



RBSP-ECT and complementary datasets

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Overview

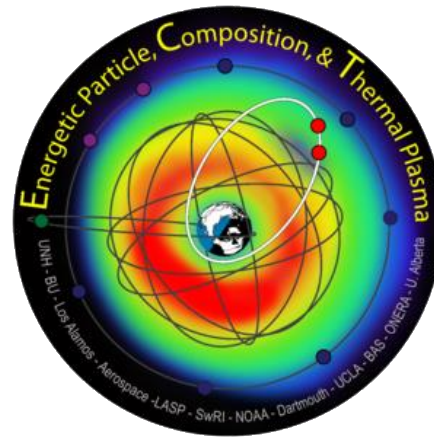


RBSP-ECT

- Unified data products for the ECT suite, combining electron measurements from HOPE, MagEIS and REPT
- Coverage for whole mission from 2012-2019
- **Complete electron spectra throughout the inner magnetosphere from 10s eV to 10s MeV**

REACH

- Set of 32 dosimeter payloads on IRIDIUM-Next
- Data is now publicly releasable
- Data coverage from 2017-2019, 2022-Present





RBSP-ECT

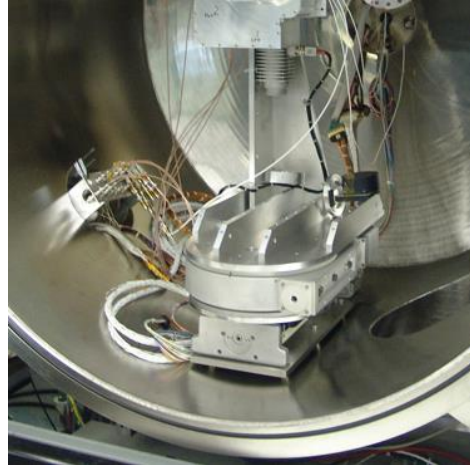


Energetic Particle Composition and Thermal Plasma (ECT) Suite

Instrumentation

- Van Allen Probes: 2 satellites in GTO orbit (Aug 2012 – Jul/Oct 2019)
- HOPE (Helium Oxygen Proton Electron Spectrometer)
 - ~ 10 eV – 50 keV electrons, ions
- MagEIS (Magnetic Electron Ion Spectrometer)
 - 20 keV – 4 MeV electrons
 - 60 keV – 20 MeV Protons
- REPT (Relativistic Electron Proton Telescope)
 - 1.8 MeV – 20 MeV electrons
 - 20 MeV – 2 GeV Protons

HOPE



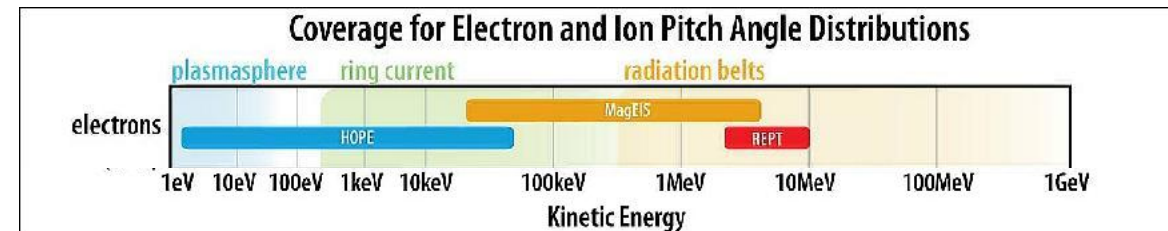
MagEIS



REPT



Image Credits: NASA



Mauk et al. (2012)

* Additional measurements of electrons/ions from RBSPICE (20 keV-1MeV) and protons from RPS (50-2000 MeV)

More than 7 years of electron spectra throughout the inner magnetosphere from 10s eV to 10s MeV

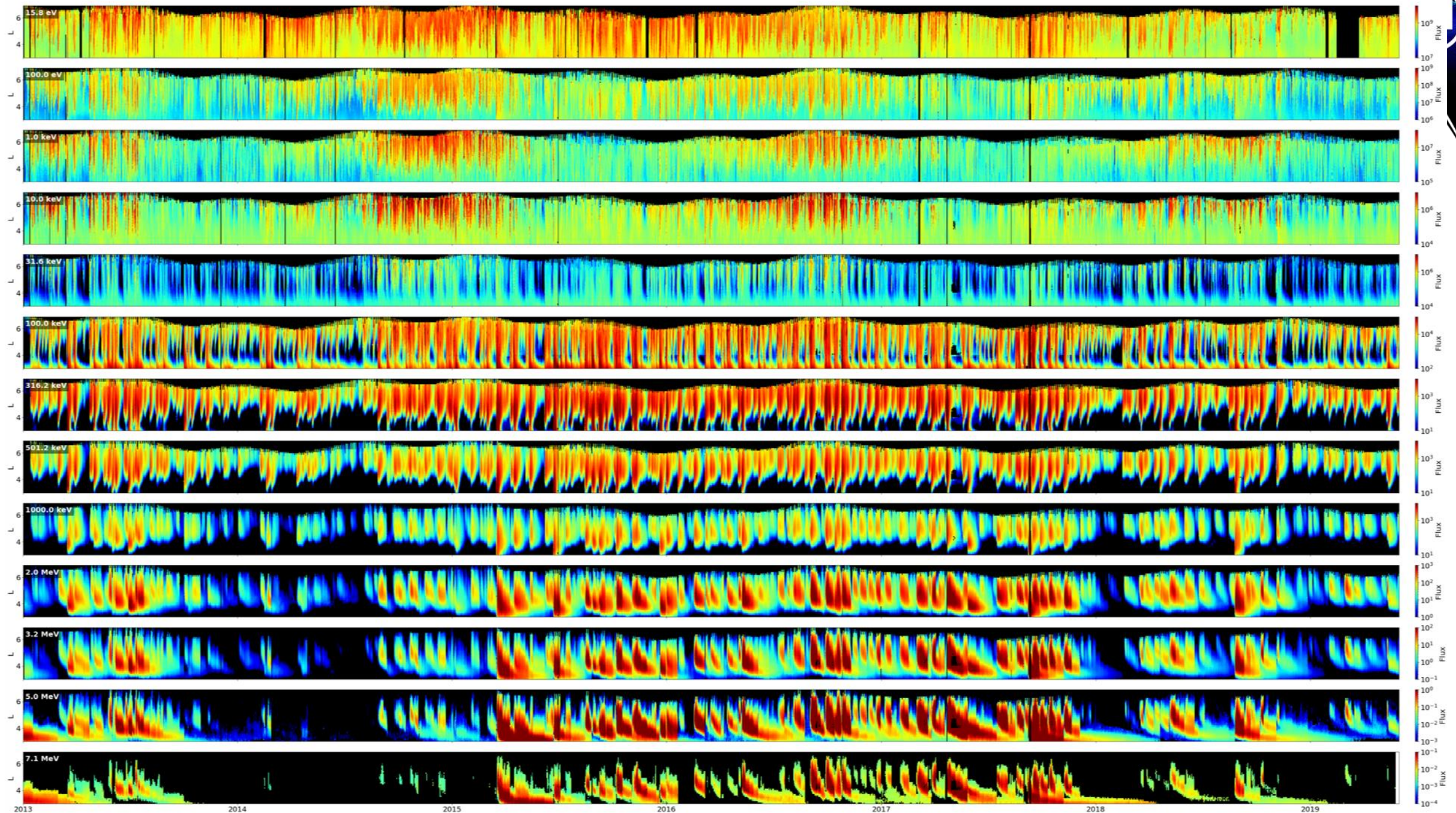


ECT

Combined Data Products

- Combine the HOPE, MagEIS, REPT electron data in one reliable, easy-to-use format
 - Removing low count, out-of-family channels
 - Providing data quality flags
 - Combined spectra (ex. FESA) and spline fit spectra (ex. FESA_FIT)
- Data products:
 - L2: Spin-Averaged Fluxes
 - L3: Pitch Angle Resolved Fluxes
- 3-minute averaged electron fluxes (label in middle of bin)
- The spin-averaged data product (L2) and Pitch angle resolved product (L3) is available for entire mission (2012-2019):
 - https://rbsp-ect.newmexicoconsortium.org/data_pub/rbspa/ECT/level2/
 - https://rbsp-ect.newmexicoconsortium.org/data_pub/rbspa/ECT/level3/
 - Accompanying paper in JGR:
 - <https://doi.org/10.1029/2019JA026733> (L2)
 - <https://doi.org/10.1029/2020JA028637> (L3)

