



Differential Algebra Space Toolbox (DAST)

Workshop on Nonlinear Propagation of Uncertainties using Differential Algebra September 22nd, 2015 Escape Dance Room, ESTEC

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 - VPT Architecture
- Applications: the Apophis case
- Advanced Applications
- Conclusions





Introduction





Introduction

What if all these aspects are combined together in a tool for uncertainty analyses in space dynamics?



Differential Algebra Space Toolbox (DAST)



Inside DAST two sub-modules are introduced:

- Software Framework (SF)
- Uncertainty Propagation Tool (UPT)











09/22/2015







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Purpose of SF Module

- The SF includes the implementation of more advanced features as:
 - DA operations between vectors and/or matrices of DA
 - DA implementation of propagation schemes (RK78)
- The SF also provides astrodynamics routines, which allow:
 - Defining a state vector at a certain epoch, with respect to given coordinate system, reference frame and units
 - Writing dynamical models and evaluating them using DA propagators.
 - Solving typical astrodynamics problems (DA Kepler solver)









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SF Core Routines



- Algebraic Vector
 - Any arithmetic type (double, DA, ...)
 - Component-wise operations
 - Intrinsic functions evaluations
 - Vector norm, cross & dot products
 - Algebraic Matrix
 - Any arithmetic type (double, DA, ...)
 - Common matrix operations
 - Dynamical Model
 - Abstract class, right-end-side (RHS) of ODEs
 - Evaluation of RHS of ODEs
- Propagator (variable and fixed step)
 - Abstract class
 - Propagate Dynamical Models
- RK78



SF Astrodynamics Routines

- State Vector
 - State (Algebraic Vector)
 - o Epoch
 - o Coordinate, Frame, Units
 - Unit and state transformations routines
- Guidance Model
 - Abstract class
 - Specialized guidance models
 - Lift-Off, Pitch Push-Over, Bilinear Tangent
 - Custom (user-supplied)
- AstroDynamical Model
 - Abstract class, derived from Dynamical Model
 - o Dimension, parameters, frame, units
 - Specialized dynamical models
 - R2BP, CR3BP, nBP, Reentry, Rel. Dynamics, Euler, Ascent dynamics, Rendezvous
 - Custom (user-supplied)









```
// Define state vector
StateVector<DA> X0( 6, t0, CARTESIAN, ECI, STDUNITS ); // create a state vector of DAs
X0.state(0) = -6.86740024261e+03 + sigma[0]*DA(1);
X0.state(1) = -5.31170788774e+02 + sigma[1]*DA(2);
X0.state(2) = 8.76076142159e-02 + sigma[2]*DA(3);
X0.state(3) = 5.79118186169e-01 + sigma[3]*DA(4);
X0.state(4) = -7.57879982355e+00 + sigma[4]*DA(5);
X0.state(5) = 3.27557629871e-01 + sigma[5]*DA(6);
// Convert to Keplerian coordinates
StateVector<DA> X0_kepl = ConvertState( X0, KEPLERIAN, ECI, STDUNITS );
// Define propagation interval
double sma = cons( X0_kepl.state(0) ); // retrieve semi-major axis
double T = 2*kPI*sqrt( sma*sma*sma/GM ); // compute the orbital period
double tf = t0 + 5.0*T;
                                       // define final propagation time
// Propagation
double tol = 1e-9;
                                        // specify the required tolerance
StateVector<DA> XF = AstroPropagator( dynR2BP, X0, tf, tol );
// Output
```

return 0;

}









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Purpose of UPT Module

- The UPT is aimed at performing uncertainty propagations and statistical analyses.
- The UPT comes with a set of useful Matlab routines, which allow:
 - An easy interaction with the SF even for standard users (no need to write or compile code)
 - Managing simulation results
 - Easy graphical representations of the performed analyses













UPT Interface Functions:

- "UPTmodel.m" for dynamical model definition (model, reference frame, units, etc.)
- "UPTmethod.m" for propagation method definition and all the required settings (order, number of samples, etc.)







- UPT MEX Core Routines:
 - "UPT.cpp" to set up the DA environment and perform DA propagation
 - "UPTaux.cpp" contains auxiliary functions for conversions between C++ objects and Matlab variables







- UPT Analysis Functions:
 - "UPTrun.m" to collect all the inputs and call mex files.
 - "UPTeval.m" for additional evaluations of the final DA map (different distribution, number of samples).
 - "evalCDA.m" for fast evaluations of a compiled DA map (for advanced users).







UPT I/O Functions:

- "load_CDA.cpp" for loading a compiled DA from file (for advanced users)
- "compiledDA.cpp" for saving compiled DA in an easy to handle format (for advanced users)
- Other additional routines for plotting results ("drawCube.m", "plot_gaussian_ellipsoid.m", etc).



