C7CPFC2	NECESSARY	7.251	15:16	11°26'	19.53	0.03	0.13522	114.3	108.0	52.6

CAMPAIGN

Tracker2: 40 cms aperture! -> limitations
 NEOCP current magnitudes -> very few available, even less if only from priority list/criteria
 Full moon period (Started bit too late...)
 Last week: <https://minorplanetcenter.net/mpec/K22/K22CE7.html>
<https://minorplanetcenter.net/mpec/K22/K22CE2.html>
<https://minorplanetcenter.net/mpec/K22/K22CE0.html>
 Not all get MPECed -> currently selecting -> trying priority list, already indept. confirmed, high NEO score



MPEC 2022-C140 : 2022 CE3
 Issued 2022 February 8, 01:51 UT
 MPEC 2022-C147 : 2022 CM3
 Issued 2022 February 8, 02:02 UT



Da: Julia de Leon <jmlc@iac.es>
 Date: mer 16 feb 2022 alle ore 14:55
 Subject: Re: Observation Campaign
 To: Dotto, Elisabetta <elisabetta.dotto@inaf.it>, Perna, Davide <davide.perna@inaf.it>
 Cc: Marcel Popescu <popescu.marcel1983@gmail.com>, Javier Licandro <jlicandr@iac.es>, David Morate <damog@iac.es>

Dear Elisabetta,

we tried to observe 2022 CE3 on the night of Feb. 10 using the NOT and our service program, but unfortunately the weather was awful. Same thing at the Teide Observatory, where we tried with the 1.5m TCS... What a bad luck!

I know you were mostly interested in doing a taxonomical classification, but Marcel Popescu was able to get some astrometry using a robotized 25 cm telescope installed at his institution in Bucharest. He observed both 2022 CE3 and 2022 CM3 on the night of Feb. 9 and has reported results to the MPC (attached files here), which have been accepted.

Orbital elements:
 2022 CE3
 Epoch 2022 Jan. 21.0 TT = JDT 2459600.5
 M 12.52622 (2000.0) P Q
 n 0.44632914 Peri. 315.00504 +0.27146466 -0.96116546
 a 1.6957659 Node 119.18397 +0.89826054 +0.23448811
 e 0.5108843 Incl. 3.26195 +0.34559360 +0.14552076
 P 2.21 H 24.22 G 0.15 U 7
 Earth MOID = 0.0123 AU
 Veres

Residuals in seconds of arc

220206 F52 (1.1- 0.3+)	220207 L01 0.8- 0.6-	220207 B49 1.1- 0.8+
220206 F52 0.0 0.0	220207 L01 1.2- 0.3+	220207 C95 0.0 0.5-
220206 F52 0.0 0.0	220207 L01 0.7- 0.5-	220207 C95 0.6+ 0.1-
220207 703 0.5- 0.3+	220207 L01 0.1- 0.2+	220207 B49 0.1+ 0.1+
220207 703 0.2+ 0.2-	220207 L01 0.5- 0.4-	220207 K87 0.0 0.5-
220207 152 0.4+ 0.1-	220207 A17 0.0 0.1+	220207 130 0.4- 0.7-
220207 152 0.3+ 0.2+	220207 A17 0.1+ 0.0	220207 C95 0.2+ 0.0
220207 152 0.4+ 0.2-	220207 104 0.0 0.1-	220207 130 0.1- 0.1+
220207 152 0.3+ 0.0	220207 104 0.1+ 0.1-	220207 K30 0.0 0.3+
220207 734 0.0 0.3+	220207 104 0.1- 0.1-	220207 Z66 1.9+ 1.3-
220207 734 1.6- 1.1-	220207 A17 0.1+ 0.1-	220207 K87 0.5- 0.6-
220207 734 0.1- 0.3+	220207 204 0.1+ 0.2+	220207 K30 0.4+ 0.4+
220207 734 0.2+ 0.0	220207 204 0.0 0.0	220207 K74 0.1- 0.0
220207 734 0.5+ 0.1+	220207 204 0.1- 0.0	220207 Z66 2.4+ 1.4-
220207 734 0.4- 0.4+	220207 G15 0.1+ 0.4+	220207 K30 0.4- 0.1+
220207 734 0.3- 0.1+	220207 G15 1.1- 0.1+	220207 K74 0.1+ 0.1+
220207 152 0.2+ 0.2+	220207 K63 0.3- 1.1-	220207 Z66 2.4+ 1.5-
220207 152 0.1+ 0.2+	220207 033 0.1+ 0.0	220207 K87 0.2+ 0.3+
220207 152 0.1+ 0.3-	220207 033 0.3+ 0.0	220207 K74 0.3+ 0.2-
220207 152 0.1- 0.0	220207 033 0.0 0.1+	220207 G14 0.4- 0.4+
220207 152 0.4+ 0.0	220207 K63 0.1+ 0.0	220207 Z33 0.0 0.2-
220207 152 0.3+ 0.0	220207 G15 0.1- 0.7+	220207 G14 0.9- 0.6+
220207 152 0.1+ 0.3-	220207 K63 0.0 1.2-	220207 Z33 0.0 0.2-
220207 152 0.2+ 0.1+	220207 K51 0.2+ 0.9+	220207 G14 0.9- 1.4+
220207 T08 0.4- 0.8+	220207 K51 0.0 0.3+	220207 Z33 0.4+ 0.2+
220207 T08 2.0+ 1.3-	220207 K51 0.2- 0.1+	220207 Z33 0.1- 0.2+
220207 T08 0.1+ 0.3+	220207 B49 0.3+ 0.1+	
220207 T08 0.5- 0.6+	220207 B49 0.0 0.3+	

