



Observations of NEOs with the TRAPPIST telescopes in support of size determination using polarimetry and radar

Ferrais Marin¹, Jehin Emmanuël², Petrescu Elisabeta²

¹ Florida Space Institute, University of Central Florida, USA;

²Space Sciences, Technologies & Astrophysics Research (STAR) Institute University of Liège, Belgium;

³Royal Observatory of Belgium, Brussels, Belgium.

contact: marin.ferrais@uliege.be; ejehin@uliege.be; elisabeta.petrescu@uliege.be

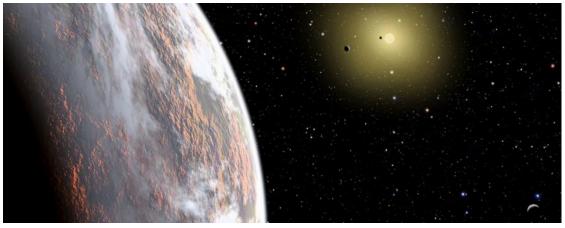
TRAnsiting Planets and Planetesimals Small Telescope

What do we have in common?

The love for astronomy

TRAPPIST

Robotic telescopes to study planetary systems



https://www.trappist.uliege.be/

Credit background: Basile Fauconnier



MEET THE TRAPPIST TEAM – COMETS AND ASTEROIDS



Emmanuël Principal Investigator



Sandrine IT Manager



Mathieu - Ph.D



Said - Ph.D



Elisabeta - Ph.D



Damien



Jean Co-Investigators

Aravind - Postdoc



Marin - Postdoc

Credit background: Basile Fauconnier

TRAPPIST – SOUTH Inaugurated: 8 June 2010 Site: Chile – La Silla Altitude: 2315 m Weather: 275 clear nights/year Funded by: ULiège & FNRS

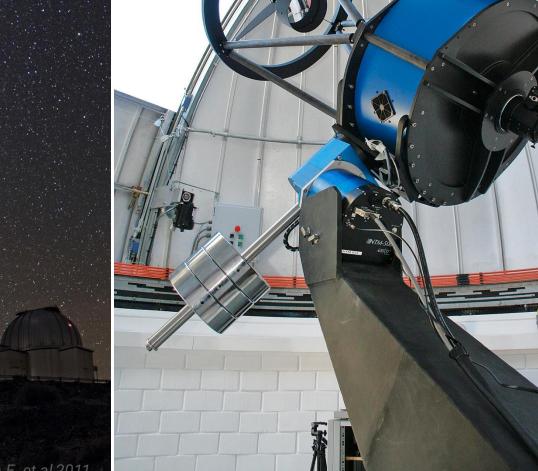
Credit image: Emmanu*ë*l Jehin

TRAPPIST – SOUTH

Telescope MPC: I40 Telescope design: Ritchey-Chrétien Diameter: 600 mm Focal length: 4800 mm

Mount: Astelco NTM-500 (German equatorial) Pointing altitude: 5 degrees Tracking accuracy: 1"/4 min

Camera: FLI ProLine PL3041-BB FOV: 22' x 22' Pixel size: 15 um Pixel scale: 0.64 arcsec/px Filter wheel 1: B, V, R $_c$, I $_c$, V+R, I+z, Solan z, NaI, H2O⁺ Filter wheel 2: OH, NH, CN, C₃, C₂, CO⁺, BC, GC, RC



Credit background: ESO Credit image right: Emmanu*ë*l Jehin TRAPPIST – NORTH Inaugurated: 6 October 2016 Site: Morocco – Oukaïmeden Altitude: 2751 m Weather: 220 clear nights/year Funded by: ULiège & FNRS

Credit image: Basile Fauconnier





TRAPPIST – NORTH

Telescope MPC: Z53 Telescope design: Ritchey-Chrétien Diameter: 600 mm Focal length: 4800 mm

Mount: Astelco NTM-500 (German equatorial) Pointing altitude: 5 degrees Tracking accuracy: 1"/4 min

Camera: Andor IKONL BEX2 DD FOV: 20' x 20' Pixel size: 13.5 um Pixel scale: 0.60 arcsec/px Filter wheel 1: B, V, R $_c$, I $_c$, V+R, I+z, Solan z, Nal Filter wheel 2: OH, NH, CN, C₃, C₂, H α , BC, GC, RC