UPT Summary

- Summarizing, the main steps required for a typical analysis are:
- 1) Call to **UPTmodel** for dynamical model definition.
- 2) Call to **UPTmethod** for propagation method definition.
- 3) Call to **UPTrun** to perform DA propagations.
- 4) Eventually, call to **UPTeval** for additional evaluations of the final DA map.



Let's see how it works with a simple example...







STEP 1 Model declaration:

Performed using the UPTmodel function.

model = UPTmodel('Model', 'R2BP', ... 'MainAttractor', 'EARTH',... 'InitialState', state, ... 'Coordinate', 'RECTANGULAR', ... 'Frame', 'J2000',... 'FrameCenter', 'EARTH', ... 'InitialEpoch', t0, ... 'FinalEpoch', t0, ... 'FinalEpoch', tf,... 'LengthUnits', 'KM', ... 'TimeUnits', 'SEC',... 'AngleUnits', 'RAD');

Additional fields may be required for different dynamical models











UPT Example: Two-Body Dynamics

Other interesting features:

- Linearized Dynamics (LD)
 - Propagation of initial covariance through STM
 - The STM is directly the linear part of a DA final state
 - No need to integrate variational equations or compute finite difference





Other interesting features:







Applications: the Apophis case

Scenario

On its next approach on April 13, 2029, the asteroid Apophis will come within 35000 km of Earth.



DA can be used to describe how the distance of closest approach (DCA) changes due to uncertainties on the initial state.



Apophis Initial Conditions

- Apophis' ephemerides of June 18, 2009
 - Equinoctial elements
 - Diagonal covariance matrix

	Nominal value	σ	
а	0.922438242375914	2.29775 × 10–7	AU
P1	-0.093144699837425	3.26033 × 10–7	
P2	0.166982492089134	7.05132 × 10–7	
Q1	-0.012032857685451	5.39528 × 10–7	
Q2	-0.026474053361345	1.83533 × 10–7	
Ι	88.3150906433494	6.39035 × 10–4	deg





Apophis Closest Approach: DAMC

- 1. Initial condition initialized as DA and converted to cartesian ecliptic J2000 coordinates
- 2. Propagation forward in time to TCA, 13 April 2029 21:46 UTC using a n-body model
- 3. Taylor expansion of final state $[x_f]$ used to compute Taylor expansion of DCA

$$[\mathbf{x}_f] = \mathcal{T}_{\mathbf{x}_f}(\delta \mathbf{x}_0) \quad \longrightarrow \quad [DCA] = \mathcal{T}_{DCA}(\delta \mathbf{x}_0)$$

	DCA (AU)		Obtained DCA is
NEODYS-2	0.00025	-	NEODYS-2
DAST	0.00025198		





- 4. Propagation for further 10 days and computation of the Taylor expansion of the final distance wrt Earth $\mathcal{T}_{d_f}(\delta x_0)$
- 5. Generation of N_T samples from initial statistic
 - MC Pointwise propagation, d_f^i
 - DAMC Polynomial evaluation of $[d_f]$ to compute approximation \tilde{d}_f^i





Apophis Closest Approach: DAMC

Comparison standard MC and DA-based MC for increasing orders:





- Polynomial Bounder (PB) propagation method:
 - Use polynomial bounder on $\mathcal{T}_{d_f}(\delta x_0)$ to compute $[d_f^{LB}, d_f^{UB}]$

N.B. No need to generate samples and evaluate maps!!

Let's compare the results with those obtained with DAMC $[\tilde{d}_f^{min}, \tilde{d}_f^{max}]$:

Nsamples	ε^{LB} [AU]	ε^{UB} [AU]
1e3	-2.8108e-04	-5.6943e-04
1e6	2.6205e-05	-2.7087e-05

Minimum and maximum distance \tilde{d}_{f}^{i} among all samples





- Linearized Dynamics (LD) propagation method:
 - a) Apophis' initial states initialized as **first order** DA variables
 - b) Propagation till TCA to compute $[x_f] = \mathcal{T}_{x_f}^1(\delta x_0)$
 - c) Extraction of transition matrix Φ from $\mathcal{T}_{x_f}^1(\delta x_0)$
 - d) Compute final covariance matrix from initial covariance C_0

$$C_f^{UPT} = \Phi C_0 \Phi^T$$

NB1: No need of variational equations!!

NB2: No need to generate samples and evaluate maps!!



Apophis Closest Approach: LD

Comparison between DA-based MC and LD:





Advanced Applications

Scenario

Suppose we want to assess the effects of uncertainties in the atmosphere model and/or in thrust magnitude and direction during the ascent trajectory of a Rocket Launch Vehicle (RLV).

Polynomials obtained with a DA approach could be used to:

- Assess uncertainties at orbit insertion
- Correct thrust to improve robustness





Main Assumptions

However...