Z66 DeSS Deimos Sky Survey, Niefila Mountain. Observers J. Nomen, M. Ortega. 0.40-m f/5.5 reflector + CCD.



TARGET SELECTION

- ✓ Before new moon (to allow more useful nights)
- ✓ Approaching targets (V increases)
- ✓ Not too high declination (enable southern obs)
- ✓ Not too faint (enable spectroscopic observations)
- ✓ Not too low solar elongation
- ✓ Targets for which DESS astrometry can count
- ✓ Size of an imminent impactor (H range 22-27)
- ✓ Small telescopes need long integration times
- ✓ Use telegram for speeding up communications?



Expect to continue campaign next week (21/02) -> until? , current availability
 Improve residuals
 Time consuming for small sensor -> integrating light
 Manual procedures: T&S and visual evaluation always required:
 Confirm existence -> NEOCP -> DNE
 Uncertainty on magnitude and positions
 Features: cometary appearance, fast rotators

Shortening times from detection and first assessment
 Geographical constraint for Z66 or/and for other sites?
 Other objectives of the campaign?



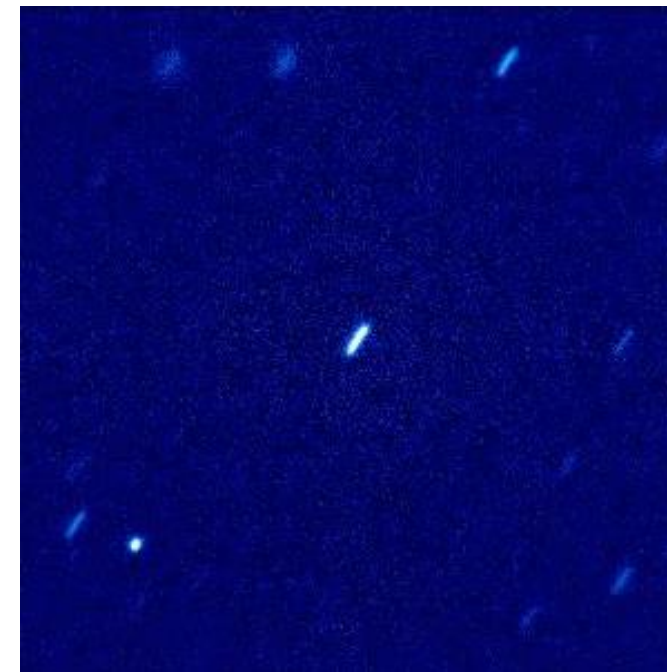
add a «blind» experiment option



In the night between 3 and 4 April 2022 DEIMOS and SpaceDys coordinated in order to perform highly automatized astrometric observations of a sample of newly discovered NEOs

The information was quickly disseminated within the consortium in order to allow rapid follow-up observations of the targets' physical properties.