Credit background: Basile Fauconnier Credit image right: Emmanu*ë*l Jehin

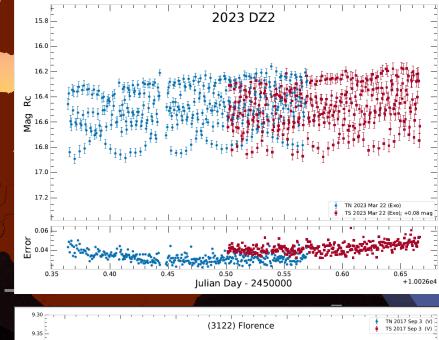
Reference: Jehin E. et al 2011

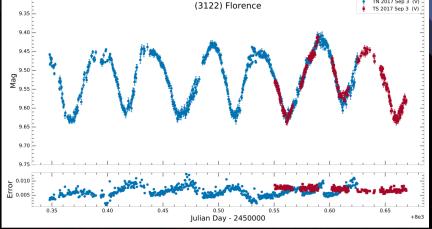


TRAPPIST – SOUTH Standard time: UTC - 4 hours Daylight saving: UTC - 3 hours

TO A TO A DOCTOR

Credit background: Basile Fauconnier





Meridian line

NEAs observed by TRAPPIST that have radar observations, and to support polarimetry measurements

The degree of linear polarization is inversely proportional to the albedo of the scattering surface (Umov law, 1905).

Polarimetry measurements (more info: Maxime Devogèle) CALIBRATION OF ALBEDO

Radar observations:
direct measurements of asteroid size

3D shape model of the asteroid

Surface roughness

RESULTS

Photometry observations: for a service of phase angles

albedo measurements

size

- H magnitude
- spin axis orientation
- spectral type
- rotation period
- photometric based shape model

Reference: Devogèle M. et al. – "Surface Heterogeneity, Physical and Shape Model of NEA (52768) 1998 OR2" (including Ferrais M., Jehin E.) PSJ February 2024

(52768) 1998 OR2 – PHA

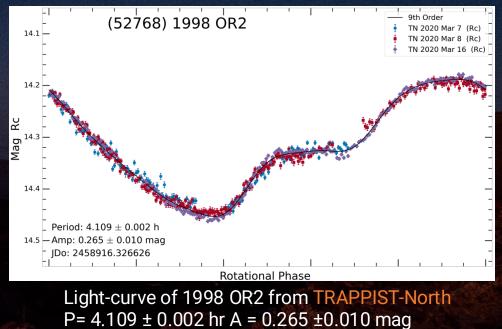
Close approach: 2020 April 29 at 16.4 LD; H = 16.04

Past characteristics:

 Rotational period values: P= 3.198 ± 0.006 hr; A = 0.29 ± 0.01 mag. Betzler & Novaes (2009);

 Classification: Xk type – using spectrophotometric observations / Xn type - using visible and IR spectroscopy Bus & Binzel 2002; Battle et al. (2022) New characteristics:

- H = 16.17
- Rotational period: P = 4.10872 ± 0.00001
- Classification:
 - M type using photometry and radar
- Photometric and radar shape model: 2.08 ± 0.10, 1.93 ± 0.10, 1.60 ± 0.03 km
- Diameter: 1.80 ± 0.10 km
- Polarimetric observations and thermal lightcurve: heterogeneous surface

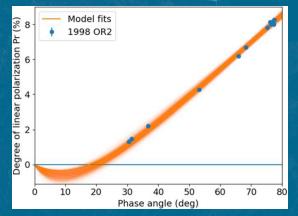


2020 fly-by using optical polarimetry, photometry and radar.

Torino Polarimeter (ToPol) on a 1.04m telescope 2020 February – April, solar phase angle 30° to 78°.

Photometric campaign: 22 telescopes, photometric filters. Observed during different apparitions between 2020 – 2022.

Radar from Arecibo Observatory 2020 April 13 – 23.



P= 4.109 ± 0.002 hr A = 0.265 ±0.010 mag					
Arecibo Observatory/NAS	SA/NSF	Arecibo Observato	ry/NASA/NSF	Arecibo Observat	ory/NASA/NSF
		e		1	
(52768) 1998 OR2 2	1020 Apr 18 UT	(52768) 1998 OR2	2020 Apr 18 UT	(52768) 1998 OR2	2020 Apr 18 UT

Ferrais M. et al. – "Optical polarimetry and radar as a tool for Planetary Defence", poster.

Devogèle M. et al. – "Surface Heterogeneity, Physical and Shape Model of NEA (52768) 1998 OR2" (including Ferrais M., Jehin E.) PSJ February 2024

(385186) 1994 AW1 – PHA

Close approach: 2015 July 15 at 23.3 LD; 2022 July 09 at 43.5LD; 2023 January 01 at 86.89LD

Characteristics:

- Binary asteroid = mutual events present
- Rotational period values: P1= 2.6 hr; diameter= 730m; P2= ~22 hr; diameter = ~300m.
- Classification: L type;

Fig. 1: Convex shape model from lightcurve inversion.