Future entryway into the ECT data

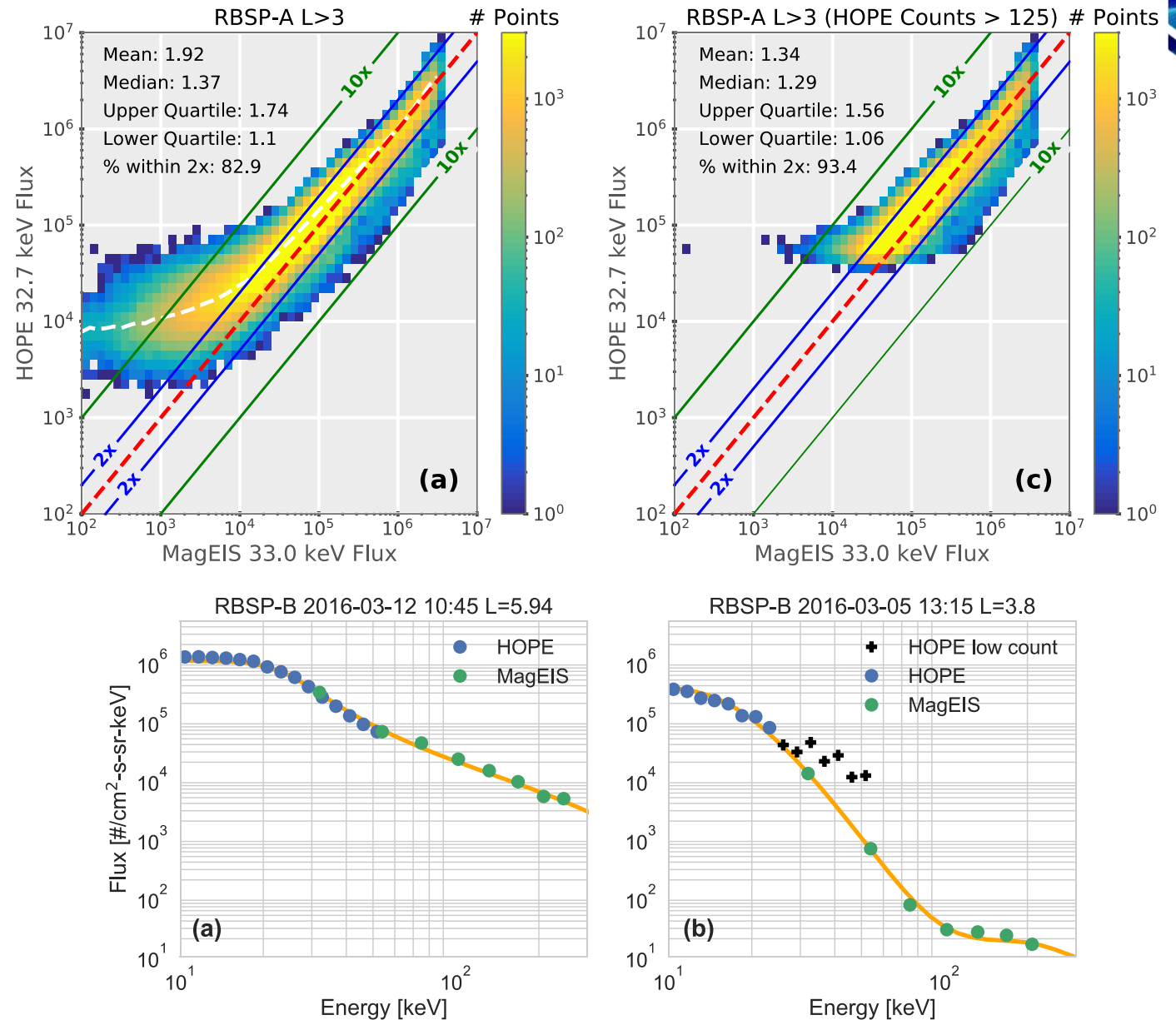




HOPE-MagEIS

Spin-Averaged

- Comparison of HOPE/MagEIS 32/54 keV channels
- Remove times where HOPE has few counts (<125 in 3 minute window)
- When low count times removed, excellent agreement (92% within a factor of 2)
- Apply correction to HOPE (based on 32 keV ratio):
 - *Inner Zone* ($L < 2.5$)
 - *Days with large disagreement* (efficiency correction)

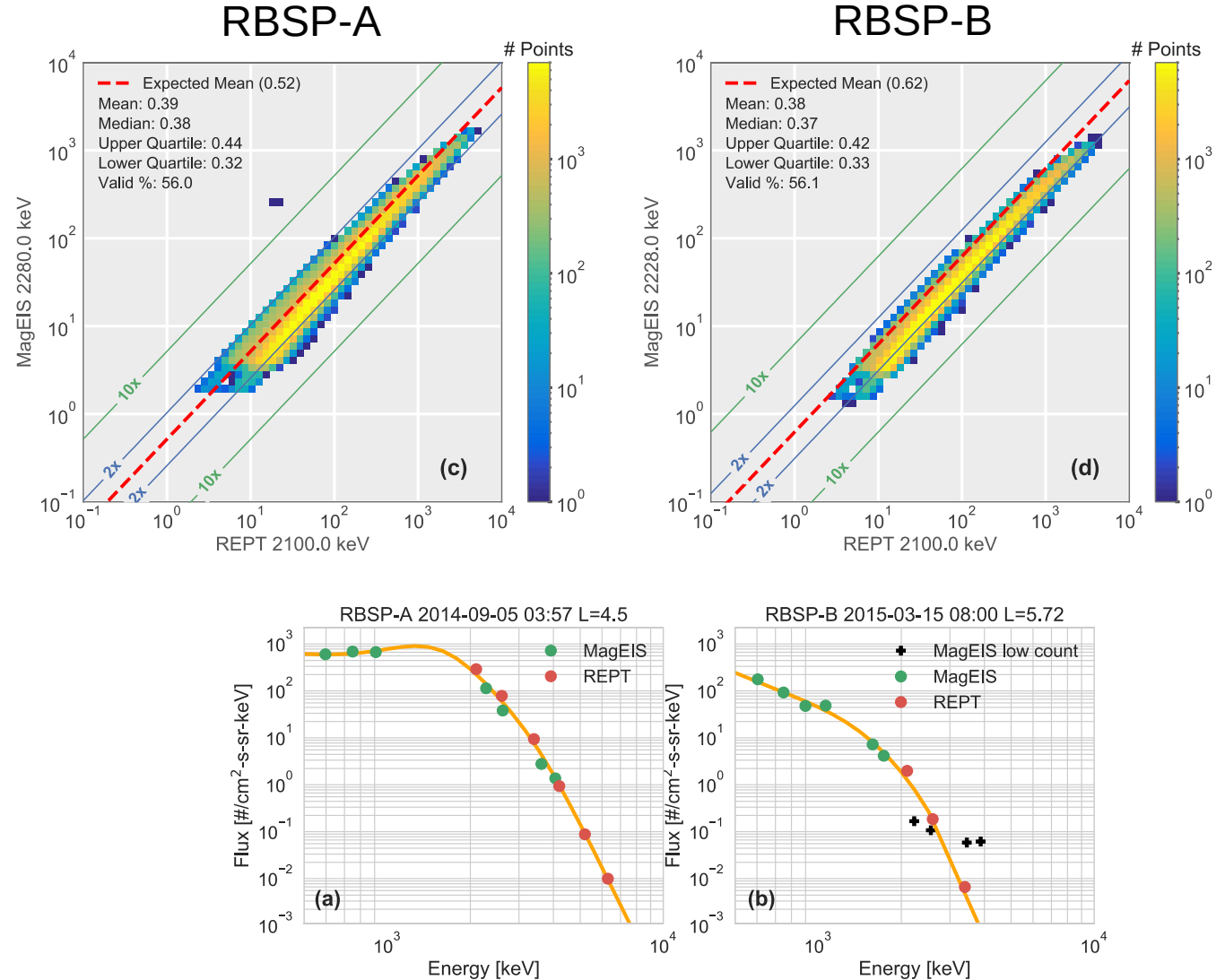


Boyd et al. (2019)

MagEIS-REPT

Spin-Averaged

- Comparison of MagEIS/REPT 2.1/2.6 MeV channels
- Removing low count times (estimated error > 75%), excellent agreement (92% within a factor of 2)
- 1.8 MeV REPT channel not included
 - Factor of ~2.5 offset
- MagEIS channels above 3 MeV not included
 - Very few points with enough counts



Boyd et al. (2019)

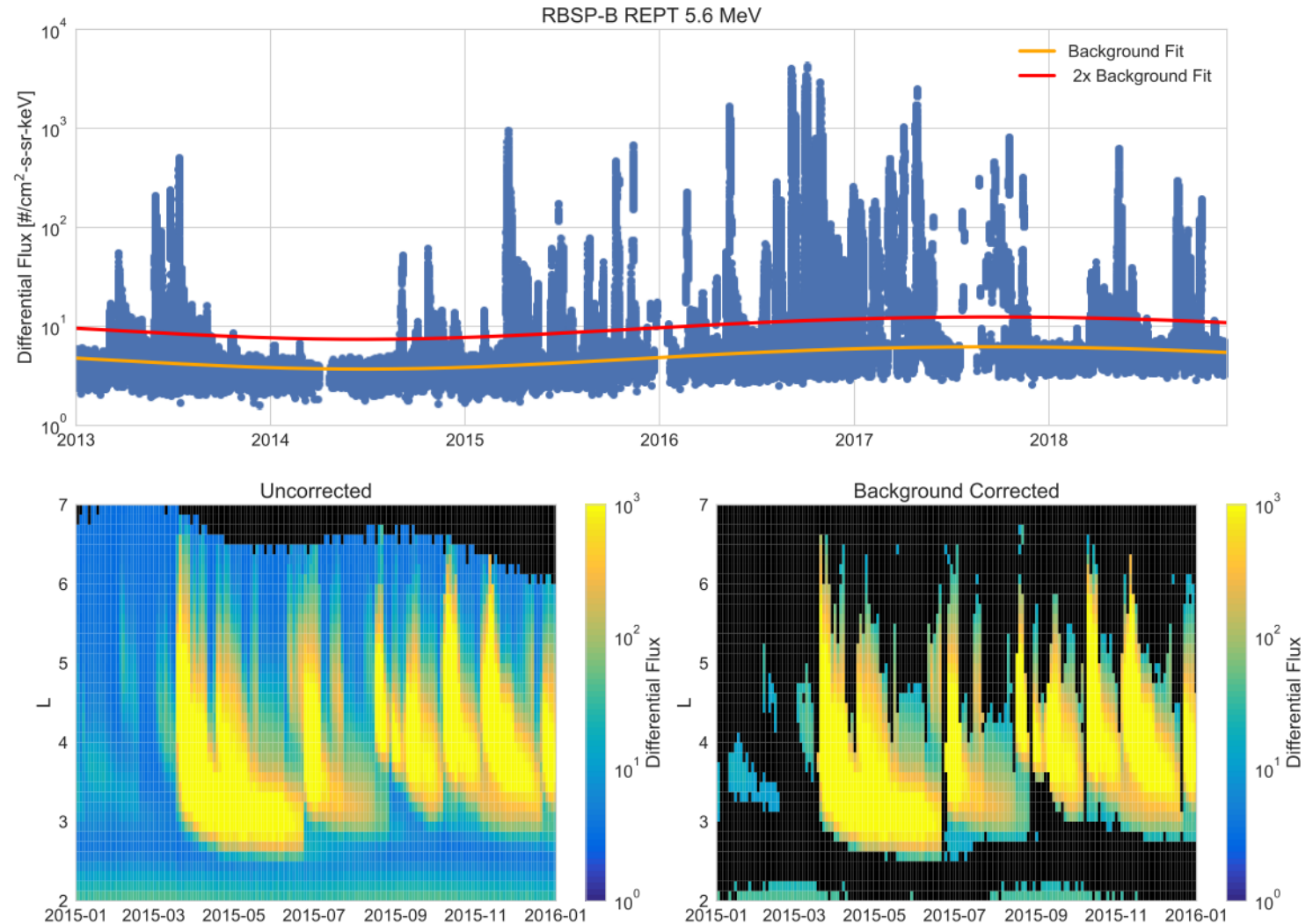
REPT Background

Spin-Averaged

- Outside of major enhancements, higher REPT channels (> 4 MeV) see background due in part to GCR
- To identify these times we apply a crude time-varying background fit:

$$E(t) = A + B \sin \frac{\pi t}{C} + D \sin \frac{\pi t}{C}$$

- Remove all times where flux < 2*fit
- At lower energy channels single count level is used

