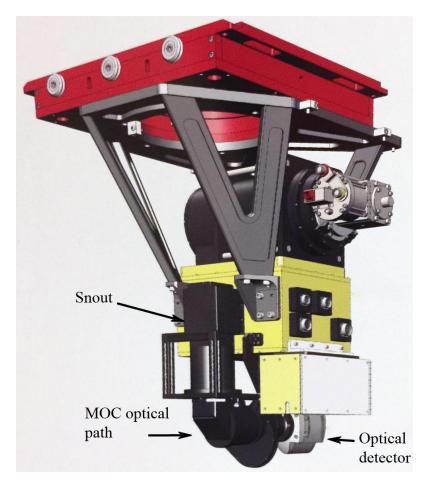
Characterizing Small Recently Discovered NEOs: Diameter and Albedo Measurements

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NEOs from the NASA-IRTF: MIRSI+MOC

- MIRSI is a thermal IR spectrograph and imager camera at the NASA-IRTF
- The new design include an optical camera, MOC
- Easily accessible and operable system for NEOs characterization and planetary defense support
- Optical+Thermal = accurate albedo and diameter

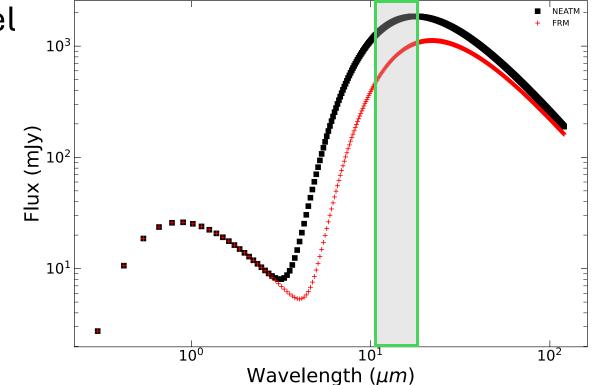


Hora et al. 2024

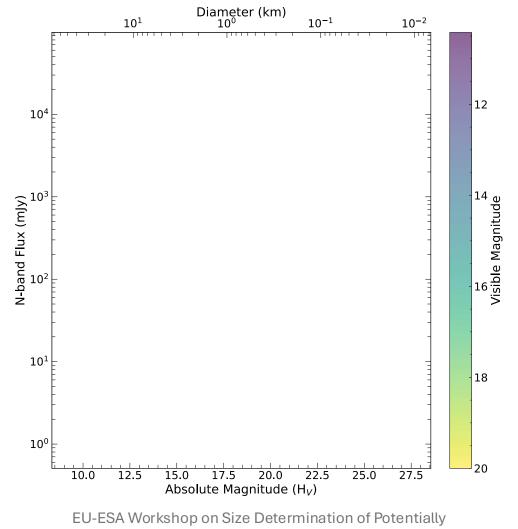
Spectral Energy Distribution of NEOs

- Standard Thermal Model
- Near-Earth Asteroid Thermal Model
- Fast Rotator Thermal Model
 - Isothermal Latitude Model

$$D = \frac{1329}{\sqrt{p_V}} 10^{-\frac{H_V}{5}} \text{MOC}$$



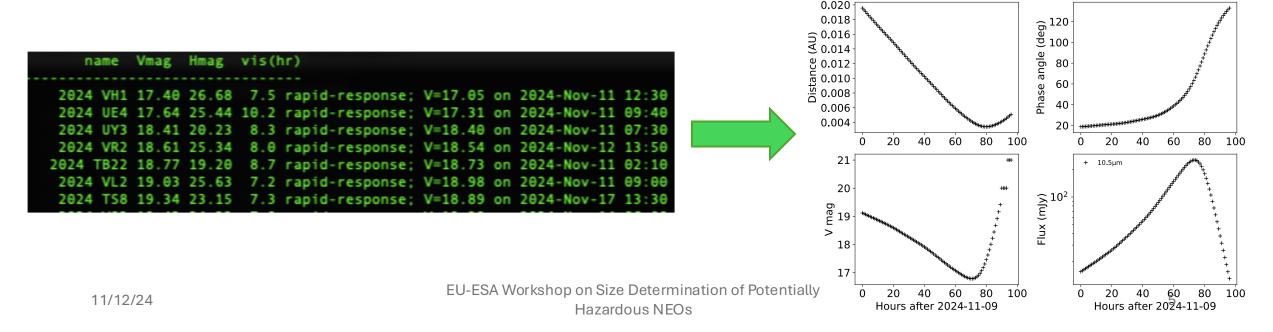
Predicted SED of Observable NEOs



Hazardous NEOs

MIRSI as a Support for Planetary Defense

- The MIRSI-NEO program, typically consist of characterizing ~30 known NEOs/year
 - 3-4 hours observing blocks
- If available, we prioritize recently discovered NEOs