Asteroid 2022 GC1 has been the subject of a rapid DDT (Director Discretionary Time) request to TNG (Telescopio Nazionale Galileo, Canary Islands).

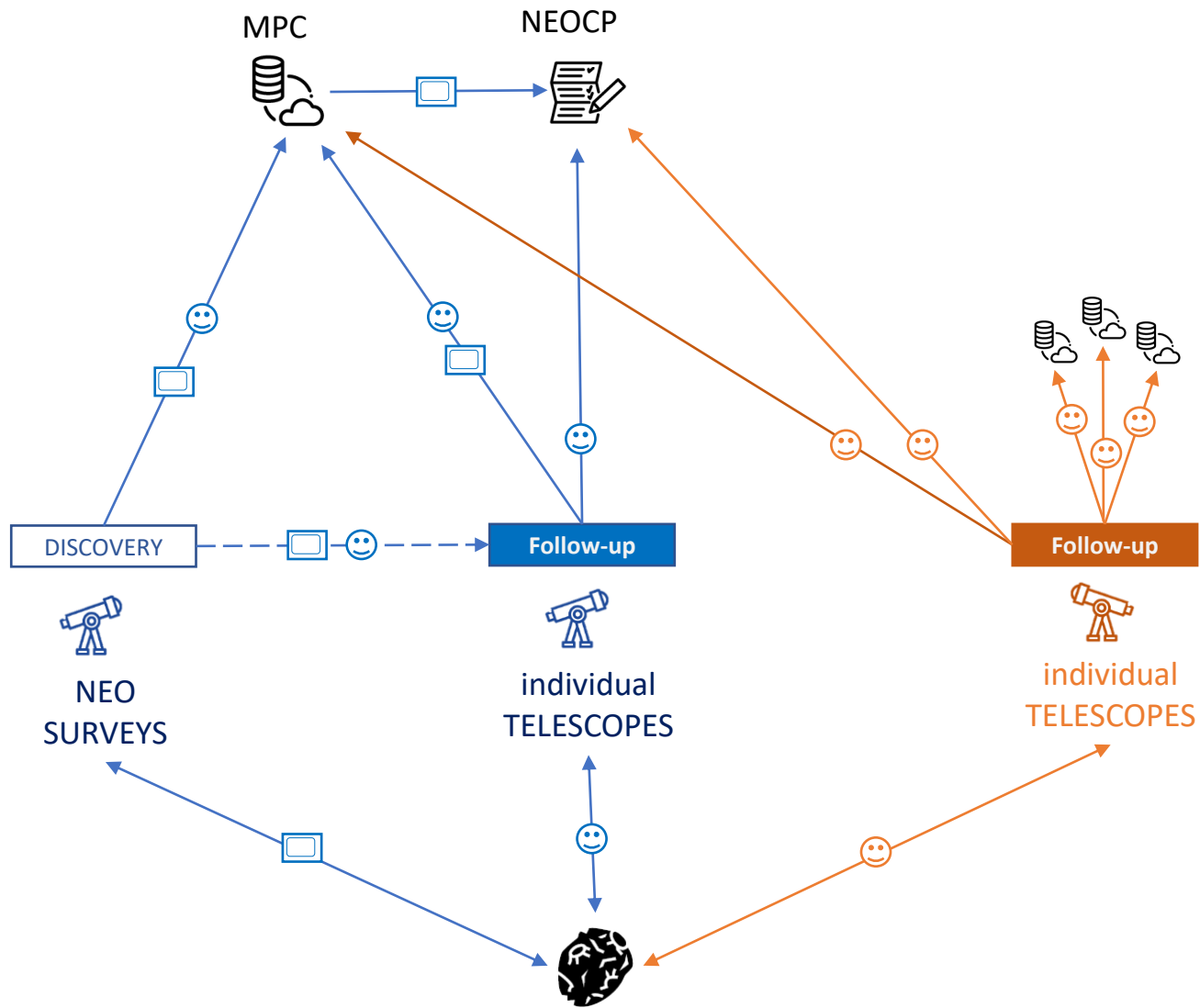


at the time of observations, 2022 GC1 was at 1.1 au from the Sun and 0.12 au from our planet.