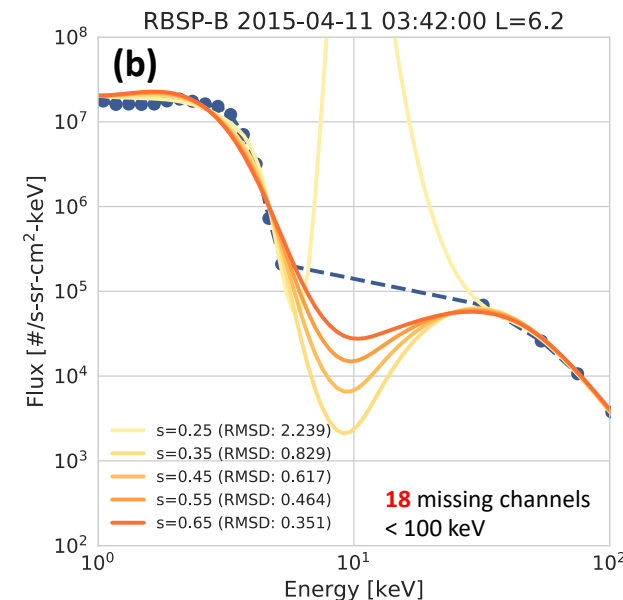
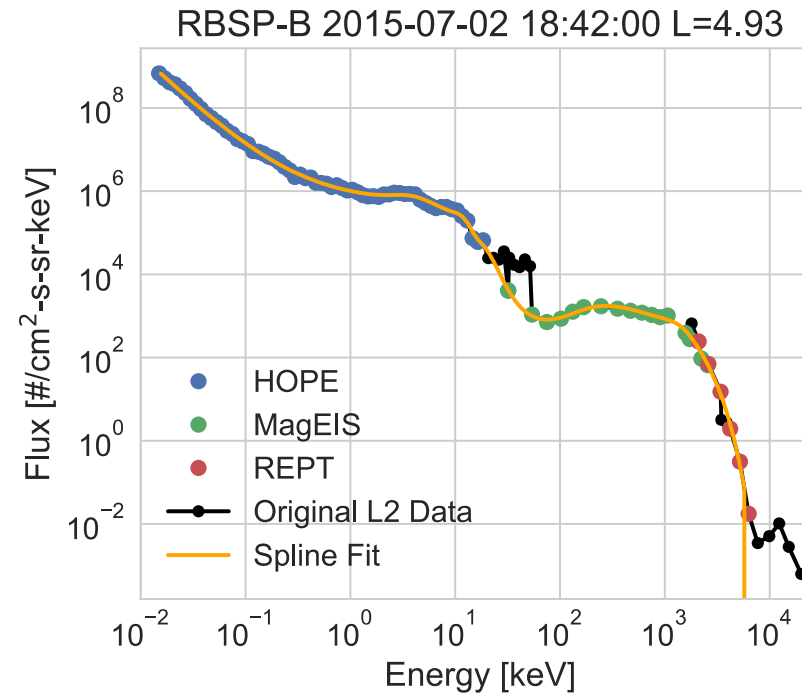


Boyd et al. (2019)



Spline Fitting

- Final step is a spline Fit of the combined spectra
- Output 127 log spaced energy channels
- Provide the knots and coefficients to allow user to reconstruct the fit
- Validation checks (increase smoothing or leave data gap in spline):
 - *Data gaps*
 - *Large 2nd derivative*
- Fit Quality Flag

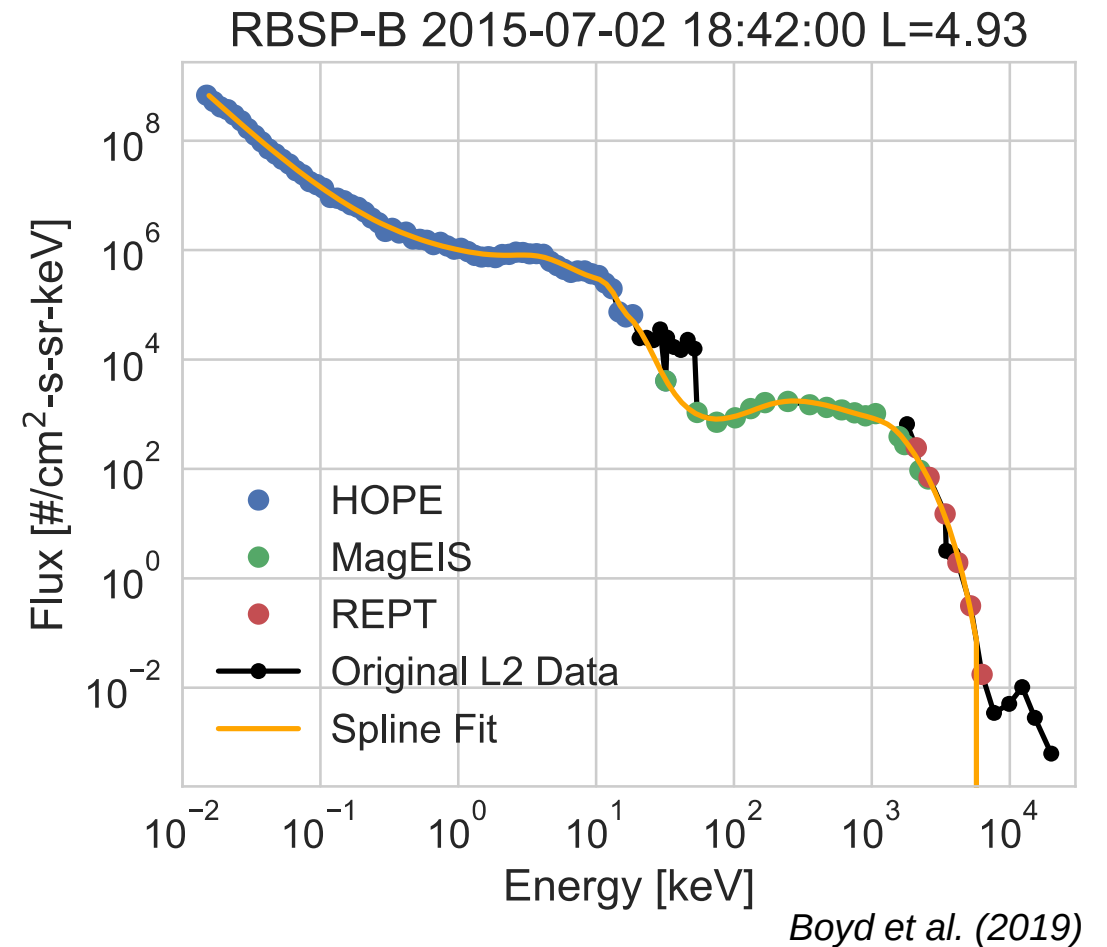


Boyd et al. (2019)



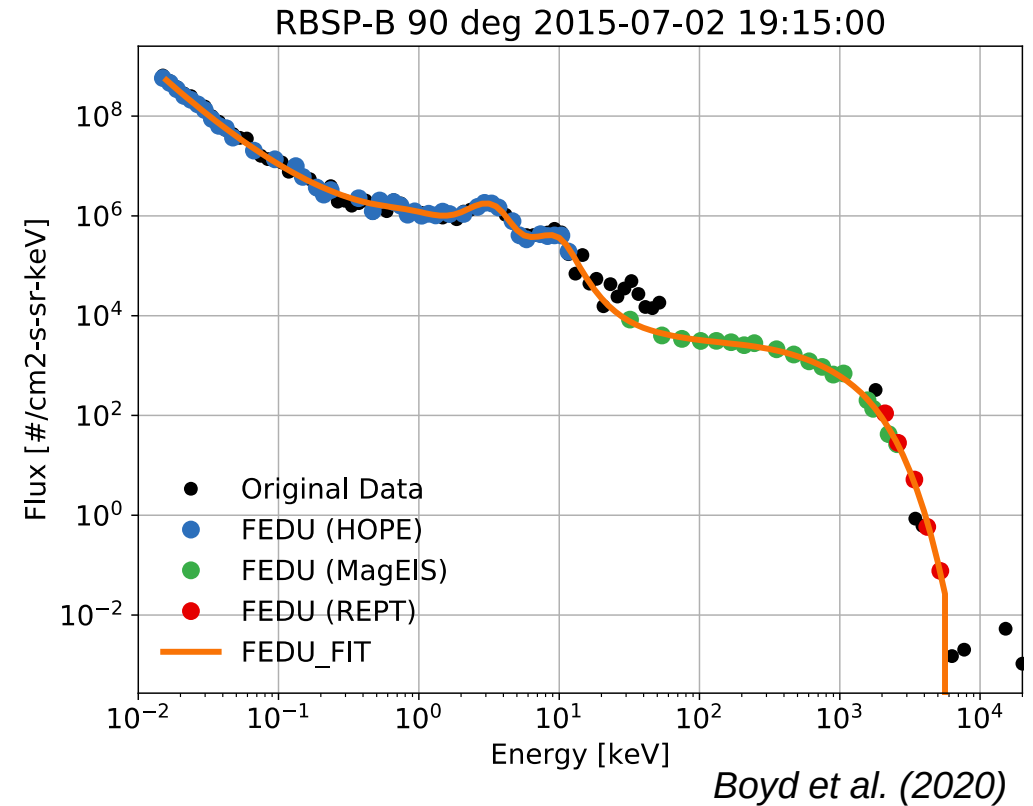
What's in the data files

- Combined Flux Spectra
 - *FESA (102 energies)*
 - *Data Quality Flags (FESA_Quality)*
- Spline Fit Spectra
 - *FESA_FIT (127 energies)*
 - *Data quality flags (FESA_FIT_Quality)*
- Support Data
 - *Magnetic Ephemeris*
 - *Spline Coefficients and Knots*