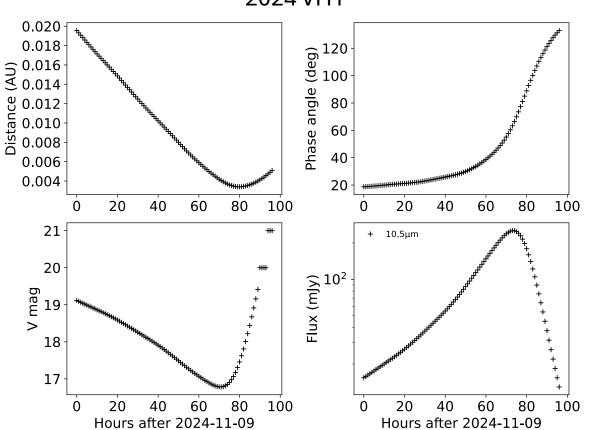


MIRSI as a Support for Planetary Defense

- The MIRSI-NEO program, typically consist of characterizing ~30 known NEOs/year
 - 3-4 hours observing blocks
- If available, we prioritize recently discovered NEOs
 - Must fall within our requested block
 - Must be bright enough for MIRSI (>120 mJy)
- So far, we have characterized 6 recently discovered NEOs
 - Simultaneous R-band photometry $\rightarrow H_V \rightarrow p_V$
 - Color measurements are also possible to determine taxonomy with MOC

Recently Discovered Asteroids

- Represents the smaller members of the NEOs population (D<100m)
- Short observing window
- Fade in brightness quickly after close approach
- Requires a fast response observing technique
 - Queue formatted telescopes (ideal)
 - Director's Discretionary Time (ToO-IRTF)



2024 VH1

MIRSI as a Support for Planetary Defense

- Diameter and albedo provide information on two critical aspects of planetary defense
 - Size
 - Composition
- International Asteroid Warning Network (IAWN; Reddy et al., 2024)
 - A global planetary defense practice
 - During the 2023 DZ2 campaign, MIRSI contributed by measuring the albedo and diameter range in a 35 minutes block
 - $19 m > D_{eff} < 26 m$
 - *p_V*~49%
 - ~6 minutes rotation period (MOC)

Advantages of MIRSI for a Fast Characterization

- MIRSI is always at operational conditions (T~5K)
- MOC provides simultaneous optical characterization
 - Minimization of H uncertainties at the time of the observation
 - Account for irregularly shaped objects with high amplitude
- Data acquisition, reduction, and analysis in (almost) real-time

Disadvantages of MIRSI for a Fast Characterization

- The target must be bright enough to be detected by MIRSI with its current sensitivity (>~120 mJy)
- MOC is a guiding camera with FOV ~1'x1'
 - Uncertainty in the NEO ephemeris could difficult it's identification
 - Inaccurate position of the NEO in the MIRSI FOV would make challenging the blind stacking process
 - Include Opihi to recover the NEO (i.e., 2024 JT3 found 110+ arcseconds away from nominal)
- The observability window of recently discovered NEO is quite narrow upon close approach...

Lessons Learned & Takeaways

- Optimizing Interrupt Time Allocation at NASA-IRTF
 - High-demand facilities should incorporate better coordination and "last-minute" communication between scheduled observers
 - Clearly stating in the observing proposal whether the program could be relocated if high-priority circumstances require so
 - Observers are under no obligation to provide access
- MIRSI is a "ready-to-go" optical and mid-IR camera to measure albedos and diameters of newly discovered asteroids accurately
 - We have tested the MIRSI+MOC system under the scenario of a possible impact to improve our planetary defense strategies
- Github repository for reduction pipeline: <u>https://github.com/jhora99/MIRSI</u>