Feasibility study on the recycling of a close to end-of-life optical Earth-observing satellite constellation for orbital debris detection

Introduction[•]

The growing amount of space debris is a phenomenon that dates back to the first man-made object launched into space: the Soviet spacecraft Sputnik-1. Due to the exploration of space only being a recent technological advancement with scientists and engineers operating under the assumption that space was "self-cleansing" (Schefter, 1982), the 60s and 70s were plagued with space missions that gave no regard to their sustainability.

By the late 1970s, data analysis and population simulations of existing debris catalogues helped bring to light the worsening situation. One particular threat was the cascading effect of debris colliding with each other and creating more debris: this was coined the "Kessler Syndrome" after NASA astrophysicist Donald J. Kessler.

Despite knowledge of this phenomenon and its repercussions, the ever-growing numbers of yearly space launches alongside various events such as in-space collisions and anti-satellite tests have all contributed to the population of objects larger than 10 com being at around 30 000.



▲ Figure 1: Evolution of space debris (from ESA)

However, a 1cm large piece of debris would be enough to cause serious damage and potentially disable a spacecraft; current estimations place the population of debris of this size at around 1 100 000. These pieces of debris pose a real threat to active satellites. As an example, the ISS has had to perform on average 1.3 collision avoidance manoeuvres a year since its launch; this statistic is predicted to increase with time.

In order to perform the collision avoidance manoeuvres, satellite organizations require accurate debris orbit predictions. However, ground-based observations have limitations:

- The atmosphere prevents the observation of objects smaller than around 2 cm. • Accurate prediction techniques require long observation periods and accurate object
- identification to reduce errors.

This work aims to address the second point, by studying the feasibility of recycling a soon-to-bedisposed Earth observing constellation, rescoping it for debris observation. This study will focus on comparing the recycled constellation's observation access time to that of current groundbased observation capabilities. The difference in observation access times will determine the "value for money" of this recycling, which can be divided into two points:

- Observation access times of a piece of debris will be longer, leading to better orbit predictions with fewer errors
- There will be no need to launch new payload to achieve these new observation access times which will help contribute to the hope of a space circular economy

-Systems selection

In order to compare the observation access time of debris between a recycled constellation and a ground station, they need to be defined. The constellation, Cosmo-SkyMed, has been chosen based on the following: • It must be in LEO.

• It must possess optical observation capabilities with a reasonable resolution.

The ground stations have been chosen from ESA's Space Surveillance and Tracking segment; one optical and one radar: OGS & TIRA.

Name	φ (°)	λ (°)			
OGS	28.30096	-16.5118			
TIRA	50.6166	7.1296			
▲ Table 1: Coordinates of the chosen ground station					

Name	φ (°)	λ	λ (°)		h (m)	
OGS	28.30096	-16.	51182	239	96	
TIRA	50.6166	50.6166 7.1296		264		
▲ Table 1: Coordinates of the chosen ground stations						
Name	Altitude (km)	Ecc.	Inc. (°)	RAAN (°)	AoP (°)	
Cosmo-SkyMed 1	619.6	≈0.0001	97.86	29	84	
Cosmo-SkyMed 1 Cosmo-SkyMed 2	619.6 619.6	≈ 0.0001 ≈ 0.0001	97.86 97.86	29 29	84 89	
Cosmo-SkyMed 1 Cosmo-SkyMed 2 Cosmo-SkyMed 3	619.6 619.6 619.6	≈ 0.0001 ≈ 0.0001 ≈ 0.0001	97.86 97.86 97.86	29 29 44	84 89 90	
Cosmo-SkyMed 1 Cosmo-SkyMed 2 Cosmo-SkyMed 3	619.6 619.6 619.6	≈ 0.0001 ≈ 0.0001 ≈ 0.0001	97.86 97.86 97.86	29 29 44	84 89 90	

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Method



figure 7.

Element	Field Bounds	Step
а	7250 km ± 50 km	2 km
е	0.0095	n/a
i	98° - 100°	1 °
Ω	TBD ± 15°	0.5°
ω	TBD ± 15°	0.5°

and outside of a 35° satellite-to-Sun exclusion zone.

radar constraints are removed, as focal length and sunlight are no longer a factor. For optical observations, the debris must be sunlit, and the ground station must not.

Note: the SightGS and SightSat functions in the algorithm are the line-of-sight computations. The prop function is for object propagation.



However, in the time being, if solving the focal length limitation problem can be achieved, by either rescoping satellites with long-range radars or variable focal length optical cameras on board, the recycling of an Earth-observing constellation after its intended mission for debris observation would have the following benefits: not only would it lead to less errors in debris orbit determination and the ability to keep a better track on them, but it would also contribute to a circular space economy and prevent the need to launch new payload into space.