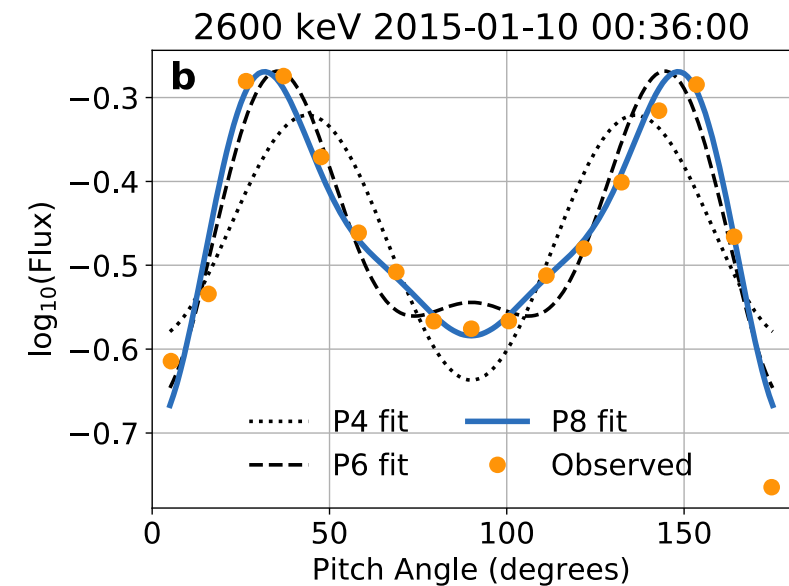
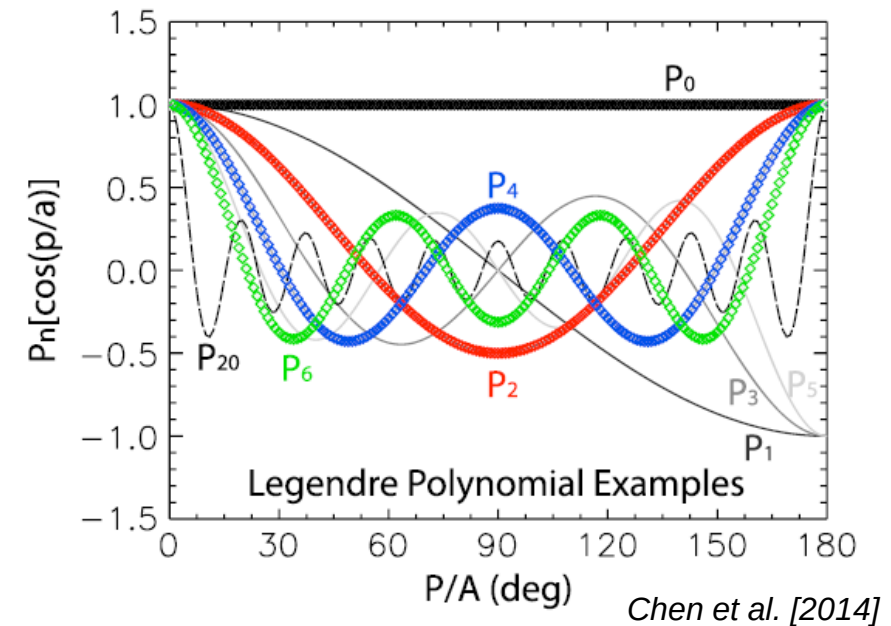
Pitch Angle Resolved Data Product

- Follow same basic methodology as L2
 - *HOPE*
 - Remove low count channels
 - Correction factor based on ratio at 90 degrees
 - *MagEIS*
 - HIGH channels where FEDU_ERROR < 75
 - *REPT*
 - Same background correction (if spin-averaged below background all PA bins are set to 0)
- Each of the instruments have different pitch angle bins, so need to fit angular distributions output at 35 pitch angles (5,10,15...170, 175)
- Calculate omnidirectional flux (FEDO, FEDO_FIT) and equatorial flux (FEDU_Eq)



Pitch Angle Fits

- Legendre Polynomials with $\log_{10}(\text{flux})$
- Most fits only use even terms
- Estimate symmetry (determine whether to include odd coefficients):
 - Mean ratio of 1st/last 4 PA pairs (e.g. 8°/172°)
 - If < 0.75 or < 1.3 include odd coefficients, otherwise only the even
- Each PA is fit using same spline routine
- FEDU_FIT (127 energies, 35 pitch angles)



Boyd et al. (2020)

Caveats



- The combined products will be updated as final instrument data is available
- Inner Zone ($L < 2.5$)
 - *All fluxes above 1 MeV are not included (this does not mean these are always 0)*
- Early Mission (before Sept 2013)
 - *Changing instrument modes/energy channels can lead to unexpected results*
- Pitch Angle shapes in FEDU_FIT can be inconsistent
- **Use the data quality flags!**

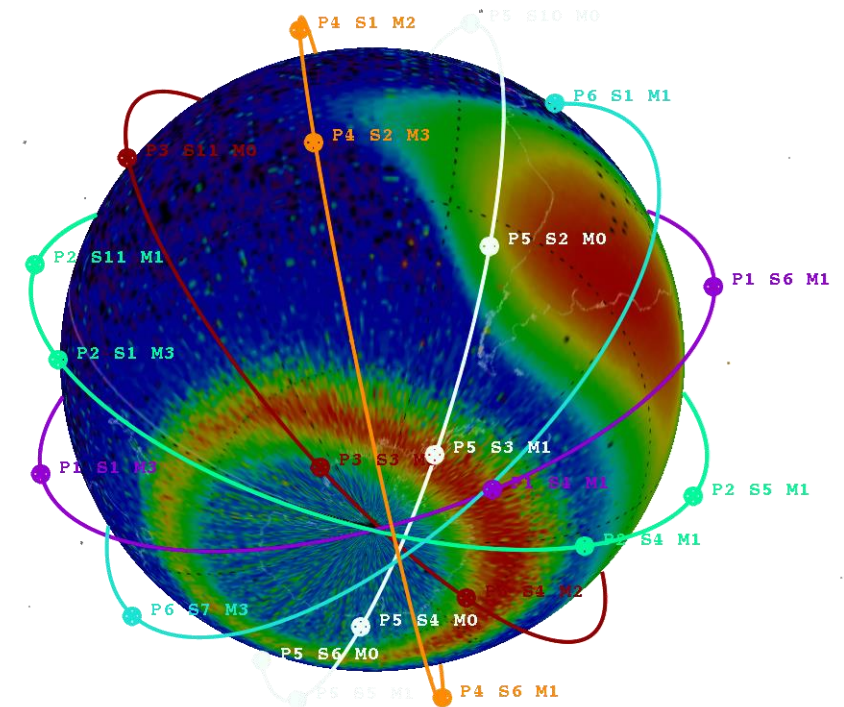
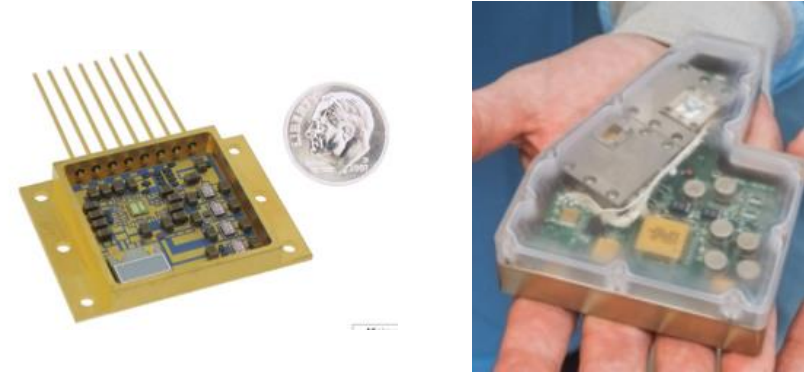