In the night between 5th and 6th of April, despite a high proper velocity (about 200 arcsec/h), standard BVRI photometry was successfully obtained from TNG , thus closing the loop within days from discovery.



PRESENT



tasking
from → to

astrometric

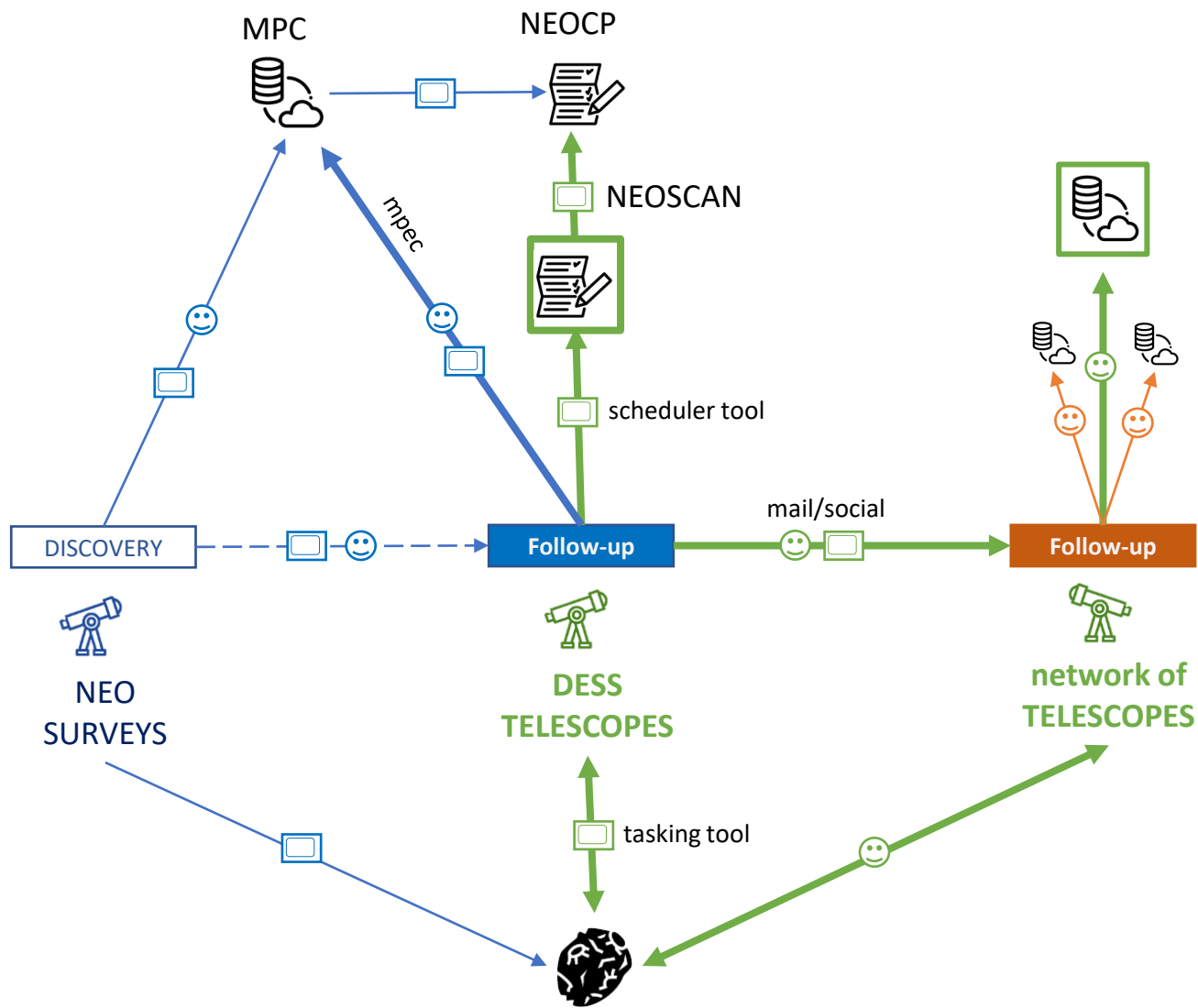
physical

human ☺

software 🖥️

data base 🗄️

tool 📄



tasking
from → to

astrometric

physical

human ☺

software 🖥️

data base 🗄️

tool 📄

NEOROCKS

being able to timely carry out the physical characterization of an imminent impactor is a key feature for developing a rapid response system for civil protection purposes