Fig 2: Shape modelling of the primary with the lightcurves + radar data.

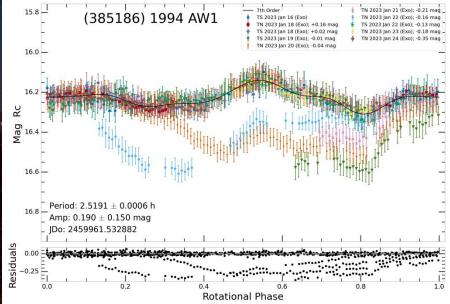
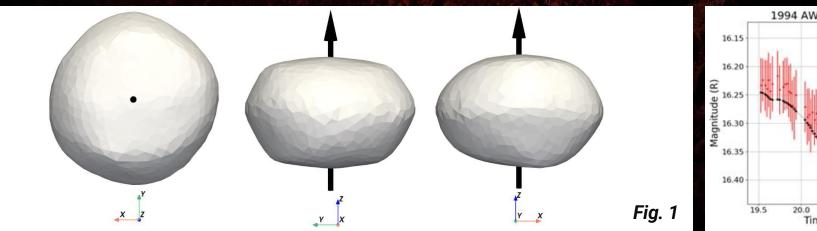




Fig. 2

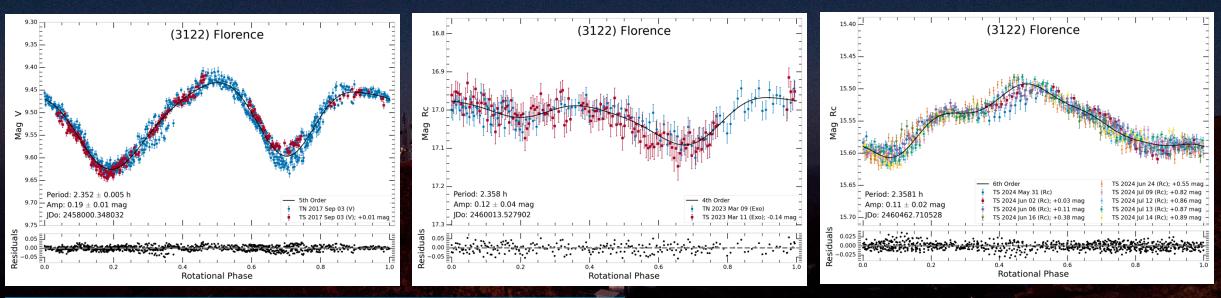


1994 AW1 - 2023-01-20 - TRAPPIST-North (Z53) 16.15 16.20 16.25 16.30 16.35 16.40 19.5 20.0 20.5 21.0 21.5 22.0 22.5 Time (UT hour when light left asteroid)

Ferrais M. et al. – "Shape model of the binary NEA 1994 AW1 from 2015 radar observations and long term lightcurve dataset", (including Devogèle M., Jehin E.), in progress.

(3122) Florence – PHA

Close approach: 2017 September 01 at 18.3 LD; 2024 October 01 at 148.5 LD



Florence, a triple system seen from TRAPPIST

Observed during three apparitions, different viewing geometries:

- 2017 September 3: A= 0.19 ± 0.01 mag
- 2023 March 09, 11: A= 0.12 ± 0.04 mag
- 2024 June 02 July 14: A= 0.11 ± 0.02 mag

Shape model in progress



Arecibo Observatory, 3122 Florence asteroid between 2017 April 2 - 5

(5189) 1990 UQ – PHA

Close approach: 2021 May 06 at 26.5 LD; 2022 July 09 at 43.5LD; 2023 January 01 at 86.89LD

Characteristics:

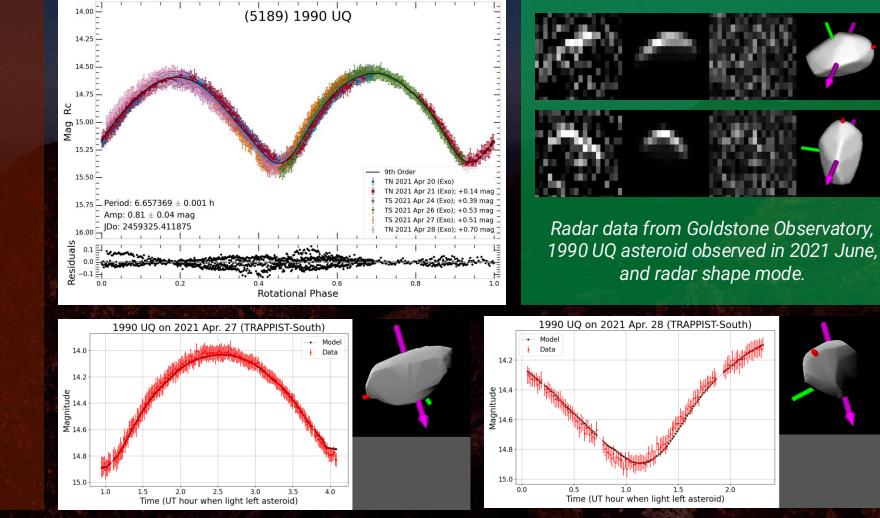
P= 6.657 ± 0.001hr A=0.81 ± 0.04

Observed in 2021 April 20-28, Exo filter.

Elongated object.

Shape model from the lightcurve inversion –Xavier Inosencio (UFC)

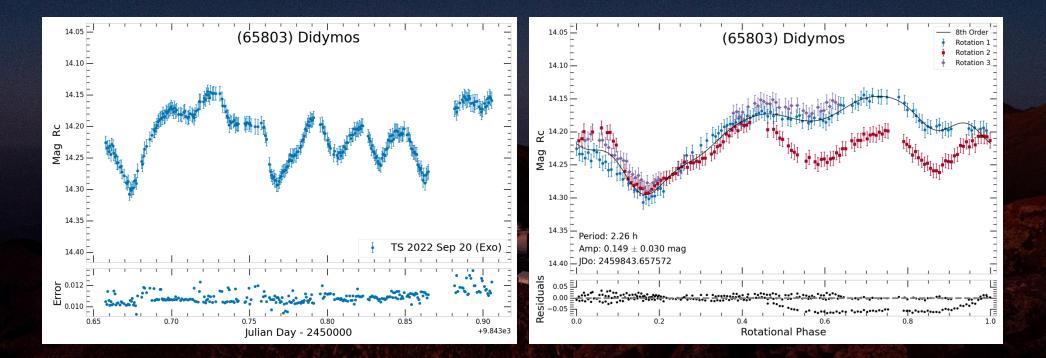
Size: 600 m Albedo: 0.4 work on going



Reference: Xavier I. et al. – "Characterizing Earth's Mini-Moon 2022 NX1 through its Analogous, 1990 UQ", (including Devogèle M., Ferrais M.), presentation.

65803 Didymos – PHA

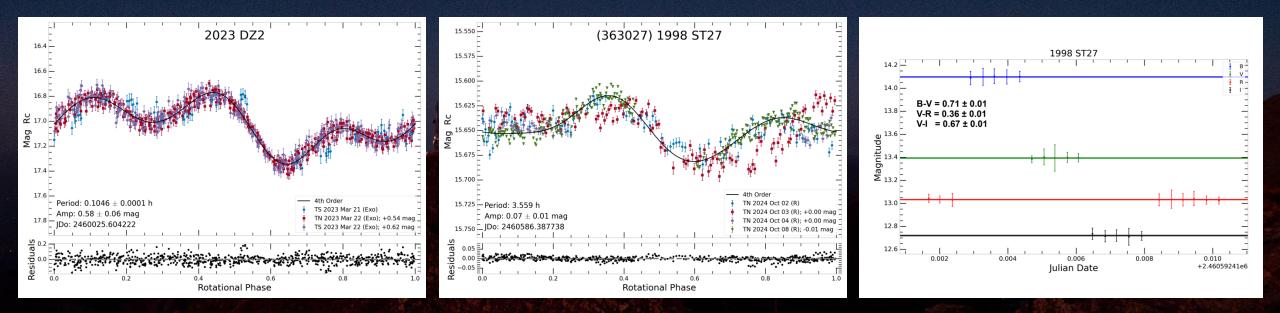
Close approach: 2022 October 04 at 27.57LD



65809 Didymos – binary system – mutual events P1=2.26h P2=11.92h Observed in R, I, B, V filters.

TRAPPIST-N: 2022 September – 2023 January TRAPPIS-S: 2022 September – 2022 November Phased curve on 2.26hr Secondary eclipse Dimorphos goes in the shadow of Didymos.