REACH

Overview

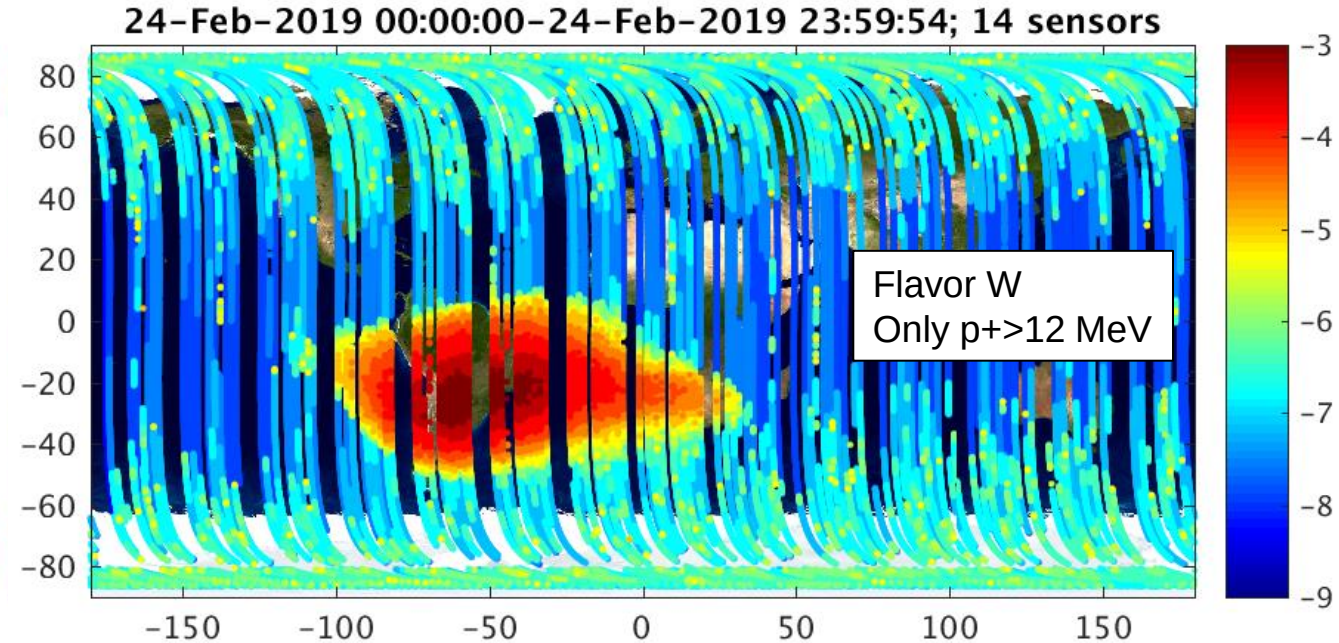
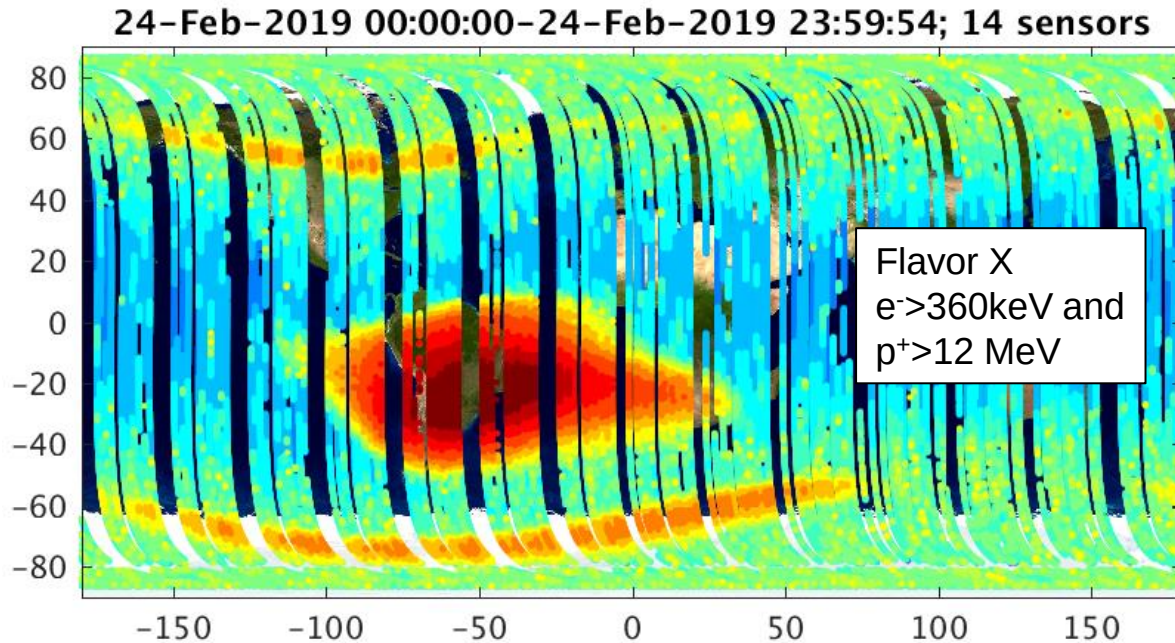
Responsive Environmental Assessment Commercially Hosted constellation

- REACH is a hosted space weather payload on the Iridium-Next constellation in LEO (780 km)
- 32 REACH payload in 6 orbital planes
 - Each payload has 2 micro-dosimeters – 64 total sensors
 - 6 different “flavors” each cover different energy range for electrons and protons
- Offers unprecedented temporal coverage of the LEO radiation environment
- First 2.5 years of REACH data (2017-2019) available on Zenodo: [10.5281/zenodo.5988170](https://zenodo.org/record/5988170)
- As of October 2022, receiving new data in real time
 - Sensors healthy, all continue to perform well
 - Data is publically releasable, pushed real time into unified data library (UDL)



REACH Constellation

REACH has 6 shielding thicknesses/electronic thresholds: we call them flavors



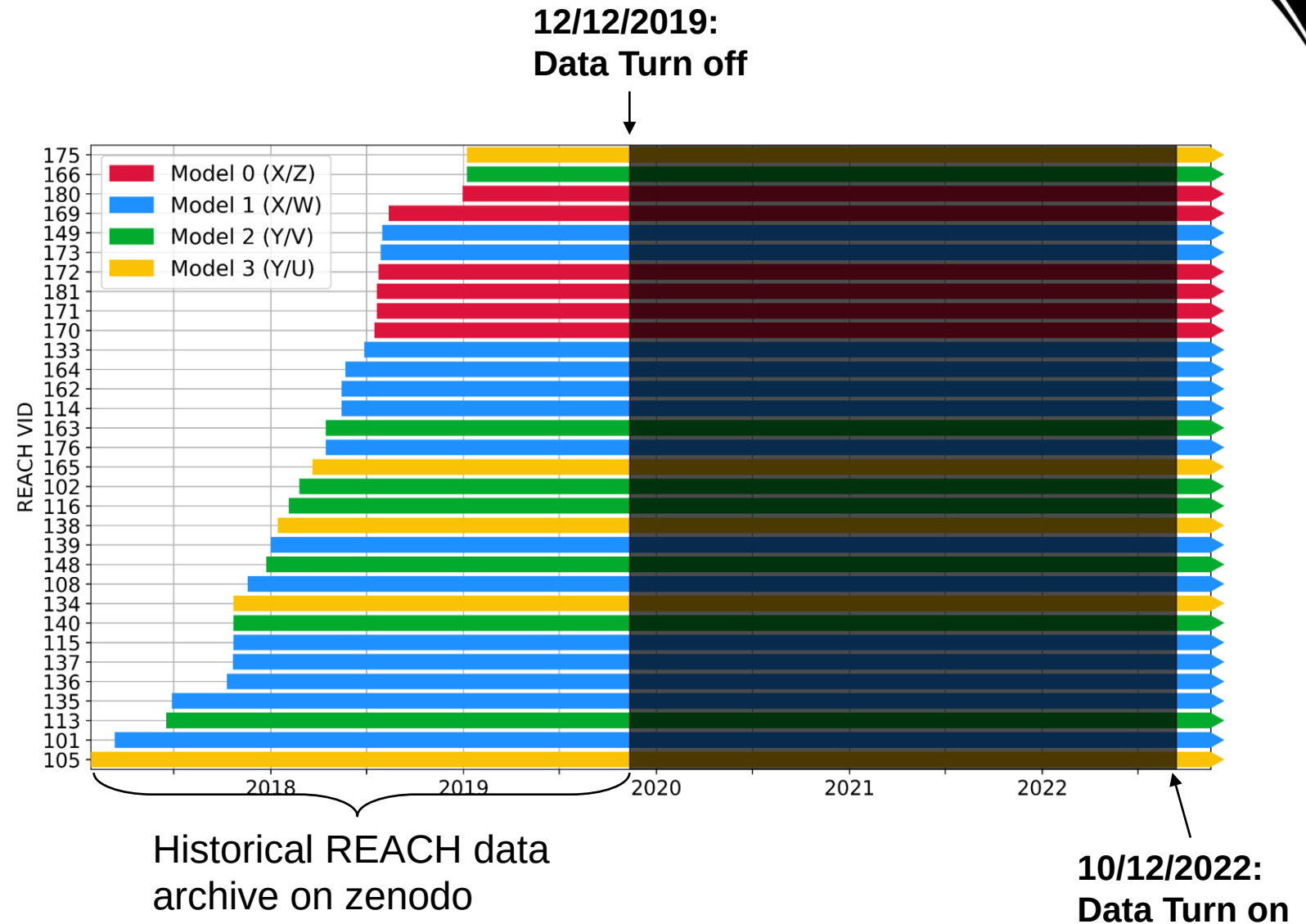
Dose rate in $\log_{10}(\text{rad/sec})$ from 14 individual sensors taken over 24 Feb 2019, southbound passes only

Flavor	Electrons	Protons	# Satellites
U	$> 4.97 \text{ MeV}$	$> 57 \text{ MeV}$	5
V	$> 3.41 \text{ MeV}$	$> 47 \text{ MeV}$	7
W	-	$> 12 \text{ MeV}$	14
X	$> 360 \text{ keV}$	$> 12 \text{ MeV}$	20
Y	$> 1.6 \text{ MeV}$	$> 31 \text{ MeV}$	12
Z	$> 50 \text{ keV}$	$> 200 \text{ keV}$	6

With 14 equivalent sensors in LEO measuring all the time, this is an unprecedented orbital coverage

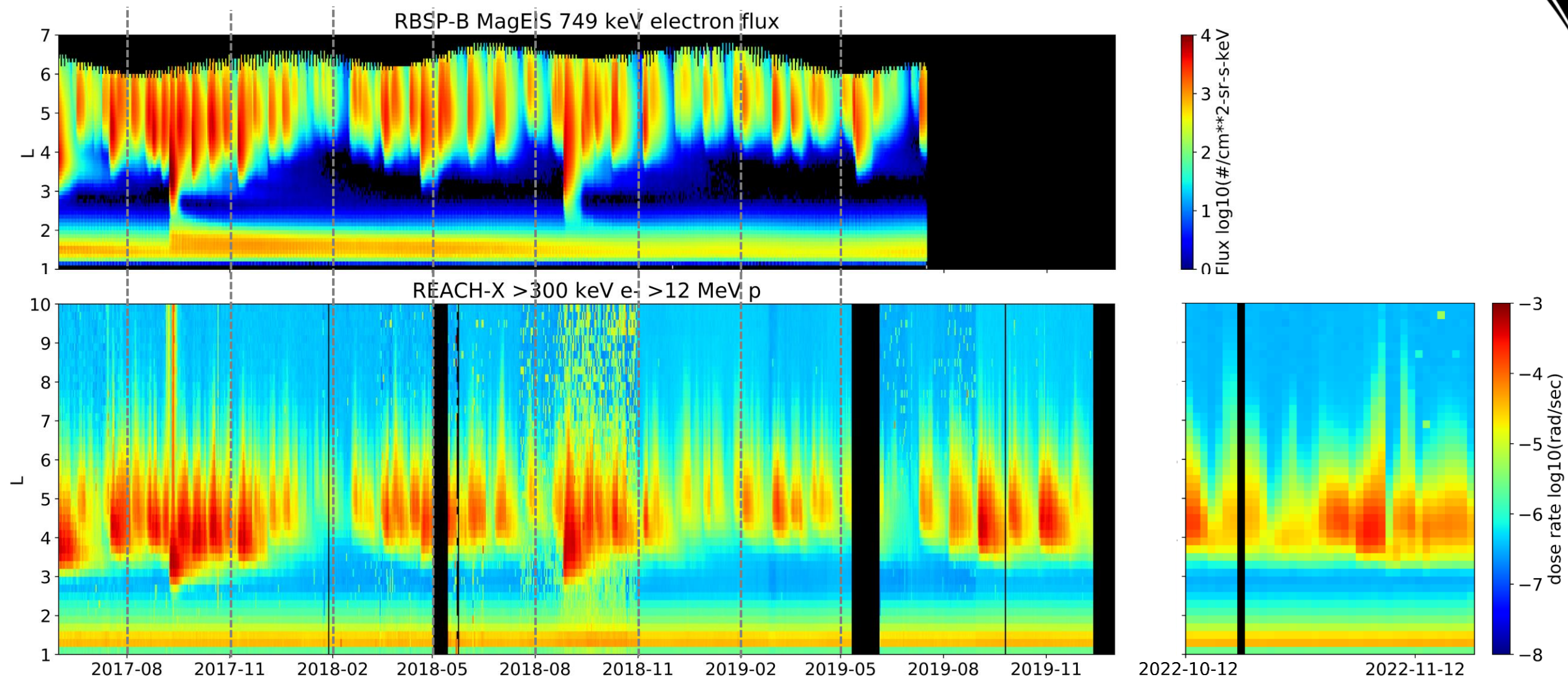
Current Status

- REACH data from 2017-Dec 2019
 - Data archived on zenodo:
[10.5281/zenodo.5988170](https://zenodo.org/record/5988170)
 - Documentation
 - csv files of 5-second average dose, fluxes
- Data turned back on 10/12/2022
 - Realtime data from all REACH payloads
 - Analysis shows all payloads healthy and performing well



Science Applications

Radiation Belt Monitoring



REACH Maps

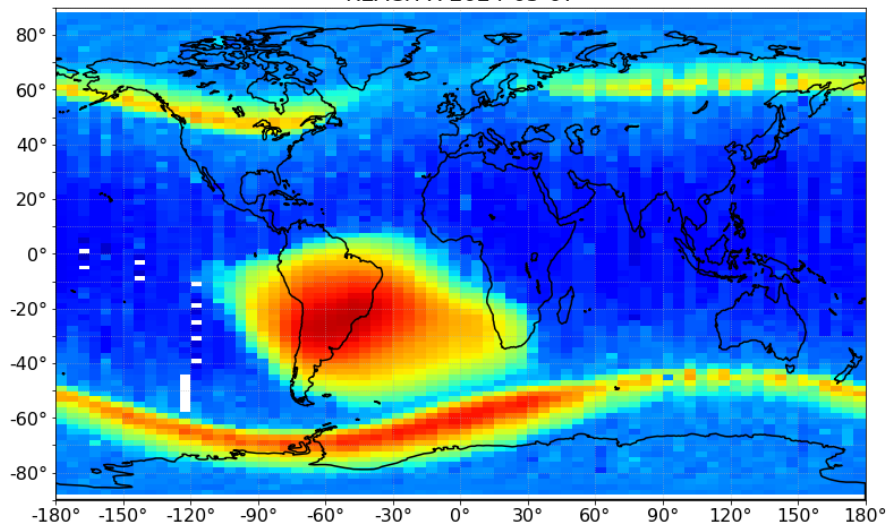
May 2024 Storm



Pre-Storm

REACH-X 2024-05-07

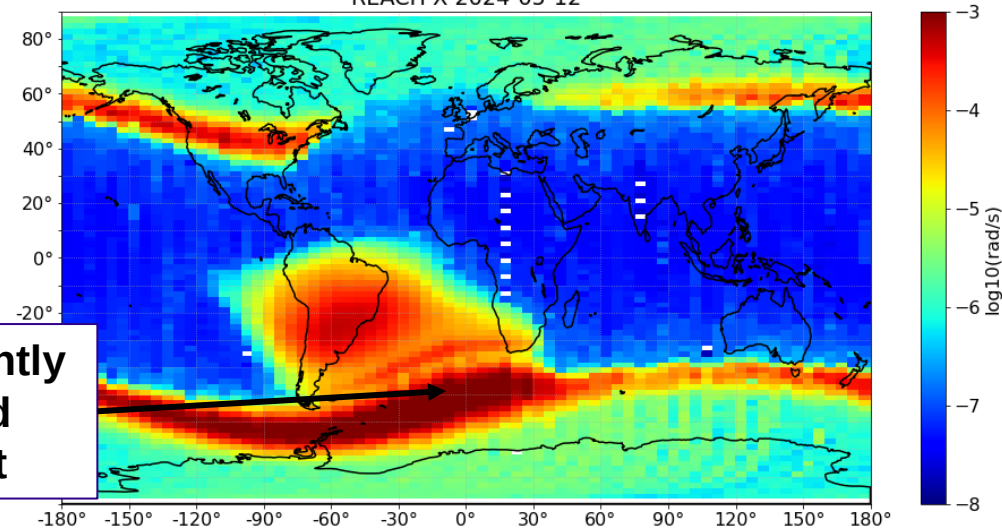
X: >300 keV
electrons,
>12 MeV
protons



Stormtime

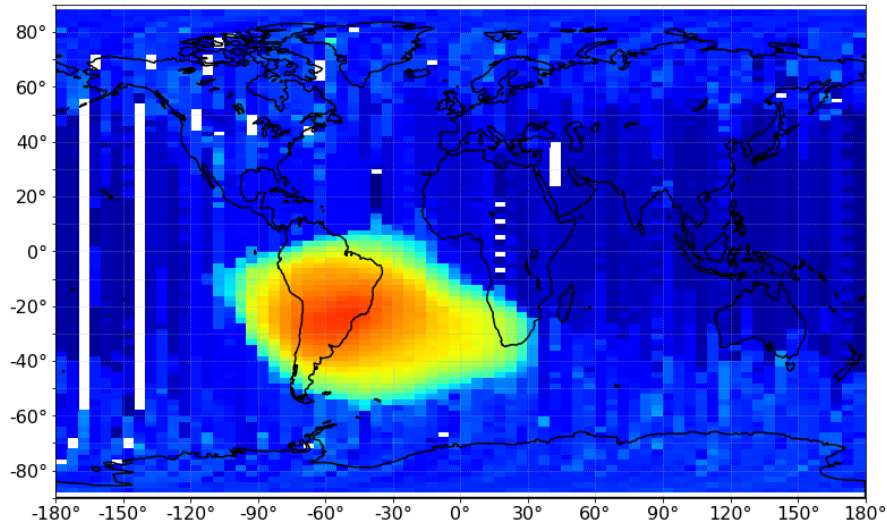
REACH-X 2024-05-12

Significantly
enhanced
outer belt

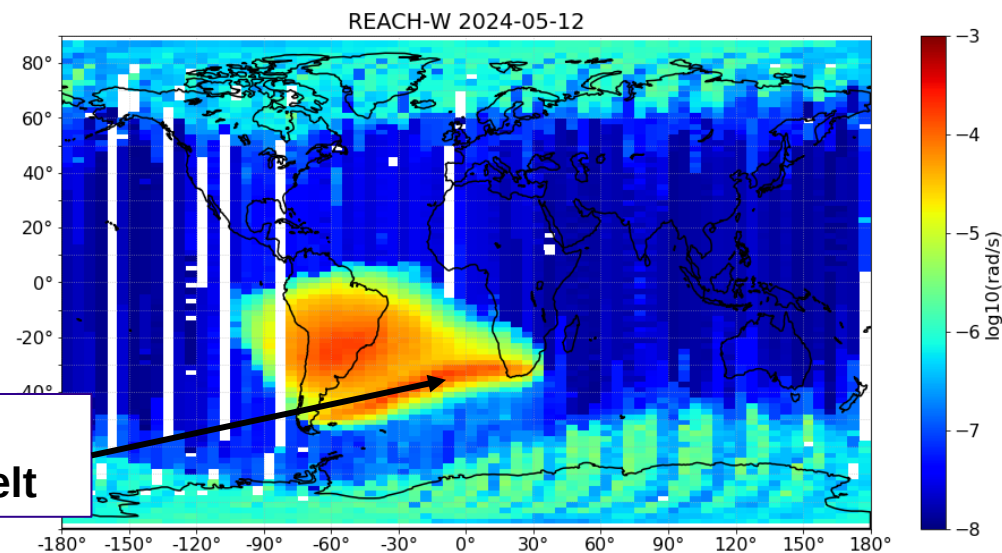


REACH-W 2024-05-07

W: >12 MeV
protons



Transient
Proton Belt



Summary



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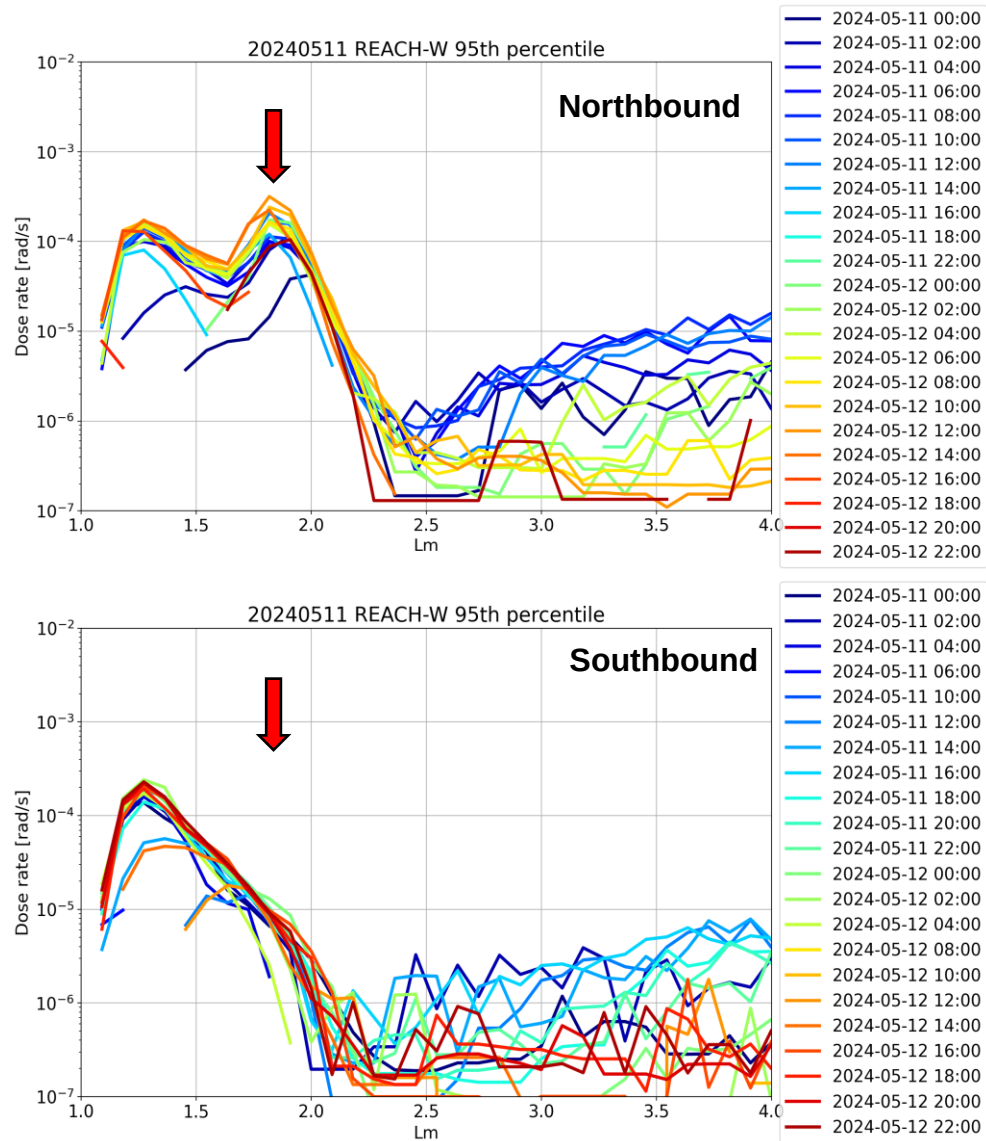
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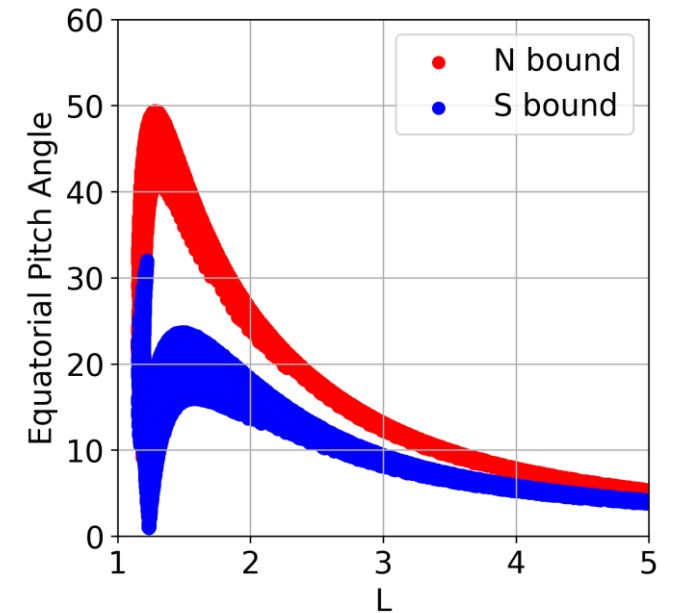
Backup

Proton Observations

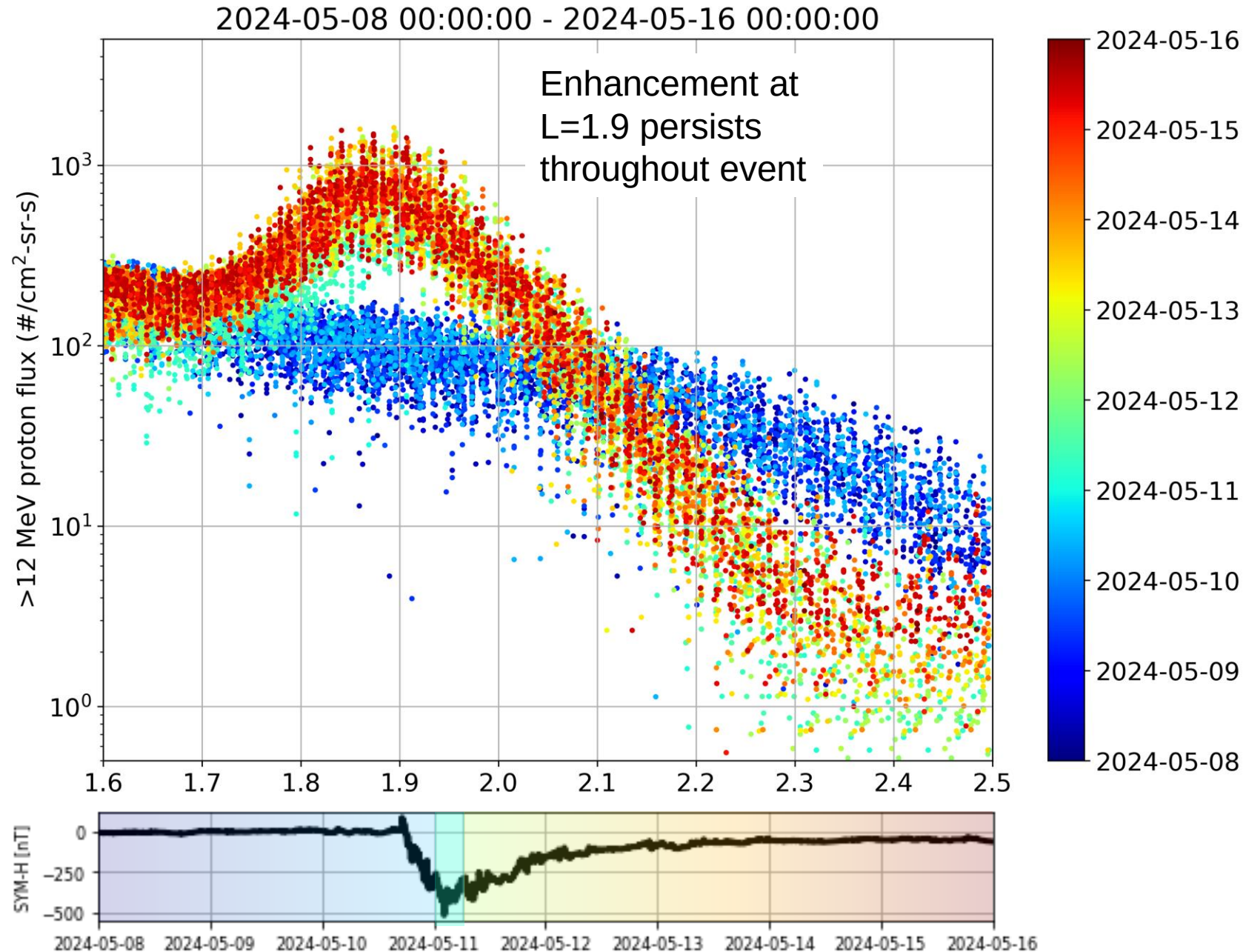
May 2024 Storm



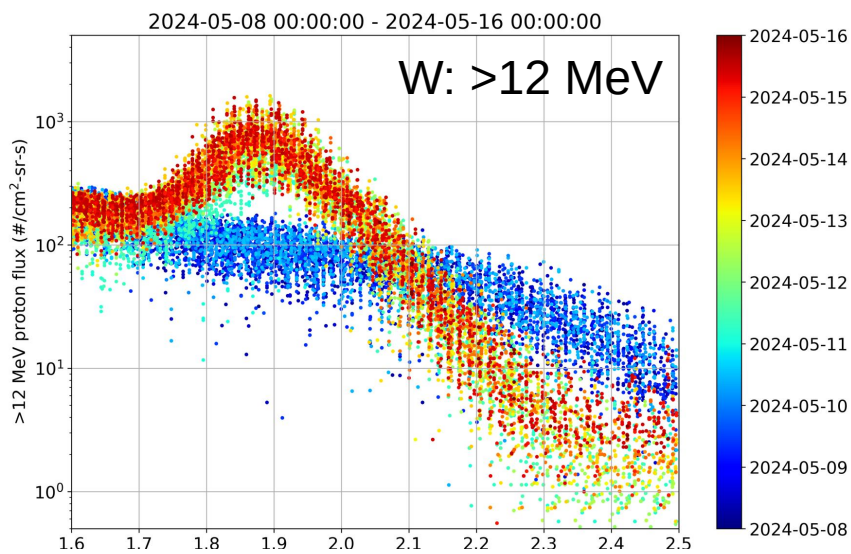
- A new transient Proton Belt, similar in magnitude to the SAA is only observed during northbound passes
- For REACH, northbound passes correspond to pitch angles closer to 90 degrees