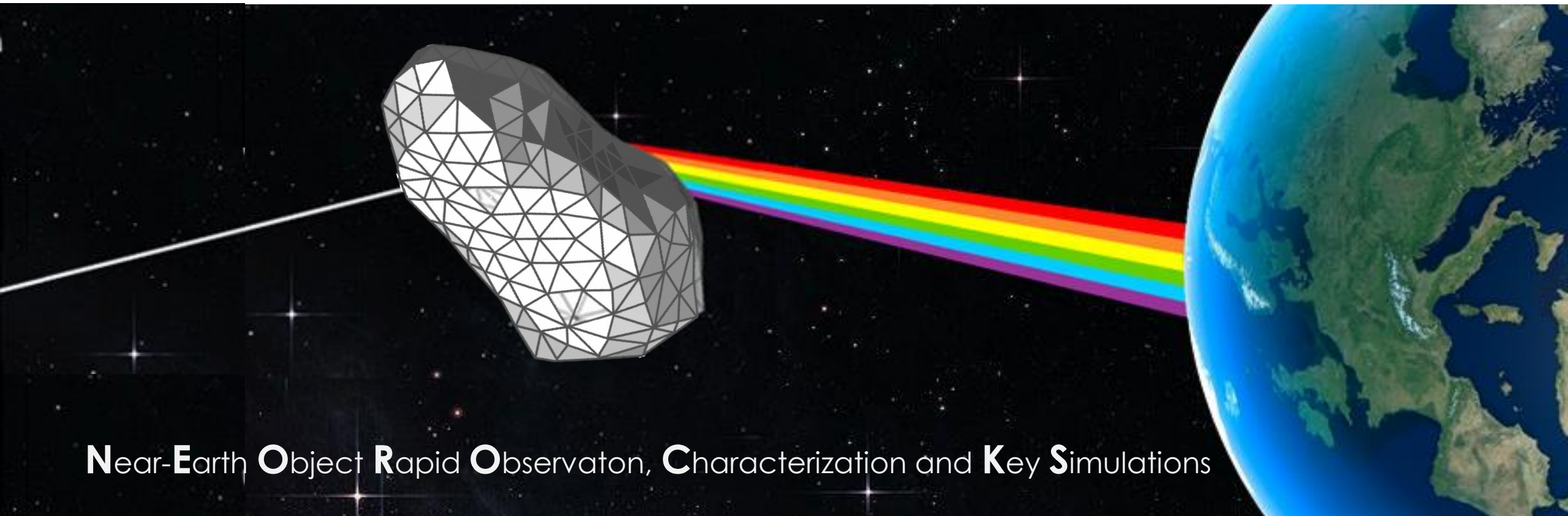
we have demonstrated that using European existing assets it is possible to set-up an integrated imminent impactor rapid response system encompassing both, dynamical and physical characterization.

Future is to turn an experiment into a prototype through e.g.

- Extending the network of telescopes that can be directly tasked for astrometry
- Automating as far as possible the tasking of large telescopes for physical characterization (e.g. use social media?)
- Obtaining priority tasking of large telescopes through institutional empowerment (EU, IAU, UNOOSA...)
- Taking into account the need for secure communication
- (...)

The NEOROCKS Rapid Response Experiment

lessons learned and future developments



Near-**E**arth **O**bject **R**apid **O**bservation, **C**haracterization and **K**ey **S**imulations

Elisabetta Dotto, Ettore Perozzi and the NEOROCKS Team



EU-ESA Workshop on NEO Imminent Impactors Warning Coordination 12-14 December 2022
ESOC - Darmstadt