Main Assumptions

- Within each phase:
 - Thrust magnitude does not depend on altitude (or atmospheric pressure) and is assumed constant (only direction changes);
 - Mass flow rate is constant;
 - Lift/Drag coefficients are constant (independent of Mach number);
 - Aerodynamic reference area is constant;

Possible uncertain model parameters

- Initial state (except for the first phase)
- C_L , C_d , ρ_0 , T_{max}





Main Assumptions

Summary:

Phase	Uncertain states	Model uncertain parameters	Guidance uncertain parameters	Total number of variables
А	-	$C_L, C_D, \rho_0, T_{max}$	$\Delta\Theta, \Delta\Psi, i_T$	7
В	$R, \lambda, \delta, V, \gamma, \chi$	$C_L, C_D, \rho_0, T_{max}$	-	10
С	$R, \lambda, \delta, V, \gamma, \chi$	$C_L, C_D, \rho_0, T_{max}$	-	10
D	$R, \lambda, \delta, V, \gamma, \chi$	$C_L, C_D, \rho_0, T_{max}$	Θ_0	11





Management of Multiple Phases

The expansion of the flow from t_0 to t_f yields to a Taylor polynomial which depends on a quite large number of variables (~35/40)











Pros:

- Optimization of the number of DA variables (comp. costs reduction);
- Possibility to Taylor expand each phase.

Cons:

 Need to properly manage the interactions between one phase and another.



VEGA Launcher: Target Orbit

Consider a simplified launch trajectory of VEGA:

Departure site: Kourou launch site

Target orbit: Polar Earth Orbit (PEO)

$$h_f = 700 \text{ km}$$

 $e_f = 0$
 $i_f = 90^\circ$





VEGA Launcher: Flight Sequence

Consider a simplified launch trajectory of VEGA:

Mission flight sequence:

Phase	Duration [s]	Active propulsion	Guidance Law
А	0-90	P80	Lift-Off & Pitch Push-Over
В	90-180	Z23	Gravity Turn
С	180-270	Z9	Gravity Turn
D	270-920	AVUM	Bi-Linear Tangent







VEGA Launcher: Nominal Trajectory













```
STEP 0 Problem variables declaration:
                   ~~~~
       Generation of samples
                           *****
  nsamples = 1e5;
  & PHASE A
  xOA distr = repmat(state, nsamples, 1); % dummy distribution (fixed initial state)
  samplesA = mvnrnd([p0A, r0A], ...
                    diag([sigma pA.^2, sigma rA.^2]), ...
                    nsamples);
  pOA_distr = samplesA(:,1:4); % distribution of model uncertain parameters
  rOA_distr = samplesA(:, 5:7);
                                  % distribution of control/quidance uncertain parameters
  % PHASE B
 p0B_distr = mvnrnd(p0B, diag(sigma_pB.^2), nsamples); % distribution of model uncertain
                                                     parameters
  % PHASE C
  . . .
  % PHASE D
  . . .
```



UPT Example: RLV Ascent Dynamics

STEP 1 Model declaration:

Performed using the UPTmodel function:









UPT Example: RLV Ascent Dynamics

STEP 3 Run UPT:

Performed using the UPTrun function.

% Run simulation
[UPToutput_PHA, UPTinput_PHA] = UPTrun('Model', model_PHA, 'Method', method_PHA);

STEP 4 Analysis of additional samples:

Use the *UPTeval* function to evaluate the polynomial with the distribution generated at the beginning:

```
% Evaluation with generated samples
[ xfA_distr, ~, ~ ] = UPTeval( UPToutput_PHA, x0A_distr', p0A_distr', r0A_distr');
mean_fA = mean(xfA_distr,2); % compute mean
cov_fA = cov(xfA_distr'); % compute covariance
```



UPT Example: RLV Ascent Dynamics



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DAMC Performance

Comparison standard MC and DA-based MC (PHASE A)



With 3rd order polynomials, the break-even point is reached with 70 samples!



DAMC Performance

Comparison standard MC and DA-based MC (PHASE B)



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DAMC Performance

DAMC computational efficiency vs number of variables





Additional Analyses

 Polynomial Bounder (PB) and Linearized Dynamics (LD) methods can be used as well





- Two products have been implemented:
 - DACE: Efficient computation of Taylor expansions of arbitrary order
 - DAST: DA-based uncertainty propagations in astrodynamics
- For what concerns DAST:
 - Propagations can be run on several ready-to-use dynamical models and any DA-compatible custom dynamical model
 - Three different uncertainty propagation techniques:
 - DA-based Monte Carlo
 - Ranging estimation using a polynomial bounder
 - Linear covariance propagation



- Main advantages:
 - More efficient than standard Monte Carlo for typical number of samples
 - No need of deriving and using variational equations
 - Obtained polynomials can be used to propagate different statistics
- Main consideration:
 - Uncertainty propagation is based on a Taylor approximation of the mapping function



- Size of uncertainty set and order shall guarantee sufficient accuracy
- Computational time increases for higher order and large number of variables.





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