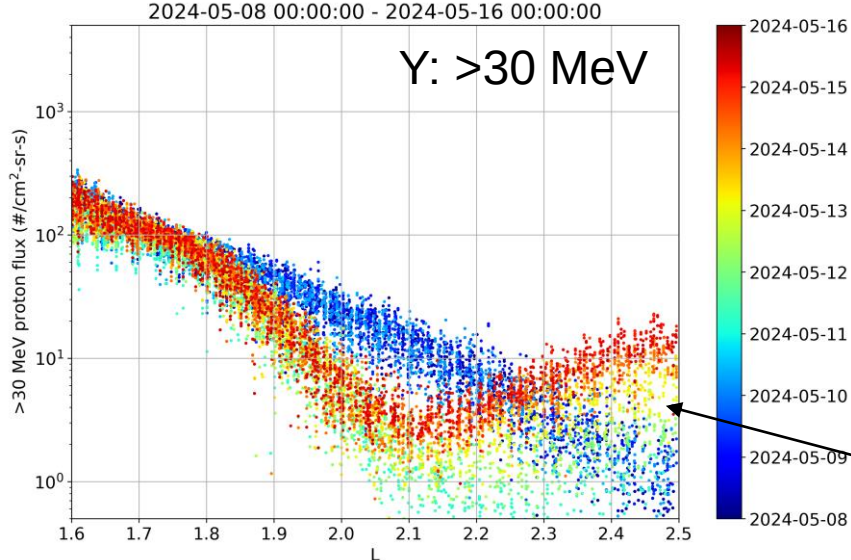
Proton Observations



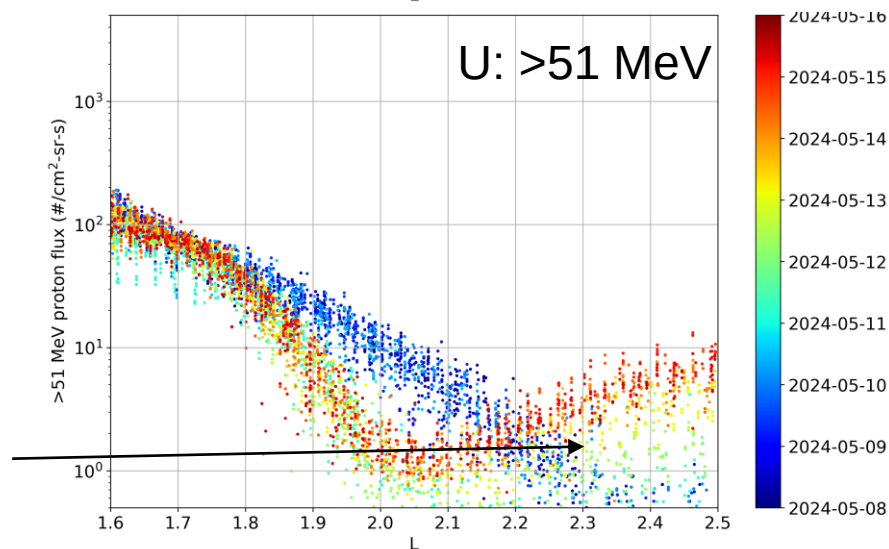
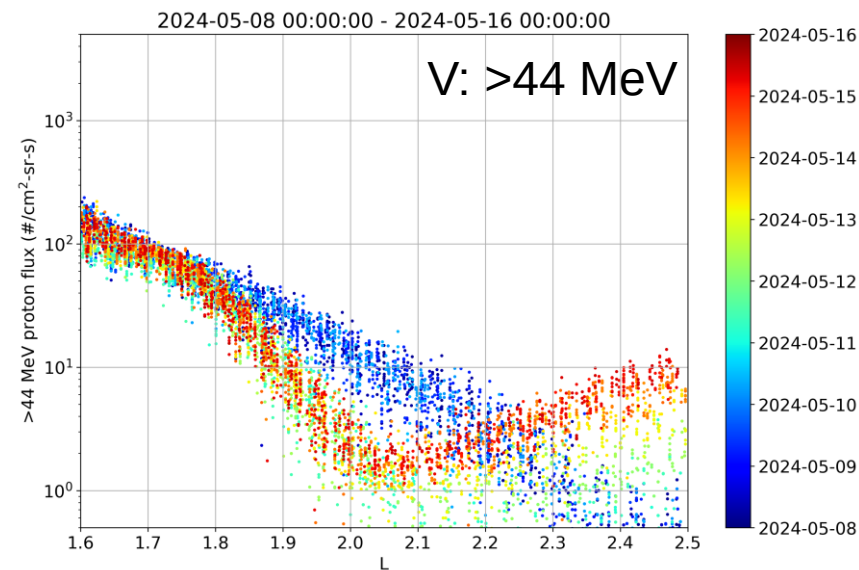
Energy Dependence



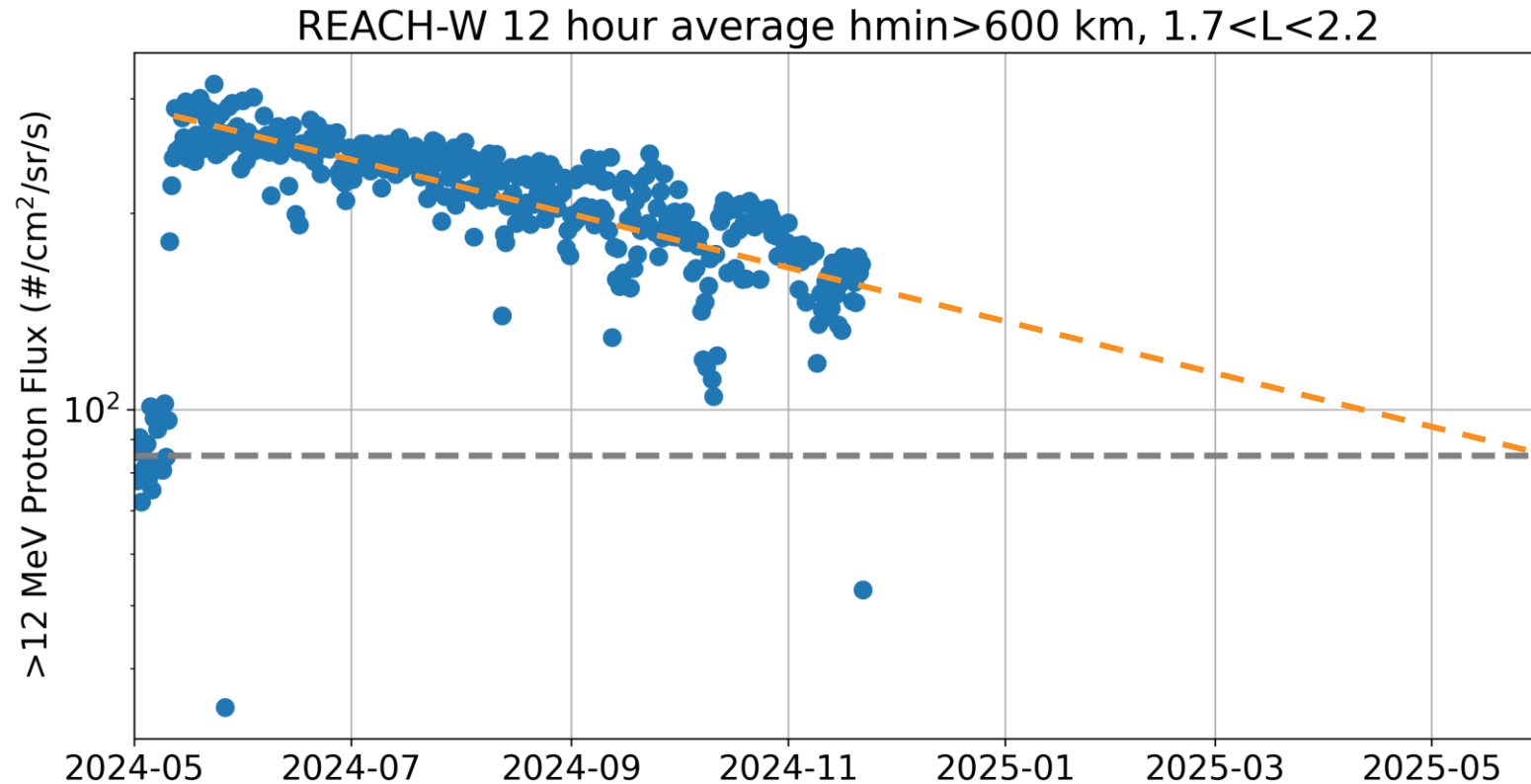
Transient
proton belt at
 $L=1.9$ is limited
to $\sim 12\text{-}30$ MeV



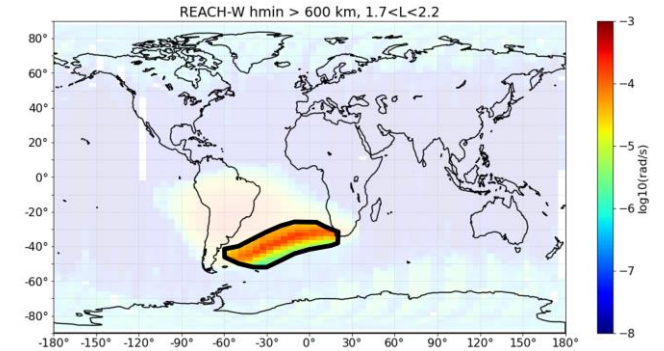
Observations of >2
MeV electrons



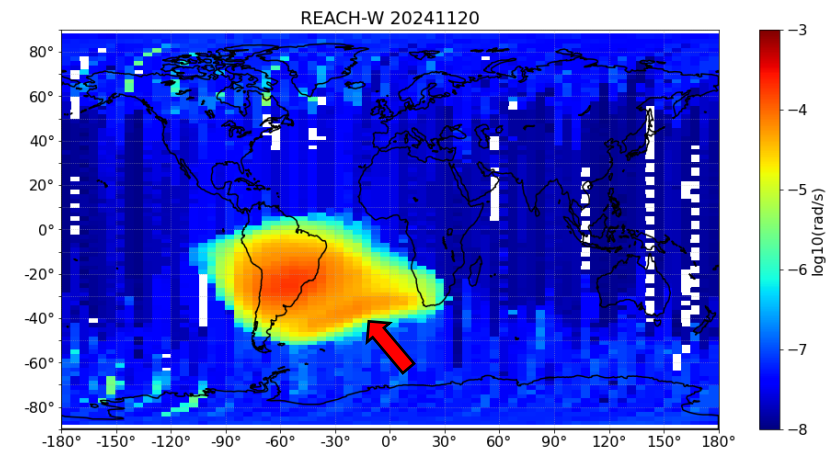
Long term observations



- Transient belt is still observable 6 months after storm
- Decay rate suggest belt will remain above pre-storm levels for ~ 1 year



Region with $h_{\min} > 600$ km
and $1.7 < L < 2.2$





Thank you