NEO lightcurves with TRAPPIST twins

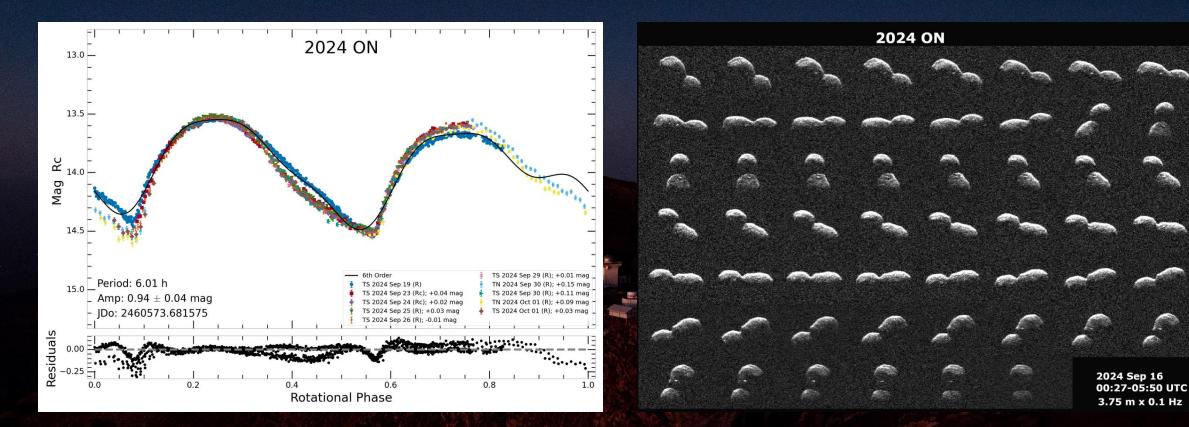


2023 DZ2 (PHA) – fast rotator: 6 min IAWN campaign Close approach: 2023 March 25 (0.45LD) Observed in R filter.

Reddy. et al – "2023 DZ2 PD Campaign" (including Ferrais M., Jehin E., Petrescu E.) PSJ June 2024 1998 ST27 (PHA) – triple asteroid: 3.5 hr Close approach: 2024 October 24 (9.26LD) Observed in R, I, B, V filters.

Petrescu. E et al – "Lightcurves of 2024ON, 2024 MK and 1998 ST27 from TRAPPIST" (including Ferrais M., Jehin E.) In preparation.

NEO lightcurves with TRAPPIST twins

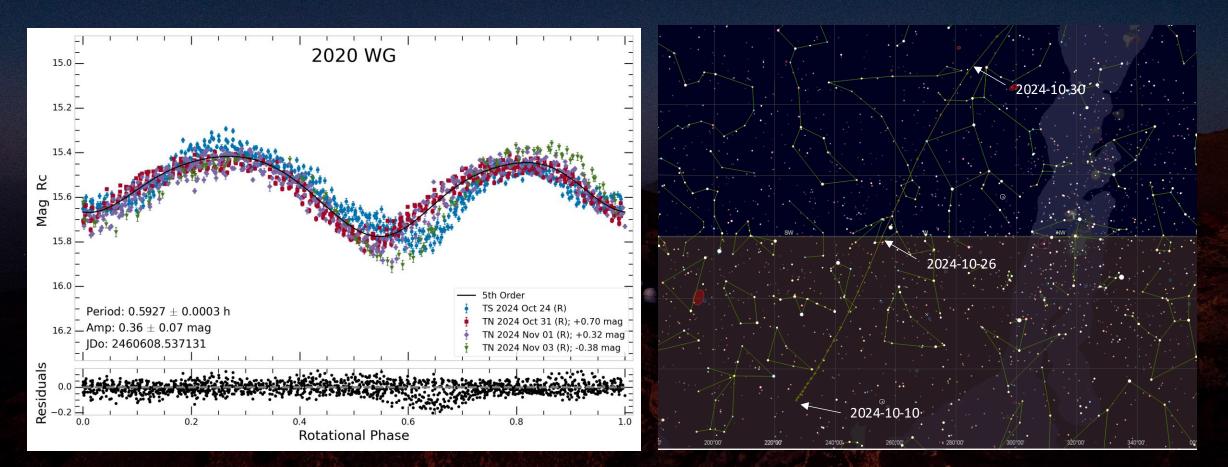


2024 ON (PHA) – contact binary P=6.01h Close approach: 2024 October 24 (2.59LD) Observed in R, I, B, V filters.

Petrescu. E et al – "Lightcurves of 2024ON, 2024 MK and 1998 ST27 from TRAPPIST" (including Ferrais M., Jehin E.) In preparation.

Radar data from Goldstone

NEO lightcurves with TRAPPIST twins



2020 WG (PHA), P=0.59 h Close approach: 2024 October 28 (8.65LD) Observed in R, I, B, V filters.

Trail of 2020 WG seen from TRAPPIST-N location. Chart made with Cartes du ciel.

TRAPPIST – "The paparaz

al nicknan

oorate — ejehin @uliege.